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The electrodynamic vacuum field theory approach and the electron inertia problem revisiting

4 5 It is a review of some new electrodynamics models of interacting charged point 6 particles and related with them fundamental physical aspects, motivated by the classical 7 A.M.Amper's magnetic and H.Lorentz force laws, as well as O. Jefimenko electromagnetic field 8 expressions. Based on the suitably devised vacuum field theory approach the Lagrangian and 9 Hamiltonian reformulations of some alternative classical electrodynamics models are analyzed 10 in details. A problem closely related to the radiation reaction force is analyzed aiming to explain 11 the Wheeler and Feynman reaction radiation mechanism, well known as the absorption radiation theory, and strongly dependent on the Mach type interaction of a charged point 12 13 particle in an ambient vacuum electromagnetic medium. There are discussed some 14 relationships between this problem and the one derived within the context of the vacuum field theory approach. The R.Feynman's "heretical" approach to deriving the Lorentz force based 15 16 Maxwell electromagnetic equations is also revisited, its complete legacy is argued both by 17 means of the geometric considerations and its deep relation with the devised vacuum field 18 theory approach. Based on completely standard reasonings, we reanalyze the Feynman's 19 derivation from the classical Lagrangian and Hamiltonian points of view and construct its 20 nontrivial relativistic generalization compatible with the vacuum field theory approach. The 21 electron inertia problem is reanalyzed within the Lagrangian-Hamiltonian formalisms and the 22 related Feynman proper time paradigm. The validity of the Abraham-Lorentz electromagnetic 23 electron mass origin hypothesis within the shell charged model is argued. The electron stability 24 in the framework of the electromagnetic tension-energy compensation principle is analyzed.

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Cet article présente un réexamen de certains nouveaux modèles électrodynamiques 26 27 d'interaction entre particules chargées ponctuelles, et en lien avec des aspects physiques 28 fondamentales, motivés par les lois magnétiques classiques de A.M. Ampère et par les forces 29 classiques de H.Lorentz, ainsi que par les formulations du champ électromagnétique décrites par O. Jefimenko. Sur la base d'une formulation adéquate de la théorie des champs en vide 30 31 quantique, les reformulations Lagrangiennes et Hamiltoniennes de certains modèles alternatifs 32 de l'électrodynamique classique sont analysés en profondeur. Un problème étroitement lié à la 33 force de réaction de rayonnement est analysé pour objectif d'expliquer le mécanisme de 34 Wheeler et Feynman de réaction au rayonnement, bien connu comme la théorie 35 d'amortissement de radiation, et dépend fortement de l'interaction de type Mach de particules 36 ponctuelles chargées dans un milieu électromagnétique en vide ambiant. Certains rapports 37 entre ce problème et celui obtenu dans le cadre de l'approche de la théorie des champs en vide quantique sont examinés. L'approche "hérétique" de R.Feynman qui consiste à dériver la force 38 39 de Lorentz depuis équations électromagnétiques de Maxwell est également revisitée, et son 40 approche est justifiée à la fois par des considérations géométriques et sa relation profonde 41 avec l'approche de la théorie des champs à vide quantique. Sur la base de raisonnements 42 complètement standards, nous réanalysons la dérivation de Feynman des points de vue 43 Lagrangiens et Hamiltoniens classiques et construisons sa généralisation relativiste non triviale 44 compatible avec l'approche de la théorie des champs en vide quantique. Le problème de l'inertie des électrons est réanalysé dans les formalismes de Lagrange et Hamilton et dans le
paradigme de Feynman en temps propre correspondant. La validité de l'hypothèse d'AbrahamLorentz sur l'origine électromagnétique de la masse de l'électron dans le modèle de couche
électronique est soutenu. La stabilité de l'électron dans le cadre du principe de compensation
tension-énergie électromagnétique est analysée.

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51 *Keywords*: Amper law, Lorentz type force, Lorenz constraint, Vacuum field theory 52 approach, Maxwell electromagnetic equation,Lagrangian and Hamiltonian formalisms, Fock 53 many-temporal approach, Jefimenko equations, Quantum self-interactifermi model, Radiation 54 theory, Feynman's proper time approach, Abraham-Lorentz electron mass problem.

- 55 PACS: 11.10.Ef, 11.15.Kc, 11.10.-z; 11.15.-q, 11.10.Wx, 05.30.-d
- 56 Received: (05 January 2016)
 57 Revised: (12 June 2016)

Revised: (12 June 2016) Accepted (Day Month 2016)

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1. Classical relativistic electrodynamics models revisiting: Lagrangian and Hamiltonian analysis

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1.1. Introductory setting

66 The Maxwell's equations serve as foundational [1] [2] [3] [4] [5] to the whole modern 67 classical and quantum electromagnetic theory and electrodynamics. They are the cornerstone 68 of a myriad of technologies and are basic to the understanding of innumerable effects. Yet 69 there are a few effects or physical phenomena that cannot be explained [6] [7] [8] [9] [10] [11] 70 [12] [13] within the conventional Maxwell theory. It is important to note here that in [8] [14] 71 [15] [16] [17] there is argued that the Maxwell equations as themselves do not determine 72 causal related to each other electric and magnetic fields, which prove, in reality, to be 73 generated independently by an external charge and current distributions: "There is a 74 widespread interpretation of Maxwell's equations indicating that spatially varying electric and 75 magnetic fields can cause each other to change in time, thus giving rise to a propagating electromagnetic wave... However, Jefimenko's equations show an alternative point of view [3]. 76 Jefimenko says: "...neither Maxwell's equations nor their solutions indicate an existence of 77 78 causal links between electric and magnetic fields. Therefore, we must conclude that an 79 electromagnetic field is a *dual entity* always having an electric and a magnetic component 80 simultaneously created by their common sources: time-variable electric charges and currents." 81 Essential features of these equations are easily observed which are that the right hand sides involve "retarded" time which reflects the "causality" of the expressions. In other words, the 82 83 left side of each equation is actually "caused" by the right side, unlike the normal differential 84 expressions for Maxwell's equations, where both sides take place simultaneously. In the typical expressions for Maxwell's equations there is no doubt that both sides are equal to each other, 85 but as Jefimenko notes [3], "... since each of these equations connects quantities simultaneous 86 87 in time, none of these equations can represent a causal relation." The second feature is that the

expression for (electric field) E does not depend upon (magnetic field) B and vice versa. 88 Hence, it is impossible for E and B fields to be "creating" each other. Charge density and 89 90 current density are creating them both." As the Jefimenko's equations for the electric field E91 and the magnetic field B directly follow from the classical retarded Lienard-Wiechert 92 potentials, generated by physically real external charge and current distributions, one naturally 93 infers that these potentials also present suitably interpreted physical field entities mutually 94 related to their sources. This way of thinking proved to be, from the physical point of view, very 95 fruitful, having brought about a new vacuum field theory approach [18] [19] to alternative 96 explaining the nature of the fundamental Maxwell equations and related electrodynamic 97 phenomena.

98 We start from detailed revisiting the classical A.M. Ampere's law in electrodynamics and 99 show that main inferences suggested by physicists of the former centuries can be strongly 100 extended for them to agree more exactly with many modern both theoretical achievements 101 and experimental results concerning the fundamental relationship of electrodynamic 102 phenomena with the physical structure of vacuum as their principal carrier.

103 The important theoretical physical principles, characterizing the related electrodynamic 104 vacuum field structure, we discuss subject to different charged point particle dynamics, based 105 on the fundamental least action principle. In particular, the main classical relativistic 106 relationships, characterizing the charge point particle dynamics, we obtain by means of the 107 least action principle within the Feynman's approach to the Maxwell electromagnetic equations and the related Lorentz type force derivation. Moreover, for each of the least action principles 108 109 constructed in the work, we describe the corresponding Hamiltonian pictures and present the 110 related energy conservation laws. The elementary point charged particle, like electron, mass 111 problem was inspiring many physicists [20] from the past as J. J. Thompson, G.G. Stokes, H.A. 112 Lorentz, E. Mach, M. Abraham, P.A. M. Dirac, G.A. Schott and others. Nonetheless, their studies 113 have not given rise to a clear explanation of this phenomenon that stimulated new researchers to tackle it from different approaches based on new ideas stemming both from the classical 114 115 Maxwell-Lorentz electromagnetic theory, as in [1] [12] [21] [22] [23] [24] [25] [26] [27] [28] [29] 116 [30] [31] [32] [33] [34] [35] [36] [37] [38] [39], and modern quantum field theories of Yang-Mills 117 and Higgs type, as in [40] [41] [42] [43] and others, whose recent and extensive review is done 118 in [44].

119 We will mostly concentrate on detail analysis and consequences of the Feynman proper 120 time paradigm [1] [22] [45] [46] subject to deriving the electromagnetic Maxwell equations 121 and the related Lorentz like force expression considered from the vacuum field theory 122 approach, developed in works [47] [48] [49] [50] [51], and further, on its applications to the 123 electromagnetic mass origin problem. Our treatment of this and related problems, based on 124 the least action principle within the Feynman proper time paradigm [1], has allowed to construct the respectively modified Lorentz type equation for a moving in space and radiating 125 126 energy charged point particle. Our analysis also elucidates, in particular, the computations of 127 the self-interacting electron mass term in [29], where there was proposed a not proper solution 128 to the well known classical Abraham-Lorentz [52] [53] [54] [55] and Dirac [56] electron 129 electromagnetic "4/3-electron mass" problem. As a result of our scrutinized studying the 130 classical electromagnetic mass problem we have stated that it can be satisfactory solved within 131 the classical H. Lorentz and M. Abraham reasonings augmented with the additional electron

stability condition, which was not taken before into account yet appeared to be very important for balancing the related electromagnetic field and mechanical electron momenta. The latter, following recent enough works [31] [35], devoted to analyzing the electron charged shell model, can be realized within there suggested *pressure-energy compensation principle*, suitably applied to the ambient electromagnetic energy fluctuations and the own electrostatic Coulomb electron energy.

138 In our investigation, we were in part inspired by works [35] [39] [43] [44] [57] [58] [59] 139 to solving the classical problem of reconciling gravitational and electrodynamic charges within 140 the Mach-Einstein ether paradigm. First, we will revisit the classical Mach-Einstein type 141 relativistic electrodynamics of a moving charged point particle, and second, we study the 142 resulting electrodynamic theories associated with our vacuum potential field dynamical 143 equations (31) and (32), making use of the fundamental Lagrangian and Hamiltonian formalisms 144 which were specially devised in [50] [51].

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1.2. Classical Maxwell equations and their electromagnetic potentials form revisiting

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149 As the classical Lorentz force expression with respect to an arbitrary inertial reference 150 frame is related with many theoretical and experimental controversies, such as the relativistic 151 potential energy impact into the charged point particle mass, the Aharonov-Bohm effect [60] 152 and the Abraham-Lorentz-Dirac radiation force [2] [5] [6] expression, the analysis of its 153 structure subject to the assumed vacuum field medium structure is a very interesting and 154 important problem, which was discussed by many physicists including E. Fermi, G. Schott, R. 155 Feynman, F. Dyson [1] [45] [46] [61] [62] [63] and many others. To describe the essence of the 156 electrodynamic problems related with the description of a charged point particle dynamics 157 under external electromagnetic field, let us begin with analyzing the classical Lorentz force 158 expression

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$$dp / dt = F_L := \xi E + \xi u \times B, \tag{1}$$

160 where $\xi \in \mathbb{R}$ is a particle electric charge, $u \in T(\mathbb{R}^3)$ is its velocity [47] [64] vector, expressed 161 here in the light speed c units,

$$E := -\partial A / \partial t - \nabla \varphi \tag{2}$$

163 is the corresponding external electric field and

$$B := \nabla \times A \tag{3}$$

is the corresponding external magnetic field, acting on the charged particle, expressed in terms 165 of suitable vector $A: M^4 \to E^3$ and scalar $\varphi: M^4 \to R$ potentials. Here, as before, the sign 166 " ∇ " is the standard gradient operator with respect to the spatial variable $r \in E^3$, " \times " is the 167 usual vector product in three-dimensional Euclidean vector space $E^3 := (R^3, <,, >)$, which is 168 naturally endowed with the classical scalar product $\langle \cdot, \cdot \rangle$. These potentials are defined on the 169 Minkowski space M^4 ; $R \times E^3$, which models a chosen laboratory reference frame K. Now, it 170 171 is a well known fact [1] [5] [37] [65] that the force expression (1) does not take into account the 172 dual influence of the charged particle on the electromagnetic field and should be considered 173 valid only if the particle charge $\xi \to 0$. This also means that expression (1) cannot be used for

studying the interaction between two different moving charged point particles, as was 174 175 pedagogically demonstrated in classical manuals [1] [5]. As the classical Lorentz force 176 expression (1) is a natural consequence of the interaction of a charged point particle with an 177 ambient electromagnetic field, its corresponding derivation based on the general principles of 178 dynamics, was deeply analyzed by R. Feynman and F. Dyson [1] [45] [46].

179 Taking this into account, it is natural to reanalyze this problem from the classical, taking only into account the Maxwell-Faraday wave theory aspect, specifying the corresponding 180 vacuum field medium. Other questionable inferences from the classical electrodynamics 181 182 theory, which strongly motivated the analysis in this work, are related both with an alternative interpretation of the well-known Lorenz condition, imposed on the four-vector of 183 electromagnetic observable potentials $(\varphi, A): M^4 \to T^*(M^4)$ and the classical Lagrangian 184 formulation [5] of charged particle dynamics under external electromagnetic field. The 185 186 Lagrangian approach latter is strongly dependent on an important Einstein notion of the proper 187 reference frame K_{τ} and the related least action principle, so before explaining it in more detail, we first to analyze the classical Maxwell electromagnetic theory from a strictly dynamical point 188 189 of view.

190 Let us consider with respect to a laboratory reference frame K, the additional Lorenz 191 condition

$$\partial \varphi / \partial t + < \nabla, A \ge 0, \tag{4}$$

193 a priori assumed the Lorentz invariant wave scalar field equation

$$\partial^2 \varphi / \partial t^2 - \nabla^2 \varphi = \rho \tag{5}$$

and the charge continuity equation 195

$$\partial \rho / \partial t + \langle \nabla, J \rangle = 0, \tag{6}$$

where $\rho: M^4 \to R$ and $J: M^4 \to E^3$ are, respectively, the charge and current densities of the 197 ambient matter. Then one can derive [18] [51] that the Lorentz invariant wave equation 198 $\partial^2 A / \partial t^2 - \nabla^2 A = J$ 199 (7)

and the classical electromagnetic Maxwell field equations [1] [2] [5] [65] [66]

$$\nabla \times E + \partial B / \partial t = 0, < \nabla, E >= \rho,$$
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(8)

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 $\nabla \times B - \partial E / \partial t = J. \langle \nabla, B \rangle = 0.$

hold for all $(t,r) \in M^4$ with respect to the chosen laboratory reference frame K_t . As it was 202 shown by O.D. Jefimenko [3] [4], the corresponding solutions to (8) for the electric 203 204 $E: M^4 \to E^3$ and magnetic $B: M^4 \to E^3$ fields can be represented (in the light speed c=1units) by means of the following causally independent to each other field expressions 205

$$E(t,r) = \frac{1}{4\pi} \int_{\mathbb{R}^3} \left[\left(\frac{\rho(t_r, r')}{|r - r'|^3} + \frac{1}{|r - r'|^2} \frac{\partial \rho(t_r, r')}{\partial t} \right) (r - r') - \frac{1}{|r - r'|^2} \frac{\partial J(t_r, r')}{\partial t} \right] d^3r',$$
(9)

$$B(t,r) = \frac{1}{4\pi} \int_{\mathbb{R}^3} \left[\frac{J(t_r, r')}{|r - r'|^3} + \frac{1}{|r - r'|^2} \frac{\partial J(t_r, r')}{\partial t} \right] \times (r - r') d^3 r',$$

where $(t,r) \in M^4$ and $t_r = t - |r - r'|$ is the retarded time. The result (9) was based on direct 207 208 derivation from the classical Lienard-Wiechert potentials [2] [3] solving the field equations (5) 209 and (7), causally depending on the corresponding charge and current distributions. Based 210 strongly on this fact in [3] [4] there was argued from physical point of view that related with 211 equations (5) and (7) electric and magnetic potentials really constitute some suitably 212 interpreted physical entities, in contrast to the usual statements [1] [2] [5] about their pure 213 mathematical origin.

