| 3 | Geothermal Gradient, Curie Point Depth and Heat Flow Determination of |
|---|---|
| 4 | Some Parts of Lower Benue Trough and Anambra Basin, Nigeria, Using    |
| 5 | High Resolution Aeromagnetic Data                                     |

**Original Research Article** 

6

1 2

7 Abstract

**Background and Objective:** This study, which is bounded within Latitude  $6^0 00' - 6^0 30'$ N 8 and Longitude  $7^{0}00'$   $7^{0}$  30'E with an approximate area of about 3025 km<sup>2</sup> within parts of 9 lower Benue trough and Anambra basin of Nigeria, aims at outlining the regional temperature 10 11 distribution and delineating areas that are geo thermally responsive by determining: the heat change per unit distance, the heat flowing from the earth's interior to the outer surface and 12 the deepest depth at which the minerals loss their magnetic properties within the study area 13 without any heat data. Materials and Methods: For the aim to be achieved, the data was 14 subjected to quantitative analysis with the aid of the WingLink, ArcGIS, Origin Pro 8, Ms 15 Excel and sulfer 10 software's. Regional-residual was applied on the total magnetic intensity 16 map and thereafter the residual divided into sixteen (16) overlapping windows. Spectral depth 17 analysis was performed upon the overlapping windows and this revealed depth due to low 18 frequency and high frequency components. The depths due to the low frequency components 19 exemplify the Curie depth point (CPD). Results: The average sedimentary thickness or the 20 average depth due to the low frequency part was ascertained to be -5 km while the 21 geothermal and heat flow varies from -25.2 °Ckm<sup>-2</sup> to -38.9 °Ckm<sup>-2</sup> and from -64.4 mWm<sup>-2</sup> to 22 -97.3 mWm<sup>-2</sup> but with average values of -32.1 °Ckm<sup>-2</sup> and -80.1 mWm<sup>-2</sup> respectively. 23 24 **Conclusion:** These results suggest alternative geothermal energy resource to be plausible within windows 2, 4, 8, 10, 12, 15 and 16 and presence of some amount of sedimentary 25 26 thickness within the area of study.

27 Keywords: WingLink, ArcGIS, CPD, Windows, Heat flow, Geothermal gradient, HRAM,
28 Raster, CSV.

29

#### 31 INTRODUCTION

Aeromagnetic technique is a type of geophysical technique in which a magnetometer is 32 towed behind an aircraft. Aeromagnetic data has been, and will continue to be, handy in the 33 geophysical and geological investigation of the earth's interior. Umeanoh (2015) asserted that 34 35 aeromagnetic data can be used in mapping magnetic basement in sedimentary rocks and delineating igneous bodies within sedimentary sections as well as locating lineaments and 36 37 structures which could be possible host to varying earth resources like groundwater, minerals and hydrocarbon. Aeromagnetic method has been used majorly for the estimation of depth to 38 basement and thickness of sediments within sedimentary basins. Although this method was 39 used in mapping igneous and metamorphic rocks and structures related to them because these 40 rocks have high magnetization compared to other rocks<sup>2, 3</sup>. 41

42 Despite being used in delineating architectural framework of the earth's subsurface geology, the aeromagnetic method, can be applied successfully in defining geothermal gradient of an 43 44 area via spectral analysi.s. It was pointed out that the assessment of variations of the Curie isotherm of an area can provide valuable information about the regional temperature 45 distribution at depth and the concentration of subsurface geothermal energy <sup>4</sup>. The Curie-46 point temperature varies from region to region, depending on the geology of the region and 47 mineralogical content of the rocks. Therefore, one normally expects shallow Curie-point 48 depths in regions that have geothermal potential, young volcanism and thinned crust <sup>4, 5</sup>. In 49 order to determine the Curie-point depths, i.e. the bottom of the magnetized rocks, and to 50 map these depths, a frequently used method is the analysis of magnetic data. 51

52

The Benue trough is based on lead-zinc mineralization, limestone deposits, coal deposits, 53 coal deposits, pyroclastics, brine spring and brine Lake<sup>1</sup>. Interestingly, the Benue trough 54 especially the lower Benue trough has been studied deeply by researchers, students and 55 private organizations using either a combination of magnetic, gravity or any other 56 57 geophysical technique for mineral exploration. Currently aeromagnetic studies has received the kind of attention other geophysical techniques have received. Many authors <sup>6-9</sup> have 58 59 carried out studies on magnetic anomalies to estimate bottom depths of the related bodies for various purposes through the application of various techniques. However, some authors <sup>4, 5</sup> 60 have also undertaken determination of Curie-point depths (CPD) within some major basins of 61 62 the world.