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214 It is worth to notice here that, inversely, Maxwell's equations (8) do not directly reduce, 215 via definitions (2) and (3), to the wave field equations (5) and (7) without the Lorenz condition 216 (4). This fact and reasonings presented above are very important: they suggest that, when it 217 comes to choose main governing equations, it proves to be natural replacing the Maxwell's 218 equations (8) with the electric potential field equation (5), the Lorenz condition (4) and the 219 charge continuity equation (6). To make the equivalence statement, claimed above, more 220 transparent we formulate it as the following proposition.

- 222 Proposition 1. The Lorentz invariant wave equation (5) together with the Lorenz condition (4) for the observable potentials $(\varphi, A): M^4 \to T^*(M^4)$ and the charge continuity 223 224 relationship (6) are completely equivalent to the Maxwell field equations (8).
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Proof. Substituting (4), into (5), one easily obtains 226 $\partial^2 \varphi / \partial t^2 = - \langle \nabla, \partial A / \partial t \rangle = \langle \nabla, \nabla \varphi \rangle + \rho,$ 227 (10)228 which implies the gradient expression 229 $< \nabla, -\partial A / \partial t - \nabla \varphi >= \rho.$ (11)

230 Taking into account the electric field definition (2), expression (11) reduces to

- 231 $\langle \nabla, E \rangle = \rho$,
- 232 which is the second of the first pair of Maxwell's equations (8).
- 233 Now upon applying $\nabla \times$ to definition (2), we find, owing to definition (3), that 234
 - $\nabla \times E + \partial B / \partial t = 0$, (13)

235 which is the first pair of the Maxwell equations (8). Having differentiated with respect to the 236 temporal variable $t \in \mathbb{R}$ the equation (5) and taken into account the charge continuity equation 237 (6), one finds that

$$\langle \nabla, \partial^2 A / \partial t^2 - \nabla^2 A - J \rangle = 0.$$
⁽¹⁴⁾

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(12)

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The latter is equivalent to the wave equation (7) if to observe that the current vector $J: M^4 \to E^3$ is defined by means of the charge continuity equation (6) up to a vector function $\nabla \times S: M^4 \to E^3$. Now applying operation $\nabla \times$ to the definition (3), owing to the wave equation (7) one obtains

$$\nabla \times B = \nabla \times (\nabla \times A) = \nabla < \nabla, A > -\nabla^2 A =$$

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$$= -\nabla(\partial \varphi / \partial t) - \partial^2 A / \partial t^2 + (\partial^2 A / \partial t^2 - \nabla^2 A) =$$
(15)

$$= \frac{\partial}{\partial t} (-\nabla \varphi - \partial A / \partial t) + J = \partial E / \partial t + J,$$

244 leading directly to

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which is the first of the second pair of the Maxwell equations (8). The final *"no magnetic charge*" equation

 $<\nabla, B>=<\nabla, \nabla \times A>=0,$

 $\nabla \times B = \partial E / \partial t + J,$

in (8) follows directly from the elementary identity $\langle \nabla, \nabla \times \rangle = 0$, thereby completing the proof.

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This proposition allows us to consider the observable potential functions 251 252 $(\varphi, A): M^4 \to T^*(M^4)$ as fundamental ingredients of the ambient vacuum field medium, by 253 means of which we can try to describe the related physical behavior of charged point particles imbedded in space-time M^4 . As there was still written by J.K. Maxwell [67]: "The conception of 254 such a quantity, on the changes of which, and not on its absolute magnitude, the induction 255 256 currents depends, occurred to Faraday at an early stage of his researches. He observed that the 257 secondary circuit, when at rest in an electromagnetic field which remains of constant intensity, does not show any electrical effect, whereas, if the same state of the field had been suddenly 258 259 produced, there would have been a current. Again, if the primary circuit is removed from the field, or the magnetic forces abolished, there is a current of the opposite kind. He therefore 260 261 recognized in the secondary circuit, when in the electromagnetic field, a " peculiar electrical 262 condition of matter' to which he gave the name of Electrotonic State." The following 263 observation provides a strong support of this reasonings within this vacuum field theory 264 approach:

265 **Observation.** The Lorenz condition (4) actually means that the scalar potential field 266 $\varphi: M^4 \to \mathbb{R}$ continuity relationship, whose origin lies in some new field conservation law, 267 characterizes the deep intrinsic structure of the vacuum field medium.

To make this observation more transparent and precise, let us recall the definition [1] [5] [65] [66] of the electric current $J: M^4 \to E^3$ in the dynamical form

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$$J := \rho u, \tag{16}$$

where the vector $u \in T(\mathbb{R}^3)$ is the corresponding charge velocity. Thus, the following continuity relationship

$$\partial \rho / \partial t + \langle \nabla, \rho u \rangle = 0 \tag{17}$$

274 holds, which can easily be rewritten Error! Reference source not found. as the integral

275 conservation law

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$$\frac{d}{dt} \int_{\Omega_t} \rho(t, r) d^3 r = 0 \tag{18}$$

for the charge inside of any bounded domain $\Omega_{t} \subset E^{3}$, moving in the space-time M^{4} with 277 respect to the natural evolution equation for the moving charge system 278

$$dr \,/\, dt := u. \tag{19}$$

280 Following the above reasoning, we obtain the following result.

282 **Proposition 2.** The Lorenz condition (4) is equivalent to the integral conservation law

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 $\frac{d}{dt}\int_{\Omega_t} \varphi(t,r) d^3r = 0,$ (20)

where $\Omega_t \subset E^3$ is any bounded domain, moving with respect to the charged point particle ξ 284 285 evolution equation

$$dr / dt = u(t, r), \tag{21}$$

which represents the velocity vector of the related local potential field changes propagating in 287 the Minkowski space-time M^4 . Moreover, for a particle with the distributed charge density 288

 $\rho: M^4 \rightarrow \mathbb{R}$, the following Umov type local energy conservation relationship 289

$$\frac{d}{dt} \int_{\Omega_t} \frac{\rho(t,r)\varphi(t,r)}{(1-|u(t,r)|^2)^{1/2}} d^3r = 0$$
⁽²²⁾

291 holds for any $t \in \mathbf{R}$.

- 292 Proof. Consider first the corresponding solutions to potential field equations (5), taking into 293
- 294 account condition (16). Owing to the standard results from [1] [5], one finds that 295

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$$= \varphi u,$$
 (23)

296 which gives rise to the following form of the Lorenz condition (4):

$$\partial \varphi / \partial t + \langle \nabla, \varphi u \rangle = 0, \tag{24}$$

298 This obviously can be rewritten [68] as the integral conservation law (20), so the expression (20) 299 is stated.

300 To state the local energy conservation relationship (22) it is necessary to combine the 301 conditions (17), (24) and find that

$$\partial(\rho\varphi) / \partial t + \langle u, \nabla(\rho\varphi) \rangle + 2\rho\varphi \langle \nabla, u \rangle = 0.$$
⁽²⁵⁾

Taking into account that the infinitesimal volume transformation $d^3r = \chi(t,r)d^3r_0$, where the 303 Jacobian $\chi(t,r) := |\partial r(t;r_0) / \partial r_0|$ of the corresponding transformation $r: \Omega_{t_0} \to \Omega_t$, induced 304 by the Cauchy problem for the differential relationship (21) for any $t \in \mathbb{R}$, satisfies the 305 306 evolution equation 307

$$d\chi / dt = <\nabla, u > \chi, \tag{26}$$

easily following from (21), and applying to the equality (25) the operator $\int_{\Omega_{t_0}} (...) \chi^2 d^3 r_0$, one 308

309 obtains that $0 = \int_{\Omega_{t_0}} \frac{d}{dt} (\rho \varphi \chi^2) d^3 r_0 = \frac{d}{dt} \int_{\Omega_{t_0}} (\rho \varphi \chi) J d^3 r_0 =$ (27)

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$$= \frac{d}{dt} \int_{\Omega_t} (\rho \varphi \chi) d^3 r := \frac{d}{dt} \mathbf{E}(\xi; \Omega_t).$$

Here we denoted the conserved charge $\xi := \int_{\Omega_{t}} \rho(t, r) d^{3}r$ and the local energy conservation 311 quantity $E(\xi;\Omega_t)$: = $\int_{\Omega_t} (\rho \varphi \chi) d^3 r = E(\xi;\Omega_{t_0}), t \in \mathbb{R}$. The latter quantity can be simplified, 312 owing to the infinitesimal Lorentz invariance four-volume measure relationship 313 $d^{3}r(t,r_{0}) \wedge dt = d^{3}r_{0} \wedge dt_{0}$, where variables $(t,r) \in \mathbf{R}_{t} \times \Omega_{t} \subset M^{4}$ are, within the present 314 context, taken with respect to the moving reference frame K_{i} , related to the infinitesimal 315 charge quantity $d\xi(t,r) := \rho(t,r)d^3r$, and variables $(t_0,r_0) \in \mathbb{R}_{t_0} \times \Omega_{t_0} \subset M^4$ are taken with 316 respect to the laboratory reference frame K_{t_0} , related to the infinitesimal charge quantity 317 $d\xi(t_0, r_0) = \rho(t_0, r_0) d^3 r_0, \qquad \text{satisfying}$ 318 the charge conservation invariance $\int_{\Omega_t} d\xi(t,r) = \int_{\Omega_t} d\xi(t_0,r_0).$ The mentioned above infinitesimal Lorentz invariance relationships 319

320 make it possible to calculate the local energy conservation quantity $\mathrm{E}(\xi;\Omega_{_0})$ as

$$\mathbf{E}(\boldsymbol{\xi};\boldsymbol{\Omega}_{t}) = \int_{\boldsymbol{\Omega}_{t}} (\boldsymbol{\rho}\boldsymbol{\varphi}\boldsymbol{\chi}) d^{3}r = \int_{\boldsymbol{\Omega}_{t}} (\boldsymbol{\rho}\boldsymbol{\varphi}\frac{d^{3}r}{d^{3}r_{0}}) d^{3}r =$$

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$$= \int_{\Omega_t} (\rho \varphi \frac{d^3 r \wedge dt}{d^3 r_0 \wedge dt}) d^3 r = \int_{\Omega_t} (\rho \varphi \frac{d^3 r_0 \wedge dt_0}{d^3 r_0 \wedge dt}) d^3 r =$$
(28)

$$= \int_{\Omega_t} (\rho \varphi \frac{dt_0}{dt}) d^3 r = \int_{\Omega_t} \frac{\rho \varphi d^3 r}{(1 - |u|^2)^{1/2}}$$

where we took into account that $dt = dt_0 (1 - |u|^2)^{1/2}$. Thus, owing to (27) and (28) the local energy conservation relationship (22) is satisfied, proving the proposition.

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The constructed above local energy conservation quantity (28) can be rewritten as

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$$E(\xi;\Omega_t) = \int_{\Omega_t} \frac{d\xi(t,r)\varphi(t,r)}{(1-|u|^2)^{1/2}} = \int_{\Omega_{t_0}} d\xi(t_0,r_0)\varphi(t_0,r_0) := \int_{\Omega_{t_0}} dE(\xi;r_0) = E(\xi;\Omega_{t_0}), \quad (29)$$

where $dE(t_0, r_0) = d\xi(t_0, r_0)\varphi(t_0, r_0)$ is the distributed in vacuum electromagnetic field energy density, related with the electric charge $d\xi(t_0, r_0)$, located initially at point $(t_0, r_0) \in M^4$.

The above proposition suggests a physically motivated interpretation of electrodynamic phenomena in terms of what should naturally be called *the vacuum potential field*, which determines the observable interactions between charged point particles. More precisely, we can *a priori* endow the ambient vacuum medium with a scalar potential energy field density function $W := \xi \varphi : M^4 \to \mathbb{R}$, where $\xi \in \mathbb{R}_+$ is the value of an elementary charge quantity, and satisfying the governing *vacuum field equations*

$$\partial^2 W / \partial t^2 - \nabla^2 W = \rho \xi, \quad \partial W / \partial t + < \nabla, A >= 0,$$

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$$\partial^2 A / \partial t^2 - \nabla^2 A = \xi \rho v, \quad A = W v$$

taking into account the external charged sources, which possess a virtual capability for disturbing the vacuum field medium. Moreover, this vacuum potential field function $W: M^4 \rightarrow R$ allows the natural potential energy interpretation, whose origin should be assigned not only to the charged interacting medium, but also to any other medium possessing interaction capabilities, including for instance, material particles, interacting through the gravity.

The latter leads naturally to the next important step, consisting in deriving the equation governing the corresponding potential field $\overline{W}: M^4 \to R$, assigned to a charged point particle moving in the vacuum field medium with velocity $u \in T(\mathbb{R}^3)$ and located at point $r(t) := R(t) \in \mathbb{E}^3$ at time $t \in \mathbb{R}$. As can be readily shown [18] [19] [50] [69], the corresponding evolution equation governing the related potential field function $\overline{W}: M^4 \to \mathbb{R}$, assigned to a moving in the space \mathbb{E}^3 charged particle ξ under the stationary distributed field sources, has the form

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 $\frac{d}{dt}(-\overline{W}u) = -\nabla\overline{W},\tag{31}$

350 where $\overline{W} := W(t,r)|_{r \to R(t)}$, u(t) := dR(t) / dt at point particle location $(t, R(t)) \in M^4$.

Similarly, if there are two interacting charged point particles, located at points r(t) = R(t) and $r_f(t) = R_f(t) \in E^3$ at time $t \in R$ and moving, respectively, with velocities u := dR(t)/dt and $u_f := dR_f(t)/dt$, the corresponding potential field function $\overline{W}' : M^4 \to R$, considered with respect to the reference frame K_t' specified by Euclidean coordinates $(t', r - r_f) \in E^4$ and moving with the velocity $u_f \in T(R^3)$ subject to the laboratory reference frame K_t , should satisfy [18] [19] with respect to the reference frame K_t' the dynamical equality

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$$\frac{d}{dt}\left[-\overline{W}\left(u^{'}-u_{f}^{'}\right)\right]=-\nabla\overline{W}^{'},$$
(32)

where, by definition, we have denoted the velocity vectors u' := dr / dt', $u'_f := dr_f / dt' \in T(\mathbb{R}^3)$. The latter comes with respect to the laboratory reference frame K_r about the dynamical equality

362

$$\frac{d}{dt}\left[-\overline{W}\left(u-u_{f}\right)\right] = -\nabla\overline{W}\left(1-|u_{f}|^{2}\right).$$
(33)

The dynamical potential field equations (31) and (32) appear to have important properties and can be used as means for representing classical electrodynamic phenomena.

(30)

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Consequently, we shall proceed to investigate their physical properties in more detail and compare them with classical results for Lorentz type forces arising in the electrodynamics of moving charged point particles in an external electromagnetic field.