63 In the present study, the objective is to estimate the average sedimentary thickness, geothermal gradient, Curie Point Depth (CPD) as well as the heat flow within the study area. 64 This aids in viewing the thermal structure of the crust. The CPD nevertheless can be defined 65 as the deepest level in the earth crust containing materials which creates discernible 66 signatures in a magnetic anomaly map <sup>11</sup>. In other words, the further the depth, the material 67 changes from a ferromagnetic material to a paramagnetic one. However, one of the important 68 69 parameters that determine the relative depth of the Curie isotherm with respect to sea level is the local thermal gradient, that is, heat flow <sup>12</sup>. This Curie Isotherm, generally, has a 70 temperature of 550°C  $\pm$  30°C <sup>11</sup>. This point is assumed to be the depth for the geothermal 71 source (magmatic chamber) where most geothermal reservoirs tapped their heat from in a 72 geothermal environment. Measurements have shown that a region with significant 73 geothermal energy is characterised by an anomalous high temperature gradient and heat flow 74 <sup>4</sup>. It is therefore expected that geothermally active areas would be associated with shallow 75 Curie point depth <sup>13</sup>. It is also a known fact that the temperature inside the earth directly 76 controls most of the geodynamic processes that are visible on the surface <sup>14</sup>. In this regard, 77 78 Heat flow measurements in several parts of African continent have revealed that the mechanical structure of the African lithosphere is variable <sup>15</sup>. 79

The very concept underlying magnetic prospecting is the existence of a magnetic dipole or
monopoles within the rocks constituting the earth<sup>1</sup>.

Magnetic force expression, F between two magnetic monopoles of strength P<sub>1</sub> and P<sub>2</sub> is given
by:

1

84

85

$$F = \frac{p_1 p_2}{\mu r^2}$$

86 Where

87  $P_1$  and  $P_2$  are dipoles

r is in meters and it is the distance between  $P_1$  and  $P_2$ 

89  $\mu$  is the free space permeability

90

91 The above Coulomb's equation is the basic underlying principle of magnetic prospecting.

92 Magnetic monopole,  $P_1$  or  $P_2$  exert force per unit pole strength and it can be expressed as:

93 
$$H = \frac{P}{r^2}$$

| 94  |   |  |  |  |  |  |  |
|-----|---|--|--|--|--|--|--|
| 95  | Where   |  |  |  |  |  |  |
| 96  |   |  |  |  |  |  |  |
| 97  | P is the magnetic monopole  |  |  |  |  |  |  |
| 98  | r is the distance between the force in question and the magnetic monopole                       |  |  |  |  |  |  |
| 99  | H = strength of the magnetic field  |  |  |  |  |  |  |
| 100 |   |  |  |  |  |  |  |
| 101 | Generally, the existence of a monopole has never been accounted for. Basically, magnetic        |  |  |  |  |  |  |
| 102 | monopoles or dipole are made up of positive and negative poles separated by a distance. The     |  |  |  |  |  |  |
| 103 | force produced and thus existing between monopoles can be estimated by vectorially adding       |  |  |  |  |  |  |
| 104 | the forces generated by each of the monopoles or dipole. The force generated by a simple bar    |  |  |  |  |  |  |
| 105 | magnetic can be compared to the force generated by a dipole.                                    |  |  |  |  |  |  |
| 106 |   |  |  |  |  |  |  |
| 107 | Magnetic materials positioned within a magnetic field will acquire magnetic force and will      |  |  |  |  |  |  |
| 108 | experience magnetic induction. Due to the inducing field, one can measure the strength of the   |  |  |  |  |  |  |
| 109 | magnetic field known as the intensity of magnetization, Ji, induced on the material and this is |  |  |  |  |  |  |
| 110 | expressed as:   |  |  |  |  |  |  |
| 111 |   |  |  |  |  |  |  |
| 112 | $J_i = kT 		3$  |  |  |  |  |  |  |
| 113 | Where   |  |  |  |  |  |  |
| 114 | $J_i$ is the magnetization  |  |  |  |  |  |  |
| 115 | k = susceptibility of the magnetic material   |  |  |  |  |  |  |
| 116 | T = inducing field.   |  |  |  |  |  |  |

#### LOCATION AND GEOLOGIC SETTINGS OF THE STUDY AREA 118

The study area is located in Enugu state and parts of Anambra state, south-east Nigeria. The 119 coordinates are Latitude 6<sup>0</sup> 00'-6<sup>0</sup> 30'N and Longitude 7<sup>0</sup>00' 7<sup>0</sup> 30'E. The study area falls 120 within the Lower Benue Trough and parts of Anambra basin. The Benue Trough generally 121 has been subdivided into three: the Upper Benue Trough at the NE Nigeria, the Middle 122 Benue Trough and the Lower Benue Trough. The Lower Benue Trough has somewhat 123

developed different tectonic history resulting in the formation of the Anambra Basin to the
west and Abakaliki Anticlinorium to the east<sup>1</sup>.