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1.2.1. Classical relativistic electrodynamics revisited

The classical relativistic electrodynamics of a freely moving charged point particle in the Minkowski space-time M^4 ; R×E³ is based on the Lagrangian approach [1] [5] [65] [66] [70] with Lagrangian function

374

$$L_0 := -m_0 (1 - |u|^2)^{1/2}, \tag{34}$$

where $m_0 \in \mathbb{R}_+$ is the so-called particle rest mass parameter with respect to the so called proper reference frame K_τ , parameterized by means of the Euclidean space-time parameters $(\tau, r) \in \mathbb{E}^4$, and $u \in T(\mathbb{R}^3)$ is its spatial velocity with respect to a laboratory reference frame K_τ , parameterized by means of the Minkowski space-time parameters $(t, r) \in M^4$, expressed here and in the sequel in light speed units (with light speed c = 1). The least action principle in the form

$$\delta S = 0, S := -m_0 \int_{t_1}^{t_2} (1 - |u|^2)^{1/2} dt$$
(35)

for any fixed temporal interval $[t_1, t_2] \subset \mathbb{R}$ gives rise to the well-known relativistic relationships

383 for the mass of the particle

$$m = m_0 (1 - |u|^2)^{-1/2},$$
(36)

385 the momentum of the particle

386 387

384

$$p := mu = m_0 u (1 - |u|^2)^{-1/2}$$
(37)

388 and the energy of the particle

389

$$\mathbf{E}_0 = m = m_0 (1 - |u|^2)^{-1/2}.$$
 (38)

390It follows from [5] [65], that the origin of the Lagrangian (34) can be extracted from the391action

392 $S := -m_0 \int_{t_1}^{t_2} (1 - |u|^2)^{1/2} dt = -m_0 \int_{\tau_1}^{\tau_2} d\tau,$ (39)

393 on the suitable temporal interval $[\tau_1, \tau_2] \subset \mathbb{R}$, where, by definition,

394

$$d\tau := dt (1 - |u|^2)^{1/2}$$
(40)

and $\tau \in \mathbb{R}$ is the so-called, proper temporal parameter assigned to a freely moving particle with respect to the proper reference frame K_{τ} . The action (39) is rather questionable from the dynamical point of view, since it is physically defined with respect to the proper reference frame K_{τ} , giving rise to the constant action $S = -m_0(\tau_2 - \tau_1)$, as the limits of integrations $\tau_1 < \tau_2 \in \mathbb{R}$ were taken to be fixed from the very beginning. Moreover, considering this particle to have charge $\xi \in \mathbb{R}$ and be moving in the Minkowski space-time M^4 under action of an electromagnetic field $(\varphi, A) \in T^*(M^4)$, the corresponding classical (relativistic) action 402 functional is chosen (see [1] [5] [47] [51] [65] [66]) as follows:

403
$$S := \int_{\tau_1}^{\tau_2} [-m_0 d\tau + \xi < A, \dot{r} > d\tau - \xi \varphi (1 - |u|^2)^{-1/2} d\tau],$$
(41)

404 with respect to the *proper reference frame*, parameterized by the Euclidean space-time 405 variables $(\tau, r) \in E^4$, where we have denoted $\dot{r} := dr / d\tau$ in contrast to the definition 406 u := dr / dt. The action (41) can be rewritten with respect to the laboratory reference frame K_i 407 as

408
$$S = \int_{t_1}^{t_2} L(r, dr / dt) dt, L(r, dr / dt) := -m_0 (1 - |u|^2)^{1/2} + \xi < A, u > -\xi \varphi,$$
(42)

409 on the suitable temporal interval $t_1, t_2] \subset \mathbb{R}$, which gives rise to the following [1] [5] [65] [66] 410 dynamical expressions

 $P = p + \xi A, \quad p = mu, \quad m = m_0 (1 - |u|^2)^{-1/2}, \tag{43}$

412 for the particle momentum and

$$\mathbf{E}_{0} = (m_{0}^{2} + |P - \boldsymbol{\xi}A|^{2})^{1/2} + \boldsymbol{\xi}\boldsymbol{\varphi}$$
(44)

for the charged particle ξ energy, where, by definition, $P \in E^3$ is the common momentum of the particle and the ambient electromagnetic field at a Minkowski space-time point $(t, r) \in M^4$.

416 The related dynamics of the charged particle ξ follows [1] [5] [65] [66] from the Lagrangian 417 equation

418 419

411

413

$$dP / dt := \nabla L(r, dr / dt) = -\nabla (\xi \varphi - \xi < A, u >).$$
(45)

420 The expression (44) for the particle energy E_0 also appears to be open to question, since the potential energy $\xi \varphi$, entering additively, has no affect on the particle "inertial" mass 421 $m = m_0 (1 - |u|^2)^{-1/2}$. This was noticed by L. Brillouin [21], who remarked that the fact that the 422 potential energy has no affect on the particle mass tells us that "... any possibility of existence 423 424 of a particle mass related with an external potential energy, is completely excluded". Moreover, 425 it is necessary to stress here that the least action principle (42), formulated with respect to the 426 laboratory reference frame K, time parameter $t \in \mathbb{R}$, appears logically inadequate, for there is a strong physical inconsistency with other time parameters of the Lorentz equivalent reference 427 428 frames. This was first mentioned by R. Feynman in [1] in his efforts to physically argue the 429 Lorentz force expression with respect to the proper reference frame K_r . This and other special 430 relativity theory and electrodynamics problems stimulated many prominent physicists of the past [1] [21] [65] [71] [72] and present [7] [23] [24] [25] [26] [44] [57] [59] [60] [73] [74] 431 432 [75] [76] [77] [78] and [79] [80] [81] [11] [82] [69] [83] [84] [85] [86] [87] to try to develop 433 alternative relativity theories based on completely different space-time and matter structure 434 principles. Some of them prove to be closely related with a virtual relationship between 435 electrodynamics and gravity, based on classical works of H. Lorentz, G. Schott, J. Schwinger, R. 436 Feynman [1] [22] [53] [54] [63] [88] and many others on the so called "electrodynamic mass" of elementary particles. Arguing this way of this mass, one can readily come to a certain paradox: 437 438 the well-known energy-mass relationship for the particle mass suitably determines the energy 439 of its gravitational field. Yet this energy should lead to an increase in the mass of the particle

that in turn should lead to increased gravitational field and so on. In the limit, for instance, an
electron must have infinite mass and energy, what we do not really observe. There also is
another controversial inference from the action expression (42). As one can easily show, owing
to (45), the corresponding expression for the Lorentz force

444

446

448

$$dp / dt = F_t := \xi E + \xi u \times B \tag{46}$$

445 holds, where we have defined here, as before,

$$E := -\partial A / \partial t - \nabla \varphi \tag{47}$$

447 the corresponding electric field and

$$B := \nabla \times A \tag{48}$$

the related magnetic field, acting on the charged point particle ξ . The expression (46), in 449 particular, means that the Lorentz force F_L depends linearly on the particle velocity vector 450 $u \in T(\mathbb{R}^3)$, and so there is a strong dependence on the reference frame with respect to which 451 452 the charged particle ξ moves. Attempts to reconcile this and some related controversies [21] 453 [1] [89] [11] [69] [13] forced Einstein to devise his special relativity theory and proceed further 454 to creating his general relativity theory trying to explain the gravity by means of geometrization 455 of space-time and matter in the Universe. Here we must mention that the classical Lagrangian 456 function L in (42) is written in terms of a combination of terms expressed by means of both the Euclidean proper reference frame variables $(\tau, r) \in E^4$ and arbitrarily chosen Minkowski 457 reference frame variables $(t, r) \in M^4$. 458

459 These problems were recently analyzed using a completely different " no-geometry" 460 approach [18] [19] [69], where new dynamical equations were derived, which were free of the 461 controversial elements mentioned above. Moreover, this approach avoided the introduction of 462 the well known Lorentz transformations of the space-time reference frames with respect to 463 which the action functional (42) is invariant. From this point of view, there are interesting for 464 discussion conclusions from [90] [91] [92] [93], in which some electrodynamic models, 465 possessing intrinsic Galilean and Poincaré-Lorentz symmetries, were reanalyzed from diverse 466 geometrical points of view. From completely different point of view the related 467 electrodynamics of charged particles was reanalyzed in [3] [4] [8] [14] [15], where all relativistic 468 relationships were successfully inferred from the classical Lienard-Wiechert potentials, solving 469 the corresponding electromagnetic equations. Subject to a possible geometric space-type 470 structure and the related vacuum field background, exerting the decisive influence on the 471 particle dynamics, we need to mention here recent works [79] [85] [13] and the closely related 472 with their ideas the classical articles [94] [95]. Next, we shall revisit the results obtained in [18] 473 [19] from the classical Lagrangian and Hamiltonian formalisms [47] [64] [66] [96] in order to 474 shed new light on the physical underpinnings of the vacuum field theory approach to the study 475 of combined electromagnetic and gravitational effects.

476

477 1.3. The vacuum field theory electrodynamics equations: Lagrangian
 478 analysis

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- 480
- 481

1.3.2. A moving in vacuum point charged particle - an alternative electrodynamic

482 **model**

In the vacuum field theory approach to combining electromagnetism and the gravity, devised in [18] [19], the main vacuum potential field function $\overline{W}: M^4 \to R$, related to a charged point particle ξ under the external stationary distributed field sources, satisfies the dynamical equation (30), namely

488

483

$$\frac{d}{dt}(-\overline{W}u) = -\nabla \overline{W} \tag{49}$$

in the case when the external charged particles are at rest, where, as above, u := dr / dt is the particle velocity with respect to some reference system.

491 To analyze the dynamical equation (49) from the Lagrangian point of view, we write the 492 corresponding action functional as

493
$$S := -\int_{t_1}^{t_2} \overline{W} dt = -\int_{\tau_1}^{\tau_2} \overline{W} (1+|\dot{r}|^2)^{1/2} d\tau,$$
(50)

494 expressed with respect to the proper reference frame K_{τ} . Fixing the proper temporal 495 parameters $\tau_1 < \tau_2 \in \mathbb{R}$, one finds from the least action principle ($\delta S = 0$) that

$$:= \partial L / \partial \dot{r} = -\overline{W}\dot{r}(1+|\dot{r}|^2)^{-1/2} = -\overline{W}u,$$
(51)

496

500

502

505

508

$$\dot{p} := dp / d\tau = \partial L / \partial r = -\nabla \overline{W} (1 + |\dot{r}|^2)^{1/2},$$

 $L := -\overline{W}(1+|\dot{r}|^2)^{1/2}.$

497 where, owing to (50), the corresponding Lagrangian function is

р

498

499 Recalling now the definition of the particle mass

 $m := -\overline{W} \tag{53}$

501 and the relationships

$$d\tau = dt(1 - |u|^2)^{1/2}, \dot{r}d\tau = udt,$$
(54)

503 from (51) we easily obtain exactly the dynamical equation (49). Moreover, one now readily 504 finds that the dynamical mass, defined by means of expression (53), is given as

$$m = m_0 (1 - |u|^2)^{-1/2}$$

which coincides with the equation (36) of the preceding section. Now one can formulate thefollowing proposition using the above results.

509 **Proposition 3.** The alternative freely moving point particle electrodynamic model (49) allows the 510 least action formulation (50) with respect to the "rest" reference frame variables, where the 511 Lagrangian function is given by expression (52). Its electrodynamics is completely equivalent to 512 that of a classical relativistic freely moving point particle, described in Subsection 1.2.1.

513

516

5141.3.3. A moving in vacuum interacting two charge system - an alternative515electrodynamic model

517 We proceed now to the case when our charged point particle ξ moves in the space-

(52)

time with velocity vector $u \in T(\mathbb{R}^3)$ and interacts with another external charged point particle ξ_f , moving with velocity vector $u_f \in T(\mathbb{R}^3)$ with respect to a common reference frame K_t . As was shown in [18] [19], the respectively modified dynamical equation for the vacuum potential field function $\overline{W}': M^{4,'} \to \mathbb{R}$ subject to the moving reference frame K_t' is given by equality (32), or

$$\frac{d}{dt}\left[-\overline{W}'(u'-u'_f)\right] = -\nabla\overline{W}',$$
(55)

where, as before, the velocity vectors u' := dr / dt', $u'_f := dr_f / dt' \in T(\mathbb{R}^3)$. Since the external charged particle ξ_f moves in the space-time M^4 , it generates the related magnetic field $B := \nabla \times A$, whose magnetic vector potentials $A : M^4 \to \mathbb{E}^3$ and $A' : M^{4,'} \to \mathbb{E}^3$ are defined, owing to the results of [18] [19] [69], as

$$\boldsymbol{\xi}\boldsymbol{A} := \boldsymbol{W}\boldsymbol{u}_{f}, \, \boldsymbol{\xi}\boldsymbol{A}' := \boldsymbol{W}'\boldsymbol{u}_{f}', \tag{56}$$

529 Whence, taking into account that the field potential

530

535

528

$$\overline{W} = \overline{W}' (1 - |u_f|^2)^{-1/2}$$
(57)

and the particle momentum $p' = -\overline{W}u' = -\overline{W}u$, equality (55) becomes equivalent to

532
$$\frac{d}{dt}(p'+\xi A') = -\nabla \overline{W}', \qquad (58)$$

533 if considered with respect to the moving reference frame K_{t}' , or to the Lorentz type force 534 equality

$$\frac{d}{dt}(p+\xi A) = -\nabla \overline{W}(1-|u_f|^2), \tag{59}$$

if considered with respect to the laboratory reference frame K_t , owing to the classical Lorentz invariance relationship (57), as the corresponding magnetic vector potential, generated by the external charged point test particle ξ_f with respect to the reference frame K_t , is identically equal to zero. To imbed the dynamical equation (59) into the classical Lagrangian formalism, we start from the following action functional, which naturally generalizes the functional (50):

541
$$S := -\int_{\tau_1}^{\tau_2} \overline{W}' (1 + |\dot{r} - \dot{r}_f|^2)^{1/2} d\tau.$$
(60)

Here, as before, \overline{W}' is the respectively calculated vacuum field potential \overline{W} subject to the moving reference frame K_{t} , $\dot{r} = u'dt'/d\tau$, $\dot{r}_{f} = u'_{f}dt'/d\tau$, $d\tau = dt'(1 - |u' - u'_{f}|^{2})^{1/2}$, which take into account the relative velocity of the charged point particle ξ subject to the reference frame K_{t} , specified by the Euclidean coordinates $(t', r - r_{f}) \in \mathbb{R}^{4}$, and moving simultaneously with velocity vector $u_{f} \in T(\mathbb{R}^{3})$ with respect to the laboratory reference frame K_{t} , specified by the Minkowski coordinates $(t, r) \in M^{4}$ and related to those of the reference frame K_{t} and K_{τ} by means of the following infinitesimal relationships: 557

549
$$dt^{2} = (dt')^{2} + |dr_{f}|^{2}, (dt')^{2} = d\tau^{2} + |dr - dr_{f}|^{2}.$$
(61)

550 So, it is clear in this case that our charged point particle ξ moves with the velocity vector 551 $u' - u'_f \in T(\mathbb{R}^3)$ with respect to the reference frame K_t in which the external charged particle 552 ξ_f is at rest. Thereby, we have reduced the problem of deriving the charged point particle ξ 553 dynamical equation to that before solved in Subsection 1.2.1.

Now we can compute the least action variational condition $\delta S = 0$, taking into account that, owing to (60), the corresponding Lagrangian function with respect to the proper reference frame K_{τ} is given as

- $L := -\overline{W}' (1 + |\dot{r} \dot{r}_f|^2)^{1/2}.$ (62)
- 558 As a result of simple calculations, the generalized momentum of the charged particle ξ equals $P := \partial L / \partial \dot{r} = -\overline{W} (\dot{r} - \dot{r}_{t})(1 + |\dot{r} - \dot{r}_{t}|^{2})^{-1/2} =$

559 $= -\overline{W}'\dot{r}(1+|\dot{r}-\dot{r}_{f}|^{2})^{-1/2} + \overline{W}'\dot{r}_{f}(1+|\dot{r}-\dot{r}_{f}|^{2})^{-1/2} =$

$$= m'u' + \xi A' := p' + \xi A' = p + \xi A,$$

560 where, owing to (57) the vectors $p' := -\overline{W}u' = -\overline{W}u = p \in E^3$, $A' = \overline{W}u_f = A \in E^3$, and 561 giving rise to the dynamical equality

562
$$\frac{d}{d\tau}(p' + \xi A') = -\nabla \overline{W}' (1 + |\dot{r} - \dot{r}_f|^2)^{1/2}$$
(64)

563 with respect to the proper reference frame K_{τ} . As $dt' = d\tau (1+|\dot{r}-\dot{r}_f|^2)^{1/2}$ and 564 $(1+|\dot{r}-\dot{r}_f|^2)^{1/2} = (1-|u'-u'_f|^2)^{-1/2}$, we obtain from (64) the equality

565 $\frac{d}{dt}(p'+\xi A') = -\nabla \overline{W}', \qquad (65)$

exactly coinciding with equality (58) subject to the moving reference frame K_{t} . Now, making use of expressions (61) and (57), one can rewrite (65) as that with respect to the laboratory reference frame K_{t} :

(63)

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$$\frac{d}{dt}(p'+\xi A') = -\nabla \overline{W}' \Rightarrow$$

$$\Rightarrow \frac{d}{dt}(\frac{-\overline{Wu'}}{(1+|u'_{f}|^{2})^{1/2}} + \frac{\xi \overline{Wu'_{f}}}{(1+|u'_{f}|^{2})^{1/2}}) = -\frac{\nabla \overline{W}}{(1+|u'_{f}|^{2})^{1/2}} \Rightarrow$$

$$\Rightarrow \frac{d}{dt}(\frac{-\overline{Wdr}}{(1+|u'_{f}|^{2})^{1/2}dt} + \frac{\xi \overline{W} \overline{d}r_{f}}{(1+|u'_{f}|^{2})^{1/2}dt}) = -\frac{\nabla \overline{W}}{(1+|u'_{f}|^{2})^{1/2}} \Rightarrow$$

$$\Rightarrow \frac{d}{dt}(-\overline{W}\frac{dr}{dt} + \xi \overline{W}\frac{dr_{f}}{dt}) = -\nabla \overline{W}(1-|u_{f}|^{2}),$$
exactly coinciding with (59):
$$\frac{d}{dt}(p+\xi A) = -\nabla \overline{W}(1-|u_{f}|^{2}).$$
(67)
Remark 1. The equation (67) allows to infer the following important and physically reasonable phenomenon: if the test charged point particle velocity $u_{f} \in T(\mathbb{R}^{3})$ tends to the light velocity $c = 1$, the corresponding acceleration force $F_{ac} := -\nabla \overline{W}(1-|u_{f}|^{2})$ is vanishing. Thereby, the electromagnetic fields, generated by such rapidly moving charged point particles, have no influence on the dynamics of charged objects if observed with respect to an arbitrarily chosen

laboratory reference frame K_r.

c = 1,

The latter equation (67) can be easily rewritten as

$$dp / dt = -\nabla \overline{W} - \xi dA / dt + \nabla \overline{W} |u_f|^2 =$$

$$= \quad \xi(-\xi^{-1}\nabla \overline{W} - \partial A / \partial t) - \xi < u, \nabla > A + \xi \nabla < A, u_f >,$$

583or, using the well-known Error! Reference source not found. identity584
$$\nabla < a, b > = < a, \nabla > b + < b, \nabla > a + b \times (\nabla \times a) + a \times (\nabla \times b),$$
(69)585where $a, b \in E^3$ are arbitrary vector functions, in the standard Lorentz type form $dp / dt = \xi E + \xi u \times B - \nabla < \xi A, u - u_f > .$ (70)587The result (70), being before found and written down with respect to the movingreference frame $K_i^{'}$ in [18] [19] [69] makes it possible to formulate the next important590proposition.