The Anambra Basin remained a stable platform supplying sediments to the Abakaliki 126 depression during a period of spasmodic phase of platform subsidence in the Turonian. 127 Following the flextural inversion of the Abakaliki area during the Santonian uplift and 128 folding, then the Anambra Basin was initiated. Four Cretaceous depositional cycles where 129 recognized in the Lower Benue and each of these was associated with the transgression and 130 regression of the sea<sup>17</sup>. The opening of the Atlantic Ocean in the Middle Albian to Upper 131 Albian gave rise to the transgression of the first sedimentary cycle. The Asu River group 132 which consist predominantly sandstone and shale was deposited at this time. Between the 133 Upper Cenomanian and Middle Turonian, the second sedimentary deposition of the Ezeaku 134 Shale occurred. The third sedimentary circle occurred from Upper Turonian to the Lower 135 136 Santonian leading to deposition of the Awgu Shale and Agbani Sandstone. The fourth and final depositional phase took place during the Campanian-Maastrichtian transgression. It was 137 138 at this time that the Nkporo Shale, Owelli Sandstones, Afikpo Sandstone, Enugu Shale as well as the coal measures including the Mamu Formation, Ajali Sandstone and Nsukka 139 Formation was deposited<sup>17</sup>. Fig.1 shows the study area and the regional geology of the Lower 140 Benue trough. The geological map (Fig 2) of the study was extracted from the regional 141 geologic map and redigitized using the Arc GIS software for enhanced interpretation of the 142 aeromagnetic map. Visually inspecting the map shows five main formations within the study 143 area. These include: Nkporo Shale Formation, Mamu Formation, Ajali Formation, Nsukka 144 Formation and Ameki Formation. The ages of the formations range from Maastrichtian to 145 Campanian and to Eocene (Ameki formation). 146

147

#### 148 MATERIALS AND METHODS

This research work made use of a digitized High Resolution Aeromagnetic (HRAM) data compiled by Fugro Airborn Service on behalf of the Nigerian Geological Survey Agency (NGSA) in 2009. The Composite Total Magnetic Intensity (CTMI) map with sheet number 301 was obtained in comma separated variable (CSV) format and in half degree sheet. The aeromagnetic data in CSV format was later transformed into a raster format (Fig 3). The high resolution survey was carried out at flight line spacing of 500 meters and at a ground 155 clearance of about 100 meters while the tie line spacing was 2 km at flight line direction of156 NE-SW.

157

The acquisition of the HRAM data initially took place in Ogun state in the year 2003.
Thereafter, between year 2004 and 2009 the rest of the country was divided into project areas
referred to as Phases I and II covering 44% and 56% respectively of the total area. The raw
data was pre-processed using Oasis montaj software by the NGSA and was transmitted as
IGRF corrected total magnetic intensity (TMI) data and was also saved in CSV file format.







## Geological Map of SHEET 301 (UDI)

170 Fig. 2: Geologic Map of the Study Area (Sheet 301 and Re-digitized using the Arc GIS)



178

179

### 180 Fig 3: AEROMAGNETIC RASTER MAP OF THE STUDY AREA (nT)

181 For effective data analysis, processing and interpretation, the WingLink, ArcGIS softwares were used for qualitative interpretation while the MS excel, Surfer 10 and Origin Pro 8 182 Geophysical software's were used for quantitative interpretation. The data which was 183 obtained in CSV format was opened and digitized in ESRI ArcGIS software for onward 184 processing and interpretation. The data was processed and converted in a format usable by 185 the WingLink visualization software with the aid of the ArcGIS. The digitization was done in 186 grid of 1 km x 1 km spacing and values of TMI, X (latitude) and Y (longitude) were picked at 187 the intersection of the grid nodes. 188

189

Manual digitization is the most elementary and least efficient method of digitisation, its accuracy when carefully done, compares favourably with other sophisticated methods using 192 computer programs<sup>14</sup>. The 1 km x 1 km grid points generated over 6000 sample points. The 193 x and y show the coordinates while the z represents the TMI value at the point. This was 194 implemented in ArcGIS 9.3 software and the xyz data was saved as MS Excel file format. 195 The data saved as excel file format was thereafter imported into the Micro soft (MS) excel 196 environment for band pass filtering. The band filtering was carried out so as to create sixteen 197 (16) overlapping spectral windows upon which Fast Fourier Transform (FFT) was performed.