Proposition 4. The alternative classical relativistic electrodynamic model (58) allows the least action formulation based on the action functional (60) with respect to the proper reference

(68)

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frame K_{τ} , where the Lagrangian function is given by expression (62). The resulting Lorentz type force expression equals (70), being modified by the additional force component $F_c := -\nabla < \xi A, u - u_f >$, important for explanation [97] [98] [99] of the well known Aharonov-Bohm effect.

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- 599 600

1.3.4. A moving charged point particle dynamics formulation dual to the classical relativistic invariant alternative electrodynamic model

601

602 It is easy to see that the action functional (60) is written utilizing the classical Galilean 603 transformations of reference frames. If we now consider the action functional (50) for a 604 charged point particle moving with respect the reference frame K_{τ} , and take into account its 605 interaction with an external magnetic field generated by the vector potential $A: M^4 \rightarrow E^3$, it 606 can be naturally generalized as

607
$$S := \int_{\tau_1}^{\tau_2} (-\overline{W}dt + \xi < A, dr >) = \int_{\tau_1}^{\tau_2} [-\overline{W}(1 + |\dot{r}|^2)^{1/2} + \xi < A, \dot{r} >] d\tau,$$
(71)

608 where $d\tau = dt(1-|u|^2)^{1/2}$.

609 Thus, the corresponding common particle-field momentum takes the form $P := \partial L / \partial \dot{r} = -\overline{W}\dot{r}(1+|\dot{r}|^2)^{-1/2} + \xi A =$

610

 $= mu + \xi A := p + \xi A,$

611 and satisfies

$$\dot{P} := dP / d\tau = \partial L / \partial r = -\nabla \overline{W} (1 + |\dot{r}|^2)^{1/2} + \xi \nabla < A, \dot{r} >=$$
(73)

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618 619

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626

 $= -\nabla \overline{W} (1 - |u|^2)^{-1/2} + \xi \nabla < A, u > (1 - |u|^2)^{-1/2},$

613 where

 $L := -\overline{W}(1+|\dot{r}|^2)^{1/2} + \xi < A, \dot{r} >$ (74)

615 is the corresponding Lagrangian function. Since $d\tau = dt(1-|u|^2)^{1/2}$, one easily finds from (73) 616 that 617 $dP/dt = -\nabla \overline{W} + \xi \nabla < A, u >$. (75)

Upon substituting (72) into (75) and making use of the identity (69), we obtain the classical expression for the Lorentz force
$$F$$
, acting on the moving charged point particle ξ :
 $dp / dt := F_L = \xi E + \xi u \times B$, (76)

621 where, by definition,

$E := -\xi^{-1} \nabla \overline{W} - \partial A / \partial t$	(77)
--	------

623 is its associated electric field and

 $B := \nabla \times A \tag{78}$

625 is the corresponding magnetic field. This result can be summarized as follows.

627 **Proposition 5.** The classical relativistic Lorentz force (76) allows the least action formulation (71)

(72)

628 with respect to the proper reference frame variables, where the Lagrangian function is given by 629 formula (74). Yet its electrodynamics, described by the Lorentz force (76), is not equivalent to 630 the classical relativistic moving point particle electrodynamics, described by means of the 631 Lorentz force (46), as the inertial mass expression $m = -\overline{W}$ does not coincide with that of (36).

632

Expressions (76) and (70) are equal up to the gradient like term $F_c := -\nabla < \xi A, u - u_f >$, which reconciles the Lorentz forces acting on a charged moving particle ξ with respect to different reference frames. This fact is important for our vacuum field theory approach since it uses no special geometry and makes it possible to analyze both electromagnetic and gravitational fields simultaneously by employing the new definition of the dynamical mass by means of the Mach-Einstein type expression (53).

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- 640 641

1.4. The A.M. Ampere's law in electrodynamics - the classical and modified Lorentz force derivations

642

The classical ingenious Andre-Marie Ampere's analysis of magnetically interacting to each other two electric currents in thin conductors, as is well known, was based [1] [5] [65] [66] on the following experimental fact: the force between two electric currents depends on the distance between conductors, their mutual spatial orientation and the quantitative values of currents. Having additionally accepted the infinitesimal superposition principle of A.M. Ampere had derived a general analytical expression for the force between two infinitesimal elements of currents under regard:

- 650
- 651

$$df(r,r') = II' \frac{(r-r')}{|r-r'|^2} \alpha(s,s';n) dl dl',$$
(79)

where vectors $r, r' \in E^3$ point at infinitesimal currents dr = sdl, dr' = s'dl' with normalized orientation vectors $s, s' \in E^3$ of two closed conductors l and l' carrying currents $I \in \mathbb{R}$ and $I' \in \mathbb{R}$, respectively and the unit vector n := (r - r')/|r - r'|, fixing the spatial orientations of these infinitesimal elements, and the function $\alpha : (S^2)^2 \times S^2 \to \mathbb{R}$ being some real-valued smooth mapping. Taking further into account the mutual symmetry between the infinitesimal elements of currents dl and dl', belonging respectively to these two electric conductors, the infinitesimal force (79) was assumed by A.M. Ampere to satisfy locally the third Newton's law:

660

$$df(r,r') = -df(r',r)$$
 (80)

661 with the mapping

662

$$\alpha(s, s'; n) = \frac{\mu_0}{4\pi} (3k_1 < s, n > < s', n > +k_2 < s, s' >),.$$
(81)

664 where $\langle \cdot, \cdot \rangle$ is the natural scalar product in E^3 and $k_1, k_2 \in \mathbb{R}$ are some still undetermined 665 real and dimensionless parameters. The assumption (80) is evidently looking very restrictive 666 and can be considered as reasonable only subject to a stationary system of conductors under

667 regard, when the mutual action at a distance principle [1] [5] can be applied. Owing to himself J.C. Maxwell [67]: "... we may draw the conclusions, first, that action and reaction are not 668 669 always equal and opposite, and second, that apparatus may be constructed to generate any 670 amount of work from its own resources. For let two oppositely electrified bodies A and B671 travel along the line joining them with equal velocities in the direction AB, then if either the 672 potential or the attraction of the bodies at a given time is that due to their position at some 673 former time (as these authors suppose), B, the foremost body, will attract A forwards more 674 than B attracts A backwards. Now let A and B be kept asunder by a rigid rod. The combined system, if set in motion in the direction AB, will pull in that direction with a force which may 675 676 either continually augment the velocity, or may be used as an inexhaustible source of energy."

Based on the fact that there is no possibility to measure the force between two infinitesimal current elements, A.M. Ampere took into account (80), (81) and calculated the corresponding force exerted by the whole conductor l' on an infinitesimal current element of other conductor under regard:

681

$$dF(r) := \int_{l'} df(r, r') =$$

682

$$=\frac{H'\mu_{0}}{4\pi}\int_{l}\frac{(r-r')}{|r-r'|^{2}}(3k_{1} < dr, \frac{r-r'}{|r-r'|} > < dr', \frac{r-r'}{|r-r'|} > +k_{2}\frac{r-r'}{|r-r'|} < dr, dr' >) =$$
(82)

$$=\frac{H'\mu_{0}}{4\pi}\int_{l}\nabla_{r'}\left(\frac{1}{|r-r'|}\right)(3k_{1} < dr, r-r' >< dr', r-r' >+k_{2} < dr, dr' >),$$

683 which can be equivalently transformed as684

$$dF(r) = \frac{H'\mu_0}{4\pi} \int_{l} \nabla_{r'} \left(\frac{1}{|r-r'|}\right) (3k_1 < dr, r-r' > < dr', r-r' > +k_2 < dr, dr' >) =$$

$$=\frac{H'\mu_{0}}{4\pi}\int_{l}\nabla_{r'}\left(\frac{1}{|r-r'|}\right)[k_{1}(3 < dr, r-r' > < dr', r-r' > - (83)]$$

685

$$- < dr, dr' >) + (k_1 + k_2) < dr, dr' >] =$$

$$= -k_1 \frac{\mu_0 I}{4\pi} < dr, \nabla \int_{I'} \left(\frac{I' dr'}{|r - r'|} \right) > -(k_1 + k_2) < \nabla, \int_{I'} < dr, \frac{I' dr'}{|r - r'|} >,$$

686 owing to the integral identity

687

688
$$\int_{l} \nabla_{r'} \left(\frac{1}{|r-r'|} \right) (3 < dr, r-r' > < dr', r-r' > - < dr, dr' >) = < dr, \nabla > \int_{l'} \frac{dr'}{|r-r'|}, \quad (84)$$

689 which can be easily checked by means of integration by parts. If to introduce the vector 690 potential

691

692

$$A(r) := \frac{\mu_0 I'}{4\pi} \int_{I'} \frac{dr'}{|r-r'|},$$

693 generated by the conductor l' at point $r \in E^3$, belonging to the infinitesimal element dl of the 694 conductor l, the resulting infinitesimal force (83) gives rise to the following expression: 695

$$dF(r) = k_1(-I < dr, \nabla)A(r) + I\nabla < dr, A(r) >) - (2k_1 + k_2)I\nabla < dr, A(r) >=$$

696 $= k_1 I dr \times (\nabla \times A(r)) - (2k_1 + k_2) I \nabla < dr, A(r) >=$

$$= k_1 J(r) d^3 r \times B(r) - (2k_1 + k_2) \nabla < J d^3 r, A(r) >,$$

697 where we have taken into account the standard magnetic field definition 698 699 (87) $B(r) := \nabla \times A(r)$ 700 and the corresponding current density relationship 701 $J(r)d^3r := Idr.$ 702 (88) 703 There are, evidently, many different possibilities to choose the dimensionless parameters $k_1, k_2 \in \mathbb{R}$. In his analysis A.M. Ampere had chosen the case when $k_1 = 1, k_2 = -2$ 704 705 and obtained the well known nowadays magnetic force expression 706 707 $dF(r) = J(r)d^3r \times B(r),$ (89)

708 which easily reduces to the *classical Lorentz expression*

709 710

 $df_L(r) = \xi u \times B(r) \tag{90}$

for a force exerted by an external magnetic field on a moving with a velocity $u \in T(\mathbb{R}^3)$ point particle with an electric charge $\xi \in \mathbb{R}$.

If to take an alternative choice and put $k_1 = 1, k_2 = -1$, the expression (86) yields *a* modified magnetic Lorentz type force, exerted by an external magnetic field generated by a moving charged particle with a velocity $u \in T(\mathbb{R}^3)$ on a point particle, endowed with the electric charge $\xi \in \mathbb{R}$ and moving with a velocity $u \in T(\mathbb{R}^3)$:

717 718

$$dF_L(r) = J(r)d^3r \times B(r) - \nabla < J(r)d^3r, A(r) >,$$
(91)

which was before occasionally discussed in different works [9] [10] [11] [69] [100] and recently

(85)

(86)

enough strongly obtained and analyzed in detail from the Lagrangian point of view in works[18] [19] [50] [51] in the following equivalent to (70) infinitesimal form:

722 723

$$\delta f_L(r) = \xi u \times (\nabla \times \xi \delta A(r)) - \xi \nabla < u - u_f, \delta A(r) >, \tag{92}$$

724 where $\delta A(r) \in T^*(\mathbb{R}^3)$ denotes the magnetic potential generated by an external charged point particle moving with velocity $u_f \in T(\mathbb{R}^3)$ and exerting the magnetic force $\delta f_I(r)$ on the 725 charged particle located at point $r \in \mathbb{R}^3$ and moving with velocity $u \in T(\mathbb{R}^3)$ with respect to a 726 727 common reference system K. We also need to mention here that the modified Lorentz force 728 expression (91) does not take naturally into account the resulting pure every weak electric 729 force, as the conductors l and l' are considered to be electrically neutral. Simultaneously, we 730 see that the magnetic potential has a physical significance in its own right [6] [9] [11] [50] [69] and has meaning in a way that extends beyond the calculation of force fields. 731

Really, to obtain the Lorentz type force (91) exerted by the external magnetic field generated by *the whole conductor* l' on an infinitesimal current element dl of the conductor l, it is necessary to integrate the expression (92) along this conductor loop l':

735

$$dF_{L}(r) := \int_{i}^{I} \delta f_{L}(r) = J(r)dr \times (\nabla \times \int_{i}^{I} \delta A(r)) - \nabla < J(r)dr, \int_{i}^{I} \delta A(r) > +$$

$$+ \nabla \int_{i}^{I} < u', \xi \delta A(r) >= J(r)dr \times (\nabla \times A(r)) - \nabla < J(r)dr, \int_{i}^{I} \delta A(r) > +$$

$$+ \nabla \int_{i}^{I} < dr', \xi \delta A(r) / dt >= J(r)dr \times B(r) - \nabla < J(r)dr, \int_{i}^{I} \delta A(r) > +$$

$$+ \nabla \int_{S(i')}^{I} < dS(i'), \nabla \times \xi \delta A(r) / dt >= J(r)dr \times B(r) - \nabla < J(r)dr, \int_{i}^{I} \delta A(r) > +$$

$$+ \nabla \int_{S(i')}^{I} < dS(i'), \xi \delta B(r) / dt >= J(r)dr \times B(r) - \nabla < J(r)dr, \int_{i}^{I} \delta A(r) > +$$
(93)

736

738

$$+\xi \nabla (d\Phi(r)/dt) = J(r)dr \times B(r) - \nabla < J(r)dr, A(r) > -\rho(r)d^3r \nabla \overline{W} =$$
$$= J(r)dr \times B(r) - \nabla < J(r)dr, \int_{i} \delta A(r) > +\rho(r)d^3r(-\nabla \overline{W} - \partial A(r)/\partial t) =$$

$$= J(r)dr \times B(r) - \nabla < J(r)dr, \int_{l} \delta A(r) > +\rho(r)d^{3}rE(r),$$

737 that is the equality

$$dF(r) = \rho(r)d^3rE(r) + J(r)d^3r \times B(r) - \nabla < J(r)d^3r, A(r) >,$$
(94)

739 where, by definition, the electric field $E(r) := -\nabla \overline{W} - \partial A(r) / \partial t$. Now one can easily derive from 740 (94) the searched for *Lorentz type force* expression (91), if to take into account that the whole 741 electric field E(r); 0 owing to the neutrality of the conductors.

The presented above analysis of the A.M. Ampere's derivation of the magnetic force expression (86), as well as its consequences (91) and (92) make it possible to suppose that the missed modified Lorentz type force expression (91) could also be embedded into the classical relativistic Lagrangian and related Hamiltonian formalisms, giving rise to eventually new aspects and interpretations of many observed during the past centuries looking "strange" experimental phenomena.