198

#### **DEPTH DETERMINATION**

The depth to sedimentary thickness and its morphology can be determined using spectral 200 depth analysis. Spectral depth technique has being used by several authors in determining the 201 thickness of sediments within their restricted area of study. Aside from the spectral depth 202 method, other automated or semi automated techniques, like Euler Deconvolution, Improved 203 Source Parameter imaging (ISPI), Werner Deconvolution and tilt angle, can as well be used 204 in estimating depth to basement ( that is, the thickness of sediments). Therefore, for this 205 study, Spectral analysis is the basis for depth determination using the potential field 206 aeromagnetic data of the study area. It was opined <sup>19</sup> that Spectral analysis has proved to be a 207 powerful and convenient tool in the processing and interpretation of potential field 208 geophysical data. It seeks to describe the frequency content of a signal based on a finite set of 209 data. This technique is believed to provide rapid depth estimates from regularly spaced digital 210 211 field data, no geomagnetic or diurnal corrections are necessary as these remove only low wave number components and do not affect the depth estimates which are controlled by the 212 213 high wave number components of the observed field. This technique is based on the shape of the power spectrum for buried bodies with a susceptibility contrast. It is shown for simple 214 bodies <sup>20</sup> and for complex shaped bodies <sup>7</sup> that the depth to the center of mass of the body is 215 easily obtained from the power spectrum of the magnetic field. If the spectrum is plotted on 216 217 semi-log paper, the slope of the spectrum is equal to the depth to the center of mass. Extremely complex shapes and layering can, however, complicate the spectrum. The 218 spectrum gives information primarily about the location of the top and bottom of a magnetic 219 layer <sup>21</sup>. Nevertheless, this research follows the assumption that magnetic basement is 220 composed of a randomly distributed number of structures. Then by calculating the average of 221 the spectra, the depth for all anomalous sources is determined and this is equivalent to that for 222 a single body at the same depth. 223

# 224 FAST FOURIER TRANSFORM (FFT) PERFORMED ON EACH SPECTRAL 225 WINDOW

FFT mathematical modeling was performed on the 16 spectral aeromagnetic grids (or 226 windows) of 10 km by 10 km for spectral depth analysis using the FFT function in the data 227 analysis tool. For the FFT to be implemented, the 16 grids were imported into the Microcal 228 OriginPro8 software for scientific data analysis and processing. The FFT function was used 229 to separate the TMI data of the windows (called cells) into their frequency component and 230 energy (FFT magnitude or the absolute value of the FFT complex) spectrum. Although, the 231 data was smoothed using smoothing function of the Origin Pro 8. This was necessary since 232 the computed logarithmic energy generally reveals a slowly decreasing mean value within an 233 envelope of erratic rapid variations. The FFT was then performed on the Smoothed TMI data 234 for all the cells. Hence, the discrimination of the anomalies into its energy spectrum and 235 236 frequency component. To determine the spatial frequency component a sampling interval 1000 m was used applying square window with a one-sided spectrum type normalized to 237 238 Mean Square Amplitude (MSA). This setting output the Spatial Frequency domain (cycle/m), FFT Complex, and energy (magnitude) spectrum. The log of the energy (Log E) was then 239 240 plotted against the radial frequency in MS Excel as suggested by Spector and Grant (1970). A straight line is finally visually fit to the energy spectrum, usually in the higher and lower 241 frequency of the figure. The negative of slope of this line is equal to twice the depth (depth= 242 slope/2) to the center of mass of the bodies producing the magnetic field. After the depth has 243 been calculated over one window a new calculation is made over a new window. This 244 continues over the grid until all windows have had their radial spectra calculated and the 245 depths picked. 246

#### 247 ESTIMATION OF THE CURIE POINT DEPTH (CPD)

The method of Curie Point Depth determination utilizes spectrum analysis technique to 248 separate influences of the different body parameters in the observed magnetic anomaly field 249 <sup>11</sup>. Two basic methods for estimating the CPD have been stated. They include those that 250 examine the shape of isolated magnetic anomalies <sup>7</sup> and those that examine the statistical 251 properties and patterns of the anomalies <sup>22</sup>. These two methods, however, provide the 252 relationship between the spectrum of the magnetic anomalies and the depth of a magnetic 253 254 source by transforming the spatial data into frequency domain. This research adopted the method in which the top boundary and the centroid of magnetic sources were calculated from 255 256 the spectrum of magnetic anomalies which were used to estimate the basal depth of magnetic

- source. For the possible determination of the CPD, the residual data was divided into sixteen
- 258 (16) overlapping windows as shown below:



#### Fig. 4: Sixteen (16) overlapping windows for CPD determination via spectral analysis

261

262 Spectral analysis was thereafter performed on the sixteen windows. From the spectral depth 263 analysis the following steps were then undertaken

264  $\succ$  STEP ONE: Estimation of the depth to Centroid (Z<sub>C</sub>) of the magnetic sources from 265 the slope of the low frequency component part of the energy spectrum.

266  $\succ$  STEP TWO: Estimation of the depth to the top boundary (Z<sub>t</sub>) of magnetic sources 267 from the slope of the high frequency component part of the spectral segment.

268

269 The calculation of  $Z_C$  and  $Z_T$  then lead to the calculation of the basal depth  $Z_b$  using equation 270 4 and this is assumed <sup>23,24</sup> as the CPD