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- 749 750

1.5. The vacuum field theory electrodynamics equations: Hamiltonian analysis

751

Any Lagrangian theory has an equivalent canonical Hamiltonian representation via the classical Legendre transformation [64] [66] [96] [101] [102]. As we have already formulated our vacuum field theory of a moving charged particle ξ in Lagrangian form, we proceed now to its Hamiltonian analysis making use of the action functionals (50), (62) and (71).

Take, first, the Lagrangian function (52) and the momentum expression (51) for defining the corresponding Hamiltonian function with respect to the moving reference frame K_{τ} :

$$H := < p, \dot{r} > -L =$$

758

$$= -\langle p, p \rangle \overline{W}^{-1} (1 - |p|^2 / \overline{W}^2)^{-1/2} + \overline{W} (1 - |p|^2 / \overline{W}^2)^{-1/2} =$$

$$= -|p|^2 \overline{W}^{-1} (1 - |p|^2 / \overline{W}^2)^{-1/2} + \overline{W}^2 \overline{W}^{-1} (1 - |p|^2 / \overline{W}^2)^{-1/2} =$$
(95)

$$= -(\overline{W}^2 - |p|^2)(\overline{W}^2 - |p|^2)^{-1/2} = -(\overline{W}^2 - |p|^2)^{1/2}.$$

Consequently, it is easy to show [64] [96] [102] [66] that the Hamiltonian function (95) is a conservation law of the dynamical field equation (49), that is for all $\tau, t \in \mathbb{R}$

761

764

 $dH / d\tau = dH / dt = 0,$ (96)

which naturally leads to an energy interpretation of *H*. Thus, we can represent the particleenergy as

 $\mathbf{E} = (\overline{W}^2 - |p|^2)^{1/2}.$ (97)

Accordingly the Hamiltonian equivalent to the vacuum field equation (49) can be writtenas

$$\dot{r} := dr / d\tau = \partial H / \partial p = p(\overline{W}^2 - |p|^2)^{-1/2}$$
(98)

767

$$\dot{p} := dp / d\tau = -\partial H / \partial r = \overline{W} \nabla \overline{W} (\overline{W}^2 - |p|^2)^{-1/2},$$

and we have the following result.

769

Proposition 6. The alternative freely moving point particle electrodynamic model (49) allows the
canonical Hamiltonian formulation (98) with respect to the "rest" reference frame variables,
where the Hamiltonian function is given by expression (95). Its electrodynamics is completely

773 equivalent to the classical relativistic freely moving point particle electrodynamics described in 774 Subsection 1.2.1.

775

776 In the analogous manner, one can now use the Lagrangian (62) to construct the Hamiltonian function for the dynamical field equation (58), describing the motion of charged 777 778 particle ξ in an external electromagnetic field in the canonical Hamiltonian form:

779
$$\dot{r} := dr / d\tau = \partial H / \partial P, \quad \dot{P} := dP / d\tau = -\partial H / \partial r, \tag{99}$$

 $= < P, \dot{r}_{f} > + |P|^{2} (\overline{W}^{',2} - |P|^{2})^{-1/2} - \overline{W}^{',2} (\overline{W}^{',2} - |P|^{2})^{-1/2} =$

 $= < P, \dot{r}_{f} - P \overline{W}^{',-1} (1 - |P|^{2} / \overline{W}^{',2})^{-1/2} > + \overline{W}^{'} [\overline{W}^{',2} (\overline{W}^{',2} - |P|^{2})^{-1}]^{1/2} =$

780 where

 $H := < P, \dot{r} > -L =$

 $= -(\overline{W}^{',2} - |P|^2)(\overline{W}^{',2} - |P|^2)^{-1/2} + \langle P, \dot{r}_f \rangle =$

$$= -(\overline{W}^{',2} - |P|^2)^{1/2} - \xi < A^{'}, P > (\overline{W}^{',2} - |P|^2)^{-1/2} =$$

$$= -(\overline{W}^{2} - |\xi A|^{2} - |P|^{2})^{1/2} - \xi < A, P > (\overline{W}^{2} - |\xi A|^{2} - |P|^{2})^{-1/2}$$

being written with respect to the laboratory reference frame K_{i} . Here we took into account 782

$$\xi A' := \overline{W}' u'_f = \overline{W}' dr_f / dt' = \xi A =$$

 $= \overline{W}' \frac{dr_f}{d\tau} \cdot \frac{d\tau}{dt'} = \overline{W}' \dot{r}_f (1 - |u - u_f|)^{1/2} =$ (101) $= \overline{W}' \dot{r}_{f} (1 + |\dot{r} - \dot{r}_{f}|^{2})^{-1/2} =$

784

$$= -\overline{W}'\dot{r}_{f}(\overline{W}'^{,2} - |P|^{2})^{1/2}\overline{W}'^{,-1} = -\dot{r}_{f}(\overline{W}'^{,2} - |P|^{2})^{1/2},$$

785 and, in particular,

790

$$\dot{r}_{f} = -\xi A(\overline{W}^{',2} - |P|^{2})^{-1/2}, \overline{W} = \overline{W}^{'}(1 - |u_{f}|^{2})^{-1/2},$$
(102)

where $A: M^4 \rightarrow \mathbb{R}^3$ is the related magnetic vector potential generated by the moving external 787 charged particle ξ_{f} . Equations (99) can be rewritten with respect to the laboratory reference 788 789 frame K_r in the form

$$dr / dt = u, dp / dt = \xi E + \xi u \times B - \xi \nabla < A, u - u_f >,$$
(103)

791 which coincides with the result (70).

24

(100)

792 Whence, we see that the Hamiltonian function (100) satisfies the energy conservation 793 conditions

$$dH / d\tau = dH / dt' = dH / dt = 0,$$
 (104)

795 for all τ, t' and $t \in \mathbb{R}$, and that the suitable energy expression is

796

794

 $\mathbf{E} = (\overline{W}^2 - \xi^2 |A|^2 - |P|^2)^{1/2} + \xi < A, P > (\overline{W}^2 - \xi^2 |A|^2 - |P|^2)^{-1/2},$

(105)797 where the generalized momentum $P = p + \xi A$. The result (105) differs essentially from that 798 obtained in [5], which makes use of the Einstein's Lagrangian for a moving charged point 799 particle ξ in an external electromagnetic field. Thus, we obtain the following proposition.

800

Proposition 7. The alternative classical relativistic electrodynamic model (103), which is 801 intrinsically compatible with the classical Maxwell equations (6), allows the Hamiltonian 802 803 formulation (99) with respect to the proper reference frame variables, where the Hamiltonian 804 function is given by expression (100).

806 The inference above is a natural candidate for experimental validation of our theory. It 807 is strongly motivated by the following remark.

808

805

809 **Remark 2.** It is necessary to mention here that the Lorentz force expression (103) uses the particle momentum p = mu, where the dynamical "mass" $m := -\overline{W}$ satisfies condition (105). 810 The latter gives rise to the following crucial relationship between the particle energy $E_{\!_0}$ and its 811 rest mass $m_0 = -\overline{W}_0$ (for the velocity u = 0 at the initial time moment t = 0): 812

813
$$E_0 = m_0 \frac{(1 - |\xi A_0 / m_0|^2)}{(1 - 2|\xi A_0 / m_0|^2)^{1/2}},$$
 (106)

or, equivalently, at the condition $|\xi A_0 / m_0|^2 < 1/2$ 814

815
$$m_0 = \mathcal{E}_0 \left(\frac{1}{2} \pm \frac{1}{2} \sqrt{1 - 4 |\xi A_0 / E_0|^2} + |\xi A_0 / E_0|^2 \right)^{1/2}, \qquad (107)$$

where $A_0 := A |_{t=0} \in E^3$, which strongly differs from the classical expression $m_0 = E_0 - \xi \varphi_0$, 816 following from (44) and is not depending a priori on the external potential energy $\xi arphi_0$. As the 817 quantity $|\xi A_0 / E_0| \rightarrow 0$, the following asymptotical mass values follow from (107): 818

819
$$m_0; E_0 - \frac{|\xi A_0|^4}{2|E_0|^3 E_0}, m_0^{(\pm)}; \pm \sqrt{2} |\xi A_0|.$$
(108)

The first mass value m_0 ; $E_0 - \frac{|\xi A_0|^4}{2|E_0|^3 E_0}$ is physically reasonable from the classic 820

relativistic point of view, giving rise at week enough magnetic potential to the charged particle 821 energy E_0 , yet the second mass values $m_0^{(\pm)}$; $\pm \sqrt{2} |\xi A_0|$ still need their physical interpretation, 822 as they may describe both matter and anti-matter states, consisting, at a very huge energy 823 824 modulus $|E_0| \rightarrow \infty$, of some charged particle excitations of the vacuum. It is also worth of 825 mentioning that the sign of the mass m_0 coincides with that of the energy E_0 if only the

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inequality $1 - |\xi A_0 / m_0|^2 \ge 0$ holds. 826

827

828 To make this difference more clear, we now analyze the Lorentz force (76) from the 829 Hamiltonian point of view based on the Lagrangian function (74). Thus, we obtain that the 830 corresponding Hamiltonian function

$$H := \langle P, \dot{r} \rangle - L = \langle P, \dot{r} \rangle + \overline{W} (1 + |\dot{r}|^2)^{1/2} - \xi \langle A, \dot{r} \rangle =$$

831

$$= \langle P - \xi A, \dot{r} \rangle + W(1 + |\dot{r}|^{2})^{1/2} =$$

$$= -\langle p, p \rangle \overline{W}^{-1}(1 - |p|^{2}/\overline{W}^{2})^{-1/2} + \overline{W}(1 - |p|^{2}/\overline{W}^{2})^{-1/2} =$$
(109)

$$= - \langle p, p \rangle \overline{W}^{-1} (1 - |p|^2 / \overline{W}^2)^{-1/2} + \overline{W} (1 - |p|^2 / \overline{W}^2)^{-1/2} =$$

$$= -(\overline{W}^{2} - |p|^{2})(\overline{W}^{2} - |p|^{2})^{-1/2} = -(\overline{W}^{2} - |p|^{2})^{1/2}.$$
832 Since $p = P - \xi A$, expression (109) assumes the final "*no interaction*" [5] [65] [103] [104] form
833 $H = -(\overline{W}^{2} - |P - \xi A|^{2})^{1/2},$ (110)
834 which is conserved with respect to the evolution equations (72) and (73), that is
835 $dH / d\tau = dH / dt = 0$ (111)

for all $\tau, t \in \mathbb{R}$. These equations latter are equivalent to the following Hamiltonian system 836 $\dot{r} = \partial H / \partial P = (P - \xi A)(\overline{W}^2 - |P - \xi A|^2)^{-1/2},$

837

$$\dot{P} = -\partial H / \partial r = (\overline{W}\nabla\overline{W} - \nabla < \xi A, (P - \xi A) >)(\overline{W}^2 - |P - \xi A|^2)^{-1/2}$$

as one can readily check by direct calculations. Actually, the first equation 838 $\dot{r} = (P - \xi A)(\overline{W}^2 - |P - \xi A|^2)^{-1/2} = p(\overline{W}^2 - |P|^2)^{-1/2} =$

839

$$= mu(\overline{W}^{2} - |p|^{2})^{-1/2} = -\overline{W}u(\overline{W}^{2} - |p|^{2})^{-1/2} = u(1 - |u|^{2})^{-1/2}$$

holds, owing to the condition $d\tau = dt(1-|u|^2)^{1/2}$ and definitions p := mu, $m = -\overline{W}$, postulated 840

from the very beginning. Similarly we obtain that 841

$$\dot{P} = -\nabla \overline{W} (1 - |p|^2 / \overline{W}^2)^{-1/2} + \nabla < \xi A, u > (1 - |p|^2 / \overline{W}^2)^{-1/2} =$$

842

$$= -\nabla \overline{W} (1 - |u|^2)^{-1/2} + \nabla < \xi A, u > (1 - |u|^2)^{-1/2},$$

843 coincides with equation (75) in the evolution parameter $t \in \mathbb{R}$. This can be formulated as the 844 next result.

845

846 **Proposition 8.** The dual to the classical relativistic electrodynamic model (76) allows the 847 canonical Hamiltonian formulation (112) with respect to the proper reference frame variables, where the Hamiltonian function is given by expression (110). Moreover, this formulation 848 849 circumvents the "mass-potential energy" controversy attached to the classical electrodynamic 850 model (42).

851

(112)

(113)

(114)

The modified Lorentz force expression (76) and the related rest energy relationship are characterized by the following remark.

Remark 3. If we make use of the modified relativistic Lorentz force expression (76) as an alternative to the classical one of (46), the corresponding charged particle ξ energy expression (110) also gives rise to a true physically reasonable energy expression (at the velocity $u := 0 \in E^3$ at the initial time moment t = 0); namely, $E_0 = m_0$ instead of the physically controversial classical expression $E_0 = m_0 + \xi \varphi_0$, where $\varphi_0 := \varphi|_{t=0}$, corresponding to the case (44).

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854

1.6. Conclusions

863 864 All of dynamical field equations discussed above are canonical Hamiltonian systems with respect to the corresponding proper reference frames K_r, parameterized by suitable time 865 866 parameters $\tau \in \mathbb{R}$. Upon passing to the basic laboratory reference frame K, with the time 867 parameter $t \in \mathbb{R}$, naturally the related Hamiltonian structure is lost, giving rise to a new 868 interpretation of the real particle motion. Namely, one that has an absolute sense only with 869 respect to the proper reference system, and otherwise being completely relative with respect 870 to all other reference frames. As for the Hamiltonian expressions (95), (100) and (110), one observes that they all depend strongly on the vacuum potential energy field function 871 $\overline{W}: M^4 \to R$, thereby avoiding the mass problem of the classical energy expression pointed out 872 873 by L. Brillouin Error! Reference source not found.. It should be noted that the canonical Dirac 874 quantization procedure can be applied only to the corresponding dynamical field systems 875 considered with respect to their proper reference frames.

876

Remark 4. Some comments are in order concerning the classical relativity principle. We have obtained our results relying only on the natural notion of the proper reference frame and its suitable Lorentz parametrization with respect to any other moving reference frames. It seems reasonable then that the true state changes of a moving charged particle ξ are exactly realized only with respect to its proper reference system. Then the only remaining question would be about the physical justification of the corresponding relationship between time parameters of moving and proper reference frames.

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- 885 886

The relationship between reference frames that we have used through is expressed as $d\tau = dt(1 - |u|^2)^{1/2},$ (115)

887 where $u := dr / dt \in E^3$ is the velocity vector with which the proper reference frame K_r moves 888 with respect to another arbitrarily chosen reference frame K_r . Expression (115) implies, in 889 particular, that

890

$$dt^2 - |dr|^2 = d\tau^2,$$
(116)

891 which is identical to the classical infinitesimal Lorentz invariant. This is not a coincidence, since 892 all our dynamical vacuum field equations were derived in turn [18][19] from the governing 893 equations of the vacuum potential field function $W: M^4 \rightarrow R$ in the form

894
$$\partial^2 W / \partial t^2 - \nabla^2 W = \xi \rho, \partial W / \partial t + \nabla (vW) = 0, \partial \rho / \partial t + \nabla (v\rho) = 0, \quad (117)$$

895 which is *a priori* Lorentz invariant. Here $\rho \in \mathbb{R}$ is the charge density and v := dr/dt the 896 associated local velocity of the vacuum field potential evolution. Consequently, the dynamical 897 infinitesimal Lorentz invariant (116) reflects this intrinsic structure of equations (117). If it is 898 rewritten in the following nonstandard Euclidean form:

$$dt^2 = d\tau^2 + |dr|^2$$
(118)

900 it gives rise to a completely different relationship between the reference frames K_r and K_r , 901 namely

899

$$dt = d\tau (1 + |\dot{r}|^2)^{1/2}, \tag{119}$$

where $\dot{r} := dr / d\tau$ is the related particle velocity with respect to the proper reference system. 903 904 Thus, we observe that all our Lagrangian analysis in this Section is based on the corresponding 905 functional expressions written in these "Euclidean" space-time coordinates and with respect to 906 which the least action principle was applied. So we see that there are two alternatives - the first 907 is to apply the least action principle to the corresponding Lagrangian functions expressed in the 908 Minkowski space-time variables with respect to an arbitrarily chosen reference frame K, and 909 the second is to apply the least action principle to the corresponding Lagrangian functions 910 expressed in Euclidean space-time variables with respect to the proper reference frame K_{r} .

911 This leads us to a slightly amusing but thought-provoking observation: It follows from 912 our analysis that all of the results of classical special relativity related with the electrodynamics 913 of charged point particles can be obtained (in a one-to-one correspondence) using of our new 914 definitions of the dynamical particle mass and the least action principle with respect to the 915 associated Euclidean space-time variables in the proper reference system.

916 An additional remark concerning the quantization procedure of the proposed 917 electrodynamics models is in order: If the dynamical vacuum field equations are expressed in 918 canonical Hamiltonian form, as we have done in this paper, only straightforward technical 919 details are required to quantize the equations and obtain the corresponding Schrödinger 920 evolution equations in suitable Hilbert spaces of quantum states. There is another striking 921 implication from our approach: the Einstein equivalence principle [1] [5] [65] [89] is rendered 922 superfluous for our vacuum field theory of electromagnetism and gravity.

923 Using the canonical Hamiltonian formalism devised here for the alternative charged 924 point particle electrodynamics models, we found it rather easy to treat the Dirac quantization. 925 The results obtained compared favorably with classical guantization, but it must be admitted 926 that we still have not given a compelling physical motivation for our new models. This is 927 something that we plan to revisit in future investigations. Another important aspect of our 928 vacuum field theory no-geometry (geometry-free) approach to combining the electrodynamics 929 with the gravity, is the manner in which it singles out the decisive role of the proper reference 930 frame K_r. More precisely, all of our electrodynamics models allow both the Lagrangian and 931 Hamiltonian formulations with respect to the proper reference system evolution parameter 932 $\tau \in \mathbb{R}$, which are well suited the to canonical quantization. The physical nature of this fact 933 remains is as yet not quite clear. In fact, as far as we know [5] [65] [75] [76] [89], there is no 934 physically reasonable explanation of this decisive role of the proper reference system, except for that given by R. Feynman who argued in [1] that the relativistic expression for the classical Lorentz force (46) has physical sense only with respect to the proper reference frame variables $(\tau, r) \in \mathbb{R} \times \mathbb{E}^3$. In future research we plan to analyze the quantization scheme in more detail and begin work on formulating a vacuum quantum field theory of infinitely many particle systems.

941 941 2. The Lorentz type force analysis within the Feynman proper time 942 paradigm and the radiation theory

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- 945 946

2.1. Introductory setting

947 The elementary point charged particle, like electron, mass problem was inspiring many 948 physicists [20] from the past as J. J. Thompson, G.G. Stokes, H.A. Lorentz, E. Mach, M. Abraham, 949 P.A. M. Dirac, G.A. Schott and others. Nonetheless, their studies have not given rise to a clear 950 explanation of this phenomenon that stimulated new researchers to tackle it from different 951 approaches based on new ideas stemming both from the classical Maxwell-Lorentz electromagnetic theory, as in [1] [12] [21] [22] [24] [25] [26] [27] [28] [29] [30] [31] [32] [33] 952 953 [34] [35] [36] [37] [39] [74] [105] [106] [107], and modern quantum field theories of Yang-Mills 954 and Higgs type, as in [40] [41] [43] [108] and others, whose recent and extensive review is 955 done in [44].

956 In the present work I will mostly concentrate on detail analysis and consequences of the 957 Feynman proper time paradigm [1] [22] [45] [46] subject to deriving the electromagnetic 958 Maxwell equations and the related Lorentz like force expression considered from the vacuum 959 field theory approach, developed in works [49] [50] [51], and further, on its applications to the 960 electromagnetic mass origin problem. Our treatment of this and related problems, based on 961 the least action principle within the Feynman proper time paradigm [1], has allowed to 962 construct the respectively modified Lorentz type equation for a moving in space and radiating 963 energy charged point particle. Our analysis also elucidates, in particular, the computations of 964 the self-interacting electron mass term in [29], where there was proposed a not proper solution 965 to the well known classical Abraham-Lorentz [52] [53] [54] [55] and Dirac [56] electron 966 electromagnetic "4/3-electron mass" problem. As a result of our scrutinized studying the 967 classical electromagnetic mass problem we have stated that it can be satisfactory solved within 968 the classical H. Lorentz and M. Abraham reasonings augmented with the additional electron 969 stability condition, which was not taken before into account yet appeared to be very important 970 for balancing the related electromagnetic field and mechanical electron momenta. The latter, 971 following recent enough works [31] [35], devoted to analyzing the electron charged shell 972 model, can be realized within there suggested pressure-energy compensation principle, suitably 973 applied to the ambient electromagnetic energy fluctuations and the own electrostatic Coulomb 974 electron energy.

- 975
- 976

2.2. Feynman proper time paradigm geometric analysis

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978 In this section, we will develop further the vacuum field theory approach within the 979 Feynman proper time paradigm, devised before in [49] [51], to the electromagnetic J.C. Maxwell and H. Lorentz electron theories and show that they should be suitably modified: 980 981 namely, the basic Lorentz force equations should be generalized following the Landau-Lifschitz 982 least action recipe [5], taking also into account the pure electromagnetic field impact. When 983 applied the devised vacuum field theory approach to the classical electron shell model, the 984 resulting Lorentz force expression appears to satisfactorily explain the electron inertial mass 985 term exactly coinciding with the electron relativistic mass, thus confirming the well known 986 assumption [2] [109] by M. Abraham and H. Lorentz.

987 As was reported by F. Dyson [45] [46], the original Feynman approach derivation of the 988 electromagnetic Maxwell equations was based on an *a priori* general form of the classical 989 Newton type force, acting on a charged point particle moving in three-dimensional space R^3 990 endowed with the canonical Poisson brackets on the phase variables, defined on the associated tangent space $T(\mathbf{R}^3)$. As a result of this approach there only the first part of the Maxwell 991 992 equations were derived, as the second part, owing to F. Dyson [45], is related with the charged 993 matter nature, which appeared to be hidden. Trying to complete this Feynman approach to the 994 derivation of Maxwell's equations more systematically we have observed [49] that the original 995 Feynman's calculations, based on Poisson brackets analysis, were performed on the tangent space $T(\mathbb{R}^3)$. which is, subject to the problem posed, not physically proper. The true Poisson 996 997 brackets can be correctly defined only on the *coadjoint phase space* $T^*(\mathbb{R}^3)$. as seen from the 998 classical Lagrangian equations and the related Legendre transformation [47] [64] [96] [110] 999 from $T(\mathbf{R}^3)$. to $T^*(\mathbf{R}^3)$. Moreover, within this observation, the corresponding dynamical Lorentz type equation for a charged point particle should be written for the particle 1000 1001 momentum, not for the particle velocity, whose value is well defined only with respect to the 1002 proper relativistic reference frame, associated with the charged point particle owing to the fact 1003 that the Maxwell equations are Lorentz invariant.

1004 Thus, from the very beginning, we shall reanalyze the structure of the Lorentz force 1005 exerted on a moving charged point particle with a charge $\xi \in \mathbb{R}$ by another point charged 1006 particle with a charge $\xi_f \in \mathbb{R}$, making use of the classical Lagrangian approach, and rederive 1007 the corresponding electromagnetic Maxwell equations. The latter appears to be strongly 1008 related to the charged point mass structure of the electromagnetic origin as was suggested by 1009 R. Feynman and F. Dyson.

1010 Consider a charged point particle moving in an electromagnetic field. For its description, 1011 it is convenient to introduce a trivial fiber bundle structure $\pi : M \to R^3, M = R^3 \times G$, with the 1012 abelian structure group $G := R \setminus \{0\}$, equivariantly acting on the canonically symplectic 1013 coadjoint space $T^*(M)$ endowed both with the canonical symplectic structure

1015 for all $(p, y; r, g) \in T^*(M)$, where $\alpha^{(1)}(r, g) := \langle p, dr \rangle + \langle y, g^{-1}dg \rangle_G \in T^*(M)$ is the 1016 corresponding Liouville form on M, and with a connection one-form $A: M \to T^*(M) \times G$ as

$$A(r,g) := g^{-1} < \xi A(r), dr > g + g^{-1} dg,$$
(121)

with $\xi \in G^*, (r,g) \in \mathbb{R}^3 \times G$, and $\langle \cdot, \cdot \rangle$ being the scalar product in \mathbb{E}^3 . The corresponding 1018 curvature 2-form $\Sigma^{(2)} \in \Lambda^2(\mathbb{R}^3) \otimes \mathbb{G}$ is 1019

1020
$$\Sigma^{(2)}(r) := dA(r,g) + A(r,g) \wedge A(r,g) = \xi \sum_{i,j=1}^{3} F_{ij}(r) dr^{i} \wedge dr^{j},$$
(122)

1021 where

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$$F_{ij}(r) := \frac{\partial A_j}{\partial r^i} - \frac{\partial A_i}{\partial r^j}$$
(123)

for $i, j = \overline{1,3}$ is the electromagnetic tensor with respect to the reference frame K_i , 1023 characterized by the phase space coordinates $(r, p) \in T^*(\mathbb{R}^3)$. As an element $\xi \in \mathbb{G}^*$ is still not 1024 fixed, it is natural to apply the standard [47] [64] [96] [110] invariant Marsden-Weinstein-1025 Meyer reduction to the orbit factor space $\tilde{P}_{\xi} := P_{\xi} / G_{\xi}$ subject to the related momentum 1026 mapping $l:T^*(M) \to G^*$, constructed with respect to the canonical symplectic structure (120) 1027 on $T^*(M)$, where, by definition, $\xi \in G^*$ is constant, $P_{\xi} := l^{-1}(\xi) \subset T^*(M)$ and 1028 $G_{\xi} = \{g \in G : Ad_{G}^{*}\xi\}$ is the isotropy group of the element $\xi \in G^{*}$. 1029

As a result of the Marsden-Weinstein-Meyer reduction, one finds that G_{ξ} ; G, the 1030 1031 $\overline{\omega}_{\xi}^{(2)} \in T^*(\tilde{P}_{\xi})$ and the corresponding Poisson brackets on the reduced manifold \tilde{P}_{ξ} are 1032

1033
$$\{r^{i}, r^{j}\}_{\xi} = 0, \{p_{j}, r^{i}\}_{\xi} = \delta^{i}_{j},$$

$$\{p_{i}, p_{j}\}_{\xi} = \xi F_{ii}(r)$$
(124)

for $i, j = \overline{1,3}$, considered with respect to the reference frame K. Introducing a new 1034 1035 momentum variable

$$\tilde{\pi} := p + \xi A(r) \tag{125}$$

on \tilde{P}_{ξ} , it is easy to verify that $\overline{\omega}_{\xi}^{(2)} \rightarrow \tilde{\omega}_{\xi}^{(2)} := \langle d\tilde{\pi}, \wedge dr \rangle$, giving rise to the following "minimal 1037 interaction" canonical Poisson brackets: 1038

 $\{r^{i},r^{j}\}_{\tilde{\omega}_{\xi}^{(2)}}=0,\{\tilde{\pi}_{j},r^{i}\}_{\tilde{\omega}_{\xi}^{(2)}}=\delta^{i}_{j},\{\tilde{\pi}_{i},\tilde{\pi}_{j}\}_{\tilde{\omega}_{\xi}^{(2)}}=0$ (126)

for $i, j = \overline{1,3}$ with respect to some new reference frame \tilde{K}_{j} , characterized by the phase space 1040 coordinates $(r, \tilde{\pi}) \in \tilde{P}_{\xi}$ and a new evolution parameter $t \in \mathbb{R}$ if and only if the Maxwell field 1041 compatibility equations 1042 1043

$$\partial F_{ij} / \partial r_k + \partial F_{jk} / \partial r_i + \partial F_{ki} / \partial r_j = 0$$
(127)

are satisfied on \mathbb{R}^3 for all $i, j, k = \overline{1,3}$ with the curvature tensor (123). 1044

Now we proceed to a dynamic description of the interaction between two moving 1045 charged point particles ξ and ξ_f , moving respectively, with the velocities u := dr / dt and 1046 $u_f := dr_f / dt$ subject to the reference frame K_f . Unfortunately, there is a fundamental problem 1047 in correctly formulating a physically suitable action functional and the related least action 1048

1049 condition. There are clearly possibilities such as

1050

$$S_{p}^{(t)} := \int_{t_{1}}^{t_{2}} dt \mathcal{L}_{p}^{(t)}[r; dr / dt]$$
(128)

on a temporal interval $[t_1, t_2] \subset \mathbb{R}$ with respect to the laboratory reference frame K_t , 1051

1052

$$S_{p}^{(t')} := \int_{t_{1}}^{t_{2}} dt' L_{p}^{(t')}[r; dr / dt']$$
(129)

on a temporal interval $[t_1, t_2] \subset R$ with respect to the moving reference frame K_1 and 1053

1054
$$S_{p}^{(\tau)} := \int_{\tau_{1}}^{\tau_{2}} d\tau L_{p}^{(\tau)}[r; dr / d\tau]$$
(130)

on a temporal interval $[\tau_1, \tau_2] \subset \mathbb{R}$ with respect to the proper time reference frame K_r , 1055 1056 naturally related to the moving charged point particle ξ .

1057 It was first observed by Poincaré and Minkowski [65] that the temporal differential $d\tau$ 1058 is not a closed differential one-form, which physically means that a particle can traverse many different paths in space R^3 with respect to the reference frame K_r , during any given proper 1059 time interval $d\tau$, naturally related to its motion. This fact was stressed [65] [111] [112] [113] 1060 1061 [114] by Einstein, Minkowski and Poincaré, and later exhaustively analyzed by R. Feynman, who argued [1] that the dynamical equation of a moving point charged particle is physically sensible 1062 only with respect to its proper time reference frame. This is Feynman's proper time reference 1063 frame paradigm, which was recently further elaborated and applied both to the 1064 1065 electromagnetic Maxwell equations in [23] [24] [74] and to the Lorentz type equation for a 1066 moving charged point particle under external electromagnetic field in [47] [49] [50] [51]. As it 1067 was there argued from a physical point of view, the least action principle should be applied only to the expression (130) written with respect to the proper time reference frame K_{r} , whose 1068 1069 temporal parameter $\tau \in \mathbf{R}$ is independent of an observer and is a closed differential one-form. 1070 Consequently, this action functional is also mathematically sensible, which in part reflects the Poincaré's and Minkowski's observation that the infinitesimal guadratic interval 1071

- $d\tau^{2} = (dt')^{2} |dr dr_{f}|^{2},$ 1072 (131)
- relating the reference frames K_{τ} and K_{τ} , can be invariantly used for the four-dimensional 1073 relativistic geometry. The most natural way to contend with this problem is to first consider the 1074 quasi-relativistic dynamics of the charged point particle ξ with respect to the moving 1075 reference frame $\mathbf{K}_{_{f}}$ subject to which the charged point particle $\boldsymbol{\xi}_{_{f}}$ is at rest. Therefore, it 1076 possible to write down a suitable action functional (129), up to $O(1/c^4)$, as the light velocity 1077 $c \rightarrow \infty$, where the quasi-classical Lagrangian function $L_p^{(i')}[r; dr / dt']$ can be naturally chosen 1078 1079 as
- $L_{p}^{(t')}[r; dr / dt'] := m'(r) |dr / dt' dr_{f} / dt'|^{2} / 2 \xi \varphi'(r).$ 1080 (132)

where $m'(r) \in \mathbb{R}_+$ is the charged particle ξ inertial mass parameter and $\varphi'(r)$ is the potential 1081 function generated by the charged particle ξ_f at a point $r \in \mathbb{R}^3$ with respect to the reference 1082