271

$$Z_{272} \quad Z_b = 2Z_C - Z_t$$

Where  $Z_b$  = the basal depth  $Z_c$  = the centroid depth HEAT FLOW AND THERMAL GRADIENT DETERMINATION A relation showing one dimensional heat conductive model was used for this study in order to estimate the heat flow and the geothermal gradient in the absence of a heat flow data. This model is based on the Fourier's law <sup>23</sup>. Fourier's law is mathematically expressed as <sup>23</sup>  $q = \sqrt{\frac{dT}{dZ}}$ Where q = the heat flow  $\Lambda$  = the coefficient of thermal conductivity The Curie temperature ( $\theta_c$ ) is defined <sup>10</sup> as Equating 5 and 6, we have -

dT/dZ = an assumed constant as no heat gain or loss above the crust and below the CPD 

$$\theta_{4} \quad \theta = \left[ \frac{dT}{dZ} \right] Z_{b} \qquad 6$$

$$300 \quad q = \Lambda \left[\frac{\theta}{Z_{\rm b}}\right]$$

Equation 4 was therefore used in this research work in determining the heat when the Curie point depths were known. For these estimations to be possible, a standard for curie point isotherm of 580°C and thermal conductivity of 2.5 Wm<sup>-1</sup>°C<sup>-1</sup> was used <sup>14</sup>. Finally, the geothermal gradient was determined using equation 6, hence we have: 

307 
$$\frac{dT}{dz} = \frac{\theta}{Z_{b}}$$

308 Where

309  $\frac{dT}{dZ}$  = geothermal gradient

310  $Z_b =$ the basal depth

311  $\theta$  = the standard curie point isotherm of 580<sup>°</sup>C

Also,  $\frac{dT}{dz}$  can be estimated by applying Fourier's model stated in equation 5, from equation 5 313 we have:

8

9

314

315 
$$\frac{dT}{dz} = \frac{q}{\Lambda}$$

316 Where

317 q =the heat flow

318  $\Lambda =$  thermal conductivity and is given as 2.5 Wm<sup>-1</sup>°C<sup>-1</sup> for igneous rocks

319

#### 320 RESULTS AND INTERPRETATION

321 In this work, quantitative analysis was achieved by using spectral depth analysis in computing the depth to magnetic basement on each of the spectral windows. Performed on 322 323 each of the windows is FFT. The FFT modeling aided in decomposing each of the window data set into its frequency components. Thereafter, plots showing log of energy (LogE) 324 325 versus the radial frequency were made. Based on the plots, table 1 which shows the centriod depth  $(Z_C)$  and the depth to the top boundary  $(Z_t)$  of the magnetic sources was generated. The 326  $Z_C$  is due to magnetic sources within the basement while the  $Z_T$  is due to magnetic sources 327 328 within the sedimentary section. In otherwords, Z<sub>C</sub> and Z<sub>t</sub> reflect magnetic sources due to deep and shallow seated features. Z<sub>C</sub> is a true reflection of Precambrian magnetic basement bodies 329 and these values ranges from -7.7 km to -12.0 km while Zt possibly depicts magnetic effect 330 due to short wavelength sources and these ranges from - 0.5 km m to - 3.1 km. 331

Nevertheless, the CPD were computed and this work found out that this varies between -14.9 km and -23 km while the geothermal gradient and the heat flow varies from -25.2 to -37.9 °Ckm<sup>-2</sup> and -63 to -97.3 mWm<sup>-2</sup> respectively. Summary of the various windows with their various parameters calculated are shown in Fig. 5.

## 339Table 1: Estimation of geothermal gradient and heat flow from CPD via spectral

340 analysis

|       |                |          |       |           |                | AVERA | GEOTHERM                    |            |
|-------|----------------|----------|-------|-----------|----------------|-------|-----------------------------|------------|
|       |                |          |       |           |                | GE    | AL                          | HEAT       |
| WINDO |                |          |       |           |                | DEPTH |                             | FLO        |
| W     | SLOPE          | SLOPE    |       | DEPTH(km) |                | (km)  | GRADIENT                    | W          |
|       |                |          |       |           |                |       | (mW                         |            |
|       | $\mathbf{M}_1$ | $M_2$    | Zc    | Zt        | Z <sub>b</sub> |       | $(^{\circ}\text{Ckm}^{-2})$ | $m^{-2}$ ) |
|       | -              | -        |       |           |                |       |                             |            |
|       | 23995          | 2011.4   |       |           |                |       |                             |            |
| W1    | .5             | 7        | -12   | -1        | -23            | -6.5  | -25.2                       | -63        |
|       |                | -        |       |           |                |       |                             |            |
|       | -              | 1361.9   |       |           |                |       |                             |            |
| W2    | 17762          | 1        | -8.9  | -0.7      | -17.1          | -4.8  | -33.9                       | -84.8      |
|       | -              | -        |       |           |                |       |                             |            |
|       | 23612          | 3218.7   |       |           |                |       |                             |            |
| W3    | .7             | 4        | -11.8 | -1.6      | -22            | -6.7  | -26.4                       | -65.9      |
|       | -              |          |       |           |                |       |                             |            |
|       | 18315          | -        |       |           |                |       |                             |            |
| W4    | .6             | 2817.4   | -9.2  | -1.4      | -17            | -5.3  | -34.1                       | -85.3      |
|       | -              | -        |       |           |                |       |                             |            |
|       | 19116          | 2461.3   | 0.4   |           | 10             |       |                             | 0.0 4      |
| W5    | .1             | 8        | -9.6  | -1.2      | -18            | -5.4  | -32.2                       | -80.6      |
|       | -              | -        |       |           |                |       |                             |            |
| NIC   | 20307          | 3801.7   | 10.0  | 1.0       | 10.5           | -     | 21.4                        | 70.4       |
| W6    | .3             | 6        | -10.2 | -1.9      | -18.5          | -6    | -31.4                       | -/8.4      |
|       | -              | -        |       |           |                |       |                             |            |
| W7    | 21503          | 2011.8   | 10.9  | 1         | 20.6           | 5.0   | 28.2                        | 70.4       |
| W /   | .2             | 1        | -10.8 | -1        | -20.6          | -5.9  | -28.2                       | -70.4      |
|       | -              | -        |       |           |                |       |                             |            |
| WO    | 19213          | 3322.0   | 06    | 16        | 176            | 56    | 22                          | 82.4       |
| VV ð  | .4             | 1        | -9.0  | -1.0      | -17.0          | -3.0  | -33                         | -82.4      |
|       | -              |          |       |           |                |       |                             |            |
| WO    | 22431<br>Q     | - 1704.2 | 11.2  | 0.0       | 21.5           | 6     | 77                          | 64.4       |
| VV 9  | .0             | 1704.3   | -11.2 | -0.9      | -21.3          | -0    | -27                         | -04.4      |
|       | - 19224        |          |       |           |                |       |                             |            |
| W10   | 10334          | 6231.6   | _0 2  | _3 1      | -153           | -6.1  | _38                         | _0/ 8      |
| VV 10 | .1             | 0231.0   | -9.2  | -3.1      | -13.3          | -0.1  | -30                         | -74.0      |
|       | - 15/10        | 1002 1   |       |           |                |       |                             |            |
| W11   | 2              | 8        | _77   | -0.5      | -14 9          | 1     | _38.0                       | _97 3      |
| ** 11 |                | 0        | -/./  | -0.5      | 17.7           | -+.1  | -30.7                       | 71.5       |

| W12 | -<br>17332<br>.5 | -<br>2234.3<br>3 | -8.7  | -1.1 | -16.3 | -4.9 | -35.6 | -89   |
|-----|------------------|------------------|-------|------|-------|------|-------|-------|
|     | -<br>21334       | -                |       |      |       |      |       |       |
| W13 | .1               | 5604.3           | 10.7  | -2.8 | -18.6 | -6.8 | -31.2 | -78   |
| W14 | 22332<br>.1      | 2101.3<br>8      | -11.2 | -1.1 | -21.3 | -6.2 | -27.2 | -68.1 |
| W15 | 20119            | 5770.2<br>8      | 10.1  | -2.9 | -17.3 | -6.5 | -33.5 | -83.8 |
| W16 | -<br>15773<br>.2 | 997.08           | -7.9  | -0.5 | -15.3 | -4.2 | -37.9 | -94.8 |

| WINDOW 1                      | WINDOW 2               | WINDOW 3                      | WINDOW 4                            |
|-------------------------------|------------------------|-------------------------------|-------------------------------------|
| Z <sub>b</sub> = -23          | Z <sub>b</sub> =-17.1  | <u>Z</u> <sub>b</sub> =-22    | Z <sub>b</sub> =- 17                |
| <u>Z</u> <sub>c</sub> = -12   | Z <sub>c</sub> = -8.9  | <u>Z</u> <sub>c</sub> = −11.8 | <u>Z</u> <sub>c</sub> = -9.2        |
| $Z_t = -1$                    | $Z_t = -0.7$           | $Z_t = -1.6$                  | $Z_t = -1.4$                        |
|                               |                        |                               |                                     |
| WINDOW 5                      | WINDOW 6               | WINDOW 7                      | WINDOW 8                            |
| Z <sub>b</sub> =-18           | Z <sub>b</sub> = -18.5 | Z <sub>b</sub> = -20.6        | <mark>Z</mark> <sub>b</sub> = −17.6 |
| <mark>Z</mark> ₅ = -9.6       | Z <sub>c</sub> = -10.2 | Z <sub>c</sub> = -10.8        | <mark>Z</mark> c=-9.6               |
| $Z_{t} = -1.2$                | <u>Zt</u> = -1.9       | $Z_{t} = -1.0$                | $Z_{t} = -1.6$                      |
|                               |                        |                               |                                     |
| WINDOW 9                      | WINDOW 10              | WINDOW 11                     | WINDOW 12                           |
| Z <sub>b</sub> = -21.5        | Z <sub>b</sub> =-15.3  | Z <sub>b</sub> =-14.9         | Z <sub>b</sub> = -16.3              |
| <u>Z</u> <sub>c</sub> = -11.2 | <mark>Zc</mark> = -9.2 | <u>Z</u> <sub>c</sub> = -7.7  | <mark>∠</mark> = -8.7               |
| $Z_t = -0.9$                  | $Z_{t} = -3.1$         | $Z_{t} = -0.5$                | $Z_{t} = -1.1$                      |
|                               |                        |                               |                                     |
| WINDOW 13                     | WINDOW 14              | WINDOW 15                     | WINDOW 16                           |
| Z <sub>b</sub> = -18.6        | Z <sub>b</sub> = -21.3 | Z <sub>b</sub> = -17.3        | Z <sub>b</sub> = -15.3              |
| Z <sub>c</sub> = -10.7        | Z <sub>c</sub> = -11.2 | $Z_{c} = -10.1$               | $Z_{c} = -7.9$                      |
| $Z_{t} = -2.8$                | $Z_{t} = -1.1$         | $Z_{t} = -2.9$                | $Z_{t} = -0.5$                      |
|                               |                        |                               |                                     |
|                               |                        |                               |                                     |