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frame K_i . Since the standard temporal relationships between reference frames K_i and K_j : 1083 $dt' = dt(1 - |dr_f / dt'|^2)^{1/2},$ 1084 (133)1085 $d\tau = dt' (1 - \left| dr / dt' - dr_f / dt' \right|^2)^{1/2},$ 1086 (134)give rise, up to $O(1/c^2)$, as $c \to \infty$, to dt'; dt and $d\tau$; dt', respectively, it is easy to verify 1087 that the least action condition $\delta S_p^{(i')} = 0$ is equivalent to the dynamical equation 1088 $d\pi / dt = \nabla \mathcal{L}_p^{(t')}[r; dr / dt] = \left(\frac{1}{2} \left| \frac{dr}{dt} - \frac{dr_f}{dt} \right|^2\right) \nabla m - \xi \nabla \varphi(r),$ 1089 (135)where we have defined the generalized canonical momentum as 1090 $\pi := \partial \mathcal{L}_{n}^{(t')}[r; dr / dt] / \partial (dr / dt) = m(dr / dt - dr_{f} / dt),$ 1091 (136)with the dash signs dropped and denoted by " ∇ " the usual gradient operator in E^3 . Equating 1092 the canonical momentum expression (136) with respect to the reference frame K_{t} to that of 1093 (125) with respect to the canonical reference frame $\tilde{K}_{_{f}}$, and identifying the reference frame 1094 \tilde{K} , with K_{ℓ} , one obtains that 1095 $m(dr / dt - dr_f / dt) = mdr / dt - \xi A(r),$ 1096 (137)giving rise to the important inertial particle mass determining expression 1097 1098 $m = -\xi \varphi(r),$ (138)which right away follows from the relationship 1099 $\varphi(r)dr_f / dt = A(r).$ 1100 (139)The latter is well known in the classical electromagnetic theory [2] [5] for potentials 1101 $(\varphi, A) \in T^*(M^4)$ satisfying the Lorentz condition 1102 1103 $\partial \varphi(r) / \partial t + \langle \nabla, A(r) \rangle = 0,$ (140)1104 yet the expression (138) looks very nontrivial in relating the "inertial" mass of the charged point particle ξ to the electric potential, being both generated by the ambient charged point 1105 particles ξ_{f} . As was argued in articles [49] [50], the above mass phenomenon is closely related 1106 1107 and from a physical perspective shows its deep relationship to the classical electromagnetic 1108 mass problem. Before further analysis of the completely relativistic the charge ξ motion under 1109 1110 consideration, we substitute the mass expression (138) into the quasi-relativistic action functional (129) with the Lagrangian (132). As a result, we obtain two possible action functional 1111 expressions, taking into account two main temporal parameters choices: 1112 $S_{p}^{(t')} = -\int_{t_{i}}^{t_{2}} \xi \varphi'(r) (1 + \frac{1}{2} \left| dr / dt' - dr_{f} / dt' \right|^{2}) dt'$ 1113 (141)

1114 on an interval $[t_1^{'}, t_2^{'}] \subset \mathbb{R}$, or

1115
$$S_{p}^{(\tau)} = -\int_{\tau_{1}}^{\tau_{2}} \xi \varphi'(r) (1 + \frac{1}{2} \left| dr / d\tau - dr_{f} / d\tau \right|^{2}) d\tau$$
(142)

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1116 on an $[\tau_1, \tau_2] \subset \mathbb{R}$. The direct relativistic transformations of (142) entail that

1117

$$S_{p}^{(\tau)} = -\int_{\tau_{1}}^{\tau_{2}} \xi \varphi'(r) (1 + \frac{1}{2} \left| dr / d\tau - dr_{f} / d\tau \right|^{2}) d\tau ;$$

$$; -\int_{\tau_{1}}^{\tau_{2}} \xi \varphi'(r) (1 + \left| dr / d\tau - dr_{f} / d\tau \right|^{2})^{1/2} d\tau = (143)$$

$$= -\int_{\tau_{1}}^{\tau_{2}} \xi \varphi'(r) (1 - \left| dr / dt' - dr_{f} / dt' \right|)^{-1/2} d\tau = -\int_{t_{1}}^{t_{2}} \xi \varphi'(r) dt',$$

giving rise to the correct, from the physical point of view, relativistic action functional form (129), suitably transformed to the proper time reference frame representation (130) via the Feynman proper time paradigm. Thus, we have shown that the true action functional procedure consists in a physically motivated choice of either the action functional expression form (128) or (129). Then, it is transformed to the proper time action functional representation form (130) within the Feynman paradigm, and the least action principle is applied.

1124 Concerning the above discussed problem of describing the motion of a charged point 1125 particle ξ in the electromagnetic field generated by another moving charged point particle ξ_f , 1126 it must be mentioned that we have chosen the quasi-relativistic functional expression (132) in 1127 the form (129) with respect to the moving reference frame K_t , because its form is physically 1128 reasonable and acceptable, since the charged point particle ξ_f is then at rest, generating no 1129 magnetic field.

Based on the above relativistic action functional expression

1131
$$S_{p}^{(\tau)} := -\int_{\tau_{1}}^{\tau_{2}} \xi \varphi'(r) (1 + \left| dr / d\tau - dr_{f} / d\tau \right|^{2})^{1/2} d\tau$$
(144)

1132 written with respect to the proper reference from K_{τ} , one finds the following evolution 1133 equation:

1134 $d\pi_{p} / d\tau = -\xi \nabla \varphi'(r) (1 + |dr / d\tau - dr_{f} / d\tau|^{2})^{1/2}, \qquad (145)$

1135 where the generalized momentum is given exactly by the relationship (136):

$$\pi_{p} = m(dr / dt - dr_{f} / dt).$$
(146)

1137Making use of the relativistic transformation (133) and the next one (134), the equation1138(145) is easily transformed to

$$\frac{d}{dt}(p+\xi A) = -\nabla\varphi(r)(1-\left|u_{f}\right|^{2}), \qquad (147)$$

1140 where we took into account the related definitions: (138) for the charged particle ξ mass, (139) 1141 for the magnetic vector potential and $\varphi(r) = \varphi'(r) / (1 - |u_f|^2)^{1/2}$ for the scalar electric potential 1142 with respect to the laboratory reference frame K_i. Equation (147) can be further transformed, 1143 using elementary vector algebra, to the classical Lorentz type form: 1144 $dp / dt = \xi E + \xi u \times B - \xi \nabla < u - u_f, A >,$ (148)

1145 where

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- $E := -\partial A / \partial t \nabla \varphi \tag{149}$
- 1147 is the related electric field and

1148 $B := \nabla \times A$ (150) 1149 is the related magnetic field, exerted by the moving charged point particle ξ_f on the charged 1150 point particle ξ with respect to the laboratory reference frame K_t . The Lorentz type force 1151 equation (148) was obtained in [49] [50] in terms of the moving reference frame K_t , and 1152 recently reanalyzed in [34] [50]. The obtained results follow in part [16] [17] from Ampère's 1153 classical works on constructing the magnetic force between two neutral conductors with 1154 stationary currents.

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- 1156 1157

3. The self-interaction problem: historical preliminaries

1158 The elementary point charged particle, like electron, mass problem was inspiring many 1159 physicists [20] from the past as J. J. Thompson, G.G. Stokes, H.A. Lorentz, E. Mach, M. Abraham, 1160 P.A. M. Dirac, G.A. Schott, J. Schwinger and many others. Nonetheless, their studies have not 1161 given rise to a clear explanation of this phenomenon that stimulated new researchers to tackle 1162 it from different approaches based on new ideas stemming both from the classical Maxwell-Lorentz electromagnetic theory, as in [1] [21] [22] [24] [25] [26] [34] [74] [109], and modern 1163 quantum field theories of Yang-Mills and Higgs type, as in [40] [41] [43] [108] and others, 1164 1165 whose recent and extensive review is done in [44].

1166 In the present work we mostly concentrate on detailed quantum and classical analysis 1167 of the self-interacting shell model charged particle within the Fock many-temporal approach 1168 [115] [116] and the Feynman proper time paradigm [1] [22] [45] [46] subject to deriving the 1169 electromagnetic Maxwell equations and the related Lorentz like force expression within the 1170 vacuum field theory approach, devised in works [24] [49] [50] [51] [74] [117], and further, we 1171 elaborate the obtained results to treating the classical H. Lorentz and M. Abraham [12] [27] 1172 [28] [29] [30] [31] [32] [33] [35] [36] [37] [39] [52] [53] [54] [107] [118] electromagnetic mass origin problem. For the first time the proper time approach to classical electrodynamics 1173 and quantum mechanics was possibly suggested still in 1937 by V. Fock [119], in which, in 1174 1175 particular, there was constructed an alternative proper time based Lagrangian description of a 1176 point charged particle under external electromagnetic field. A more detailed motivation of 1177 using the proper time approach was later presented by R. Feynman in his Lectures [1]. 1178 Concerning the alternative and much later investigations of the *a priori* given quantum 1179 electromagnetic Maxwell equations in the Fock space one can mention the Gupta-Bleiler [120] 1180 [121] [122] and [61] [71] [88] approaches. The first one, as it is well known [71] [121], 1181 contradicts a one of the most important field theoretical principles - the positive definiteness of the quantum event probability and is strongly based on making nonphysical use of an indefinite 1182 1183 metric on quantum states. The second one is completely non-relativistic and based on the 1184 canonical quantization scheme [71] in the case of the Coulomb gauge condition. Inspired by 1185 these and related classical results, we have stated that the self-interacting quantum mechanism 1186 of the charged particle with its self-generated electromagnetic field consists of two physically different phenomena, whose influence on the structure of the resulting Hamilton interaction 1187 1188 operator appeared to be crucial and gave rise to a modified analysis of the related classical shell 1189 model charged particle within the Lagrangian formalism. As a result of our scrutinized studying 1190 the classical electromagnetic mass problem there was demonstrated that it can be satisfactory

solved within the classical H. Lorentz and M. Abraham reasonings augmented with the 1191 1192 additional electron stability condition, which was not taken before into account yet appeared to 1193 be very important for balancing the related electromagnetic field and mechanical electron 1194 momenta. The latter, following the recent enough works [31] [35] [118] devoted to analyzing 1195 the electron charged shell model, was realized within there suggested *pressure-energy* compensation principle, suitably applied to the ambient electromagnetic energy fluctuations 1196 1197 and the self-generated electrostatic Coulomb electron energy. In the case of a point charged particle the alternative relativistic invariant approach to studying the radiation reaction force 1198 1199 was before suggested by Teitelbom [37], which was based on a formal relativistic invariant 1200 splitting of the electromagnetic energy-momentum tensor and deriving the related suitably 1201 renormalized charged particle equations of motion.

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4. The charged particle self-interaction quantum origin

1205Consider a free relativistic quantum fermionic *a priori* massless particle field described1206[121] [123] by means of the secondly-quantized self-adjoint Dirac-Weil type Hamiltonian

1207
$$H_{f} = \int_{\mathbb{R}^{3}} d^{3} x \psi^{+} < c \alpha, \frac{\hbar}{i} \nabla > \psi, \qquad (151)$$

1208 where $\alpha \in E^3 \otimes End M^4$ denotes the standard Dirac spin matrix vector representation in the 1209 Minkowski space M^4 , $c \in \mathbb{R}_+$ is the light velocity, $\langle \cdot, \cdot \rangle$ denotes the usual scalar product in 1210 the Euclidean space E^3 , $\psi: \mathbb{R}^3 \to (End \Phi)^4$ - a spinor of the quantum annihilation operators, 1211 acting in a suitable Fock space Φ endowed with the usual scalar product (\cdot, \cdot) and 1212 $\psi^+: \mathbb{R}^3 \to (End \Phi)^4$ - the respectively adjoint co-spinor of creation operators in the Fock space 1213 Φ . The following anticommuting [121] [123] operator relationships

$$\psi_i(x)\psi_i^+(y)+\psi_i^+(y)\psi_i(x) = \delta_{il}\delta(x-y),$$

1214

$$\Psi_{i}(x)\Psi_{l}(y) + \Psi_{l}(y)\Psi_{i}(x) = 0,$$
 (152)

$$\psi_{i}^{+}(x)\psi_{l}^{+}(y) + \psi_{l}^{+}(y)\psi_{i}^{+}(x) = 0$$

hold for any $x, y \in \mathbb{R}^3$ and $j, l \in \overline{1, 4}$, being compatible with the related Heisenberg operator dynamics, generated by the fermionic Hamiltonian operator (151):

1217
$$\partial \psi / \partial \overline{t} := \frac{i}{\hbar} [H_f, \psi], \ \partial \psi^+ / \partial \overline{t} := \frac{i}{\hbar} [H_f, \psi^+]$$
(153)

1218 with respect to its own laboratory reference frame $K_{\overline{t}}$, parameterized by the evolution 1219 parameter $\overline{t} \in \mathbb{R}$.

1220 It is clear that the Hamiltonian (151) possesses no information of such an important 1221 characteristic as the electric charge $\xi \in \mathbb{R}$, which generates the own electromagnetic field 1222 interacting both with it and with other ambient charged particles. As it is usually accepted, we 1223 will model a free electromagnetic field by its bosonic self-adjoint operator four-potential 1224 $(\varphi, A): \mathbb{R}^3 \to Hom \ (\Phi, \Phi^4)$, whose evolution is generated by the self-adjoint Hamiltonian

1225
$$H_{b} = 2 \int_{\mathbb{R}^{3}} d^{3}k |k|^{2} [\langle A^{+}(k), A(k) \rangle - \varphi(k)\varphi^{+}(k)], \qquad (154)$$

1226acting in the before introduced common Fock space Φ and represented by means of the field1227operators expanded into the Fourier integrals

$$\varphi(x) := \frac{1}{(2\pi)^{3/2}} \int_{\mathbb{R}^3} d^3 k \varphi(k) \exp(i < k, x >) + \frac{1}{(2\pi)^{3/2}} \int_{\mathbb{R}^3} d^3 k \varphi^+(k) \exp(-i < k, x >),$$
(155)

1228

$$A(x) := \frac{1}{(2\pi)^{3/2}} \int_{\mathbb{R}^3} d^3 k A(k) \exp(i < k, x >) + \frac{1}{(2\pi)^{3/2}} \int_{\mathbb{R}^3} d^3 k A^+(k) \exp(-i < k, x >).$$

1229 The coefficients of the expansions (155) satisfy the following [115] [116] [121] 1230 commutation operator relationships:

$$\varphi(k), \varphi^+(s)] = -\frac{c\hbar}{2|k|}\delta(k-s)$$

 $\varphi(k), A_i(s)] = 0,$

1231
$$[\varphi(k),\varphi(s)] = 0 = [\varphi^+(k),\varphi^+(s)],$$
(156)

$$[A_j(k), A_l^+(s)] = \frac{c\hbar}{2|k|} \delta_{jl} \delta(k-s).$$

$$[A_{j}(k), A_{l}(s)] = 0 = [A_{j}^{+}(k), A_{l}^{+}(s)]$$

- for all $k, s \in E^3$ and $j, l \in \overline{1,3}$, compatible with the related Heisenberg operator dynamics [121] generated by the electromagnetic field Hamiltonian (154):
- 1234 $\frac{\partial A}{\partial \tilde{t}} := \frac{i}{\hbar} [H_b, A], \frac{\partial \varphi}{\partial \tilde{t}} := \frac{i}{\hbar} [H_b, \varphi], \qquad (157)$
- 1235 with respect to its own laboratory reference frame $K_{\tilde{t}}$, parameterized by the temporal 1236 parameter $\tilde{t} \in \mathbb{R}$. In particular, based on the commutation relationships (156), one can check 1237 that the electric

$$E := -\nabla \varphi - \frac{1}{c} \frac{\partial A}{\partial \tilde{t}}$$
(158)

and magnetic

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1240

1243

- $B := \nabla \times A \tag{159}$
- 1241 fields satisfy the operator Maxwell equations in vacuum, and the following weak Lorenz type 1242 constraints

 $C_0(k)\Phi := i[< k, A(k) > -|k|\varphi(k)]\Phi = 0,$ (160)

$$C_0^+(k)\Phi := -i[\langle k, A^+(k) \rangle - |k| \varphi^+(k)]\Phi = 0$$

1244 hold in the Fock space Φ for all $k \in E^3$. As the operators $C_0(k): \Phi \to \Phi$ and $C_0^+(k): \Phi \to \Phi$

1245 are commuting both to each other for all $k \in E^3$ and with the Hamiltonian (154), that is $C_0(k), C_0(l)] = 0 = [C_0(k), C_0^+(l)],$ 1246 $C_0(k), H_b] = 0 = [C_0^+(k), H_b]$ 1247 for any $k, l \in E^3$, the constraints (160) are compatible with the evolution operator equations

1248 (157). Moreover, concerning the Hamiltonian operator (154), whose equivalent operator 1249 expression is

 $H_{b} = \frac{1}{2} \int_{\mathbb{R}^{3}} (|E|^{2} + |B|^{2}),$ (162)

1251 the following proposition holds.