## 344 Fig. 5: Succinct view of the various windows and their various depths (km)

#### 345 **DISCUSSION OF FINDINGS**

Graph of the spectral energies revealed that the depth due to the deep seated anomalous 346 sources or the centriod depth (Z<sub>c</sub>) varies from 7.7 km to 12.0 km. Conversely, the depth due 347 to the shallow bodies or due to the top boundary of magnetic sources ranges between 0.5 km 348 m to 3.1 km while the corresponding CPD ranges from 14.9 km to 23.0 km. The true or 349 average depths for each of the respective windows were calculated and this varies between 350 4.1 km and 6.1 km. Generally, a true depth of 5.3 km was ascertained within the study area. 351 The depth value of 5.3 km suggests relative sedimentary thickness. These results compare 352 favourably with the results of other researchers<sup>1</sup> within the lower Benue trough. Although 353 with information on the crustal temperature missing in his work, He obtained the depths due 354 to the deep seated sources to range from 7.2 km to 13.0 km while the depths values for the 355 shallow sources varies from 0.4 km to 3.9 km, but on the average, he obtained a true depth of 356 7 km. Within parts of lower Benue trough, a thickness value of 5.6 km is believed to exist <sup>25</sup>. 357 This result is not a far cry from what was obtained for this study. Also, the CPD values <sup>13</sup> 358 were obtained within the area. The CPD depths were not in variance with those CPD values 359 obtained for this study. However within the study area, the CPD within the study area vary 360 361 from 23.80 km to 28.70 km. Shallower CPD can be seen towards the southeastern portion of the study area and this falls within the Oji river settlement province perceived to be the 362 crystalline basement area. This is a possible reflection of the thinning of the crust under 363 Benue rift <sup>13</sup>. They further stated that the basement area is as a result of the upwelling of 364 magma on Cameroon Volcanic Line (CVL) during the tertiary period. Hence, the possibility 365 of igneous intrusive that provide appropriate geothermal energy needed for the maturation of 366 hydrocarbon found within that region. This particular findings is therefore in support of the 367 fact that geo thermally active regions are usually shallower as it consist of igneous intrusive 368 that could foster or be detrimental to hydrocarbon maturation. The heat flow within the study 369 area varies between -64.4 mWm<sup>-2</sup> and -97.3 mWm<sup>-2</sup> but with an average of about 80.1 mWm<sup>-</sup> 370 <sup>2</sup> while the geothermal gradient varies from 25.2 °Ckm<sup>-2</sup> to 38.9 °Ckm<sup>-2</sup> but with an average 371 of about 32.1 °Ckm<sup>-2</sup> existing within the area under review. The heat flow values and the 372 geothermal gradient values obtained further validate the fact that the shallowest depth 373 occurring within window eleven (11) is a good geo thermally active area. Also, windows 2, 4, 374 8, 10, 12, 15 and 16 are possible geothermal resource areas. These areas will not, thus, be 375 376 much productive in terms of oil and gas.

#### 378 CONCLUSION

The study area is characterized with high heat flow and geothermal gradient with relatively low magnetic intrusions which could be detrimental to the quantity of hydrocarbon exploration. Therefore, the high heat flow and moderate geothermal gradient observed in this study are possible tectonically induced rifting and magmatism that occurred during Pan– African Orogeny.

- 384 385 REFERENCES Umeanoh, D.C. 2015 Structural interpretation of aeromagnetic data over parts of 386 1. lower benue trough, Nigeria, Msc thesis, University of PortHarcourt, Choba, Nigeria, 387 (Unpubl). 388 2. Nettleton, L. L. 1971. Elementary Gravity and Magnetics for geologists and 389 seismologists. Society of Exploration Geophysicists, Monograph series, (1) 121. 390 3. Reynolds, L. R., Rosenbaum, J. G., Hudson, M.R., and Fishman, N.S. 1990. Rock 391 Magnetism, the distribution of magnetic minerals in the earth crust and aeromagnetic 392 anomalies: U.S. Geological Survey and aeromagnetic anomalies: U.S. Geological 393 Survey, Bulletin 1924, 24-45. 394 395 4. Tselentis, G.A. 1991. An attempt to define Curie depth in Greece from Aeromagnetic and heat flow data. PAGEOPH, 36(1), 87-101. 396 397 5. Ibrahim, A., Halil, I. and Ali, K. 2015. Curie-point depth map of Turkey. Geophys. J. 398 Int., 162(12), 633–640. 399 400 6. Vacquier, V. and Affleck, J. 1941. A computation of the average depth to the bottom 401 402 of the Earth' magnetic crust, based on a statistical study of local magnetic anomalies. *Trans. Am. geophys. Un.*, 2, 446 – 450. 403 404 Bhattacharyya, B.K. and Leu, L.K. 1975. Analysis of magnetic anomalies over 7. 405 Yellowstone National Park: mapping of Curie point isothermal surface for geothermal 406 reconnaissance, J. geophys. Res., 80, 4461 - 4465. 407 408 8. Shuey, R.T., Schellinger, D.K., Tripp, A.C. and Alley, L.B. 1977. Curie depth 409 determination from aeromagnetic spectra. *Geophys. J. R. astr. Soc.*, **50**, 75–101. 410 411 9. Connard, G., Couch, R. and Gemperle, M. 1983. Analysis of aeromagnetic 412 measurements from the Cascade Range in central Oregon. Geophysics, 48, 376–390. 413 414 415 Tanaka, A., Okuba, Y. and Matsubayashi, O. 1999. Curie point depth based on 416 10. spectrum analysis of the magnetic anomaly data in East and Southeast Asia. 417
  - 18

*Tectonophysics*, **306**, 461–470.

419 Elleta, B.E. and Udensi, E.E. 2012. Investigation of the Curie point Isotherm from 420 11 the Magnetic Fields of Eastern Sector of Central Nigeria. Geosciences, 2(4): 101-106. 421 422 Hisarlı, M. 1995. Determination of the Curie point depths in Edremit - Susurluk 423 12 424 region (Turkey). *Jeofizik*, **1**(2), 111–117. 425 13 Nuri, D. M., Timur U. Z., Mumtaz, H. and Naci, O. 2005. Curie Point Depth 426 variations to infer thermal structure of the crust at the African-Eurasian convergence 427 zone, SW Turkey. Journ.Earth planets Space. 57(20), 373-383. 428 429 430 14. Nwankwo, L.I., Olasehinde, P.I. and Akoshile, C.O. 2011. Heat flow anomalies from the spectral analysis of Airborne Magnetic data of Nupe Basin, Nigeria. Asian Journal 431 of Earth Sciences. 1. (1), 1-6. 432 433 Nur, A., Ofoegbu C.O. and Onuoha K.M. 1999. Estimation of the depth to the Curie 434 15. point Isotherm in the upper Benue trough, Nigeria. Jour. Min. Geol., 35 (1), 53 - 60. 435 436 Onvewuchi, A.R. 2011 Interpretation of Aeromagnetic and Landsat Imageries of 437 16. Okposi, Ebonyi State, SE Nigeria, Msc thesis, University of PortHarcourt, Choba, 438 439 Rivers State, (Unpubl). 440 17. Murat, R.C. 1972. Stratigraphy and paleogeography of the Cretaceous and lower Tertiary in southern Nigeria. African Geology, University of Ibadan Press, 251-266. 441 Nwozor, K.K., Chiaghanam, O.I., Egbuachor, C.J. and Onyekuru, S.O. (2012): 442 18. 443 Surface Geophysics Character of Maastrichtian - Danian Sediments in parts of Udi-Ezeagu Area, Southern Anambra Basin, Nigeria. Archives of Applied Science 444 Research, 4 (4):1609-1617. 445 19. Hahn, A.; Kind, E.G. and Mishra, D.C. 1976. Depth estimation of Magnetic Sources 446 447 by means of Fouries amplitude spectra. Geophysical Prosp. 24: 278-308. 448 Odegard, M.E. and Berg, J.W. 1975. Gravity interpretation using Fourier integral. 449 20. Geophysics, 32(4), 1-4. 450 21 Blakely, R.J. 1996. Potential theory in gravity and Magnetic Applications. Cambridge 451 University Press, New York, 435-567. 452 22 Spector, A. and Grant, F.S. 1970. Statistical models for interpreting aeromagnetic 453 data, Geophysics, 35, 293-302. 454 455 23 Kasidi, S. and Nur, A. 2012. Curie depth isotherm deduced from spectral analysis of 456 Magnetic data over sarti and environs of North-Eastern Nigeria. Scholarly J. 457 Biotechnol. 1(3), 49-56. 458 459 24 Okubo, Y.J., Graf, R., Hansen, R.O., Ogawa, K. and Tsu, H. 1985. Curie point depth 460 of the Island of Kyushu and surrounding areas. Japan Geophysic., 53, 481-491. 461

463 25 Onwuemesi, A.G. 1996. One-dimensional spectral analysis of aeromagnetic anomalies
464 and curie depth isotherm in the Anambra basin of Nigeria. J. Geodynamics, 23(2), 95465 107.