1252

1253 **Proposition 9.** The Hamiltonian operator (154) on the reduced by means of constraints (160) 1254 Fock subspace Φ is Hermitian and non-negative definite. 1255

1256 *Proof.* Really, if to define the operator

$$B(k) := A(k) - \frac{k}{|k|^2} < k, A(k) >,$$
(163)

1258 the Hamiltonian operator (154) can be rewritten equivalently as

$$H_{b} = 2 \int_{\mathbb{R}^{3}} d^{3}k |k|^{2} \left\{ < \frac{k}{|k|} \times B^{+}(k), \frac{k}{|k|} \times B(k) > + \right\}$$
(164)

1259

1257

$$+\frac{i}{|k|}\varphi(k)C_{0}^{+}(k)+\frac{1}{|k|^{2}}[_{0}^{+}(k)-i|k|\varphi^{+}(k)]C_{0}(k)\}.$$

1260 The latter, owing to the weak Lorenz type constraints (160), gives rise to the inequality

$$(f, H_b f) = 2 \int_{\mathbb{R}^3} d^3k \, |k|^2 | \left(< \frac{k}{|k|} \times B(k) f, \frac{k}{|k|} \times B(k) f > \right) =$$
(165)

1261

$$= 2 \int_{\mathbb{R}^3} d^3 k \, \| \, k \times B(k) f \, \|^2 \ge 0$$

1262 for any vector $f \in \Phi$, proving the proposition.

1263

1264**Remark 5.** The Hamiltonian operator expression (154) easily follows [116] [121] [123] from the1265well known relativistic invariant classical Fock-Podolsky electromagnetic Lagrangian

$$L_{b} := \frac{1}{2} \int_{\mathbb{R}^{3}} d^{3}x [<\nabla \varphi + \frac{1}{c} \frac{\partial A}{\partial \tilde{t}}, \nabla \varphi + \frac{1}{c} \frac{\partial A}{\partial \tilde{t}} > -$$
(166)

1266

$$< \nabla \times A, \nabla \times A > -(\frac{1}{c} \frac{\partial \varphi}{\partial \tilde{t}} + \langle \nabla, A \rangle)^{2}]$$

1267 Based on the corresponding to (166) Euler-Lagrange equations one finds that

1268
$$\frac{1}{c^2}\frac{\partial^2 A}{\partial \tilde{t}^2} - \Delta A = 0, \frac{1}{c^2}\frac{\partial^2 \varphi}{\partial \tilde{t}^2} - \Delta \varphi = 0, \qquad (167)$$

1269 whose wave solutions allow to determine the electromagnetic fields (158) and (159) and to 1270 check that the related Maxwell field equations in vacuum are satisfied if the Lorenz condition

1271
$$C_0(\overline{t}, x) := \frac{1}{c} \frac{\partial \varphi}{\partial \tilde{t}} + \langle \nabla, A \rangle = 0$$
(168)

1272 holds for all $(\tilde{t}, x) \in M^4$. Moreover, from the Lagrangian expression (166) one easily obtains by 1273 means of the corresponding Legendre transformation [64] [96] [121] the Hamiltonian operator

1274
$$H_{b} = \frac{1}{2} \int_{\mathbb{R}^{3}} d^{3}x (|E|^{2} + |B|^{2} - C_{0}^{2}) + \int_{\mathbb{R}^{3}} d^{3}x (\langle \nabla, A \rangle^{2} - \langle \nabla \varphi, \nabla \varphi \rangle),$$
(169)

1275 being equivalent in the Fock space Φ , modulo the solutions (155) of the wave equations (167), 1276 to the written above operator expression (154).

1278

1279

Taking into account the operator equations (157), one easily obtains that $C_0(k) = i[\langle k, A(k) \rangle - |k| \varphi(k)] \neq 0,$ (170)

$$C_0^+(k) = -i[\langle k, A^+(k) \rangle - |k| \varphi^+(k)] \neq 0,$$

1280 contradicting the imposed above Lorenz constraint (168). As the latter should be vanishing in 1281 the Fock space, it was suggested in [115] to reduce the Fock space Φ to a subspace, on which 1282 there are satisfied only the weak Lorenz type operator constraints (160). Concerning these 1283 constraints, imposed on the Fock space Φ , it is necessary to mention that a corresponding 1284 vacuum vector $|0\rangle \in \Phi$ does not, evidently, annihilate the operators $\varphi(k): \Phi \to \Phi$ and 1285 $A^+(k): \Phi \to \Phi^3$, as they do not form computing pairs with operators $C_0^+(k)$ and $C_0(k)$, 1286 respectively.

1288 5. The transformed Fock space, its Lorenz type reduction and the1289 Quantum Maxwell equations

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As we are interested in describing the self-interaction of the fermionic quantum particle 1291 field $\psi: \Phi \rightarrow \Phi^4$ with the related self-generated bosonic electromagnetic potentials field 1292 $(\varphi, A): \Phi \to \Phi^4$, we need, within the Fock many-temporal description approach [115] [116], 1293 1294 first to consider the fermionic particle and bosonic electromagnetic fields with respect to the 1295 common reference frame K, specified by the temporal parameter $t \in \mathbb{R}$. Secondly, we need to 1296 make use of the classical "minimum interaction" principle [47] [117], (whose sketched 1297 backgrounds are presented in Supplement, Section 9. and to apply to the Hamiltonian operator 1298 expression (151):

1299
$$H_f^{(int)} = \int_{\mathbb{R}^3} d^3 x \psi^+ < c\alpha, \frac{\hbar}{i} \nabla > \psi + \int_{\mathbb{R}^3} d^3 x (\xi \psi^+ \psi \varphi - \xi \psi^+ < c\alpha, A > \psi), \tag{171}$$

1300 in which the fermionic $\psi: \Phi \to \Phi^4$ and bosonic $(\varphi, A): \Phi \to \Phi^4$ operators are commuting *a* 1301 *priori* to each other as quantum fields of different nature. Since the whole quantum field

system consists of the fermionic particle and bosonic self-generated electromagnetic fields, itsevolution is described by means of the joint Hamiltonian operator

1304

$$H_{f-b} := H_f^{(int)} + H_b$$
 (172)

1305 via the Heisenberg equations

$$\frac{\partial \psi}{\partial t} := \frac{i}{\hbar} [H_{f-b}, \psi], \frac{\partial \psi^+}{\partial t} := \frac{i}{\hbar} [H_{f-b}, \psi^+],$$
(173)

1306

1313

1318

$$\frac{\partial A}{\partial t} \quad : \quad = \frac{i}{\hbar} [H_{f-b}, A], \frac{\partial \varphi}{\partial t} := \frac{i}{\hbar} [H_{f-b}, \varphi]$$

1307 with respect to the common temporal parameter $t \in \mathbb{R}$, as in this case there is assumed that 1308 the corresponding temporal parameters $\tilde{t} \in \mathbb{R}$ and $\overline{t} \in \mathbb{R}$ coincide, that is $\tilde{t} = \overline{t} = t \in \mathbb{R}$ and, 1309 by definition, the operator spinor $\psi(t, x) := \psi(\overline{t}, \tilde{t})|_{\tilde{t}=\tilde{t}=t}$. Simultaneously, there should be, 1310 evidently, satisfied the before derived both the electromagnetic field evolution equations (157) 1311 with respect to the own reference frame $K_{\overline{t}}$ and the modified fermionic charged particle field 1312 equations

$$\frac{\partial \psi}{\partial \overline{t}} := \frac{i}{\hbar} [H_f^{(int)}, \psi], \frac{\partial \psi^+}{\partial \overline{t}} := \frac{i}{\hbar} [H_f^{(int)}, \psi^+]$$
(174)

1314 with respect to the own reference frame $K_{\overline{i}}$.

1315 Being mostly interested in the evolution of the quantum particle fermionic field 1316 $\psi: \Phi \rightarrow \Phi$, we can get rid of the bosonic Hamiltonian impact into (174) having applied to the 1317 Fock space Φ the unitary canonical transformation

$$\Phi \to \tilde{\Phi} := U(t)\Phi, \tag{175}$$

1319 where we denoted by $U(t): \Phi \rightarrow \Phi$ the unitary operator satisfying the determining equation

1320
$$dU(t) / dt = \frac{i}{\hbar} H_b U(t)$$
(176)

subject to the bosonic Hamiltonian operator (154) and the temporal parameter $t \in \mathbb{R}$. As a consequence of the transformation (175) we obtain the effective fermionic particle field interaction Hamiltonian operator

$$\tilde{H}_{f}^{(int)} := U(t)H_{f}^{(int)}U^{*}(t) =$$
(177)

1324

$$= \int_{\mathbb{R}^3} d^3 x \psi^+ < c\alpha, \frac{\hbar}{i} \nabla > \psi + \int_{\mathbb{R}^3} d^3 x (\xi \psi^+ \psi \tilde{\varphi} - \xi \psi^+ < c\alpha, \tilde{A} > \psi),$$

1325 where, by definition,

1326

$$\tilde{A} := U(t)AU^*(t), \,\tilde{\varphi} := U(t)\varphi U^*(t)],$$
(178)

- 1327 subject to which the evolution in the transformed Fock space $\tilde{\Phi}$, induced by the Hamiltonian 1328 operator (154)
- 1329 $\tilde{H}_{b} := U(t)H_{b}U^{*}(t) = 2\int_{\mathbb{R}^{3}} d^{3}k |k|^{2} [\langle \tilde{A}^{+}(k), \tilde{A}(l) \rangle \tilde{\varphi}(k)\tilde{\varphi}^{+}(k)],$ (179)

became completely eliminated. Concerning the Hamiltonian operator (179) here we need to mention that the related commutation relationships for the operators $(\tilde{\varphi}(k), \tilde{A}(k)) : \tilde{\Phi} \to \tilde{\Phi}^4$

1332 and $(\tilde{\varphi}^+(k), \tilde{A}^+(k)): \tilde{\Phi} \to \tilde{\Phi}^4$ remain the same as (156), that is

$$\tilde{\varphi}(k), \tilde{\varphi}^+(s)] = -\frac{c\hbar}{2|k|} \delta(k-s), [\tilde{\varphi}(k), \tilde{A}_j(s)] = 0,$$

$$\tilde{A}_{j}(k), \tilde{A}_{l}^{+}(s)] = \frac{c\hbar}{2|k|} \delta_{jl} \delta(k-s),$$
(180)

1333

1337

$$[\tilde{\varphi}(k), \tilde{\varphi}(s)] = 0 = [\tilde{\varphi}^+(k), \tilde{\varphi}^+(s)]$$

 $[\tilde{A}_{j}(k), \tilde{A}_{l}(s)] = 0 = [\tilde{A}_{j}^{+}(k), \tilde{A}_{l}^{+}(s)],$

1334 for all $k, s \in E^3$ and $j, l \in \overline{1,3}$.

1335 Now, concerning the Hamiltonian operators (177) and (179), the following Heisenberg 1336 evolutions equations

$$\frac{\partial \psi}{\partial \overline{t}} := \frac{i}{\hbar} [\tilde{H}_{f}^{(int)}, \psi], \quad \frac{\partial \psi^{+}}{\partial \overline{t}} := \frac{i}{\hbar} [\tilde{H}_{f}^{(int)}, \psi^{+}]$$
(181)

1338 with respect to the own reference frame $K_{\overline{r}}$ and the Heisenberg evolution equations

1339
$$\frac{\partial \tilde{\varphi}}{\partial \tilde{t}} := \frac{i}{\hbar} [\tilde{H}_{b}, \tilde{\varphi}], \frac{\partial \tilde{A}}{\partial \tilde{t}} := \frac{i}{\hbar} [\tilde{H}_{b}, \tilde{A}]$$
(182)

1340 with respect to the own reference frame $K_{\tilde{i}}$ hold. Being further interested in the evolution 1341 equations (173), suitably rewritten in the transformed Fock space $\tilde{\Phi}$ with respect to the 1342 common temporal parameter $t \in \mathbb{R}$, we need to take into account [116] that the following 1343 functional relationships

$$\Psi(t) := \Psi(\overline{t}, \widetilde{t})|_{\overline{t}=\widetilde{t}=t}, \quad \widetilde{A}(t) := \widetilde{A}(\overline{t}, \widetilde{t})|_{\overline{t}=\widetilde{t}=t}$$
(183)

1345 hold. In particular, from (183) the evolution expressions follow

$$\partial \psi(t) / \partial t = \partial \psi(\overline{t}, \widetilde{t}) / \partial \overline{t} |_{\overline{t} = \widetilde{t} = t} + \partial \psi(\overline{t}, \widetilde{t}) / \partial \widetilde{t} |_{\overline{t} = \widetilde{t} = t}$$

1346

1344

 $\partial \tilde{A}(t) / \partial t = \partial \tilde{A}(\overline{t}, \tilde{t}) / \partial \overline{t} |_{\overline{t} = \tilde{t} = t} + \partial \tilde{A}(\overline{t}, \tilde{t}) / \partial \tilde{t} |_{\overline{t} = \tilde{t} = t},$

1347 for all $t \in \mathbb{R}$. The latter will be useful when deriving the resulting quantum Maxwell 1348 electromagnetic equations.

Before doing this, we need to take into account that the weak operator Lorenz constraints (160), rewritten in the transformed Fock space $\tilde{\Phi}$, is compatible with the evolution equations (182):

1352 $[\tilde{C}_{0}(k),\tilde{H}_{b}] = 0 = [\tilde{C}_{0}^{+}(k),\tilde{H}_{b}], \qquad (185)$

1353 yet they fail to be compatible with the evolution equations (181), that is

1354
$$[\tilde{C}_{0}(k), \tilde{H}_{f}^{(int)}] \neq 0 \neq [\tilde{C}_{0}^{+}(k), \tilde{H}_{f}^{(int)}]$$

1355 This means that we can not impose on the transformed Fock pace $ilde{\Phi}$ the constraints

(184)

$$\begin{split} \tilde{C}_{0}(k)\tilde{\Phi} &:= i(\langle k, \tilde{A}(k) \rangle - |k| \tilde{\varphi}(k))\tilde{\Phi} \neq 0, \end{split}$$
 $\tilde{C}_{0}^{+}(k)\tilde{\Phi} &:= -i(\langle k, \tilde{A}^{+}(k) \rangle - |k| \tilde{\varphi}^{+}(k))\tilde{\Phi} \neq 0 \end{split}$ (186)

invariantly for all $k \in E^3$. Notwithstanding, it is easy enough to check that the following slightly perturbed operators

$$\tilde{C}(k) := \tilde{C}_{0}(k) + \frac{i\xi \exp(-ic |k| \overline{t})}{2 |k| (2\pi)^{3/2}} \int_{\mathbb{R}^{3}} \exp(-i < k, y >) \psi^{+}(y) \psi(y) d^{3}y,$$
(187)

1359

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$$\tilde{C}^{+}(k) := \tilde{C}_{0}^{+}(k) - \frac{i\xi \exp(ic |k| \overline{t})}{2 |k| (2\pi)^{3/2}} \int_{\mathbb{R}^{3}} \exp(i < k, y >) \psi^{+}(y) \psi(y) d^{3}y,$$

1360 are commuting both to each other and with the Hamiltonian operators (177) and (179):

$$[C(k), C(s)] = 0 = [C^+(k), C(s)]$$

1361
$$[\tilde{C}(k), \tilde{H}_{f}^{(int)}] = 0 = [\tilde{C}^{+}(k), \tilde{H}_{f}^{(int)}],$$
(188)

$$[\tilde{C}(k), \tilde{H}_b] = 0 = [\tilde{C}^+(k), \tilde{H}_b]$$

- 1362 for all $k, s \in E^3$. Thus, the related evolution flows (181) and (182) in the transformed Fock space 1363 $\tilde{\Phi}$ should be considered under the modified constraints
- 1364 $\tilde{C}(k)\tilde{\Phi} = 0 = \tilde{C}^+(k)\tilde{\Phi}$ (189)

for all $k \in E^3$. Taking into account the exact expressions (187), the constraints (189) can be equivalently rewritten as

1367

1368

where for all $x \in \mathbb{R}$ and the corresponding temporal parameters \overline{t} and $\tilde{t} \in \mathbb{R}$

$$\tilde{C}(\bar{t}; \bar{t}, x) := \frac{1}{(2\pi)^{3/2}} \int_{\mathbb{R}^3} d^3k \tilde{C}(k) \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}} \exp(i < k, x > -i |k| \tilde{t}) + \frac{1}{(2\pi)^{3/2}}$$

 $\tilde{C}(\bar{t};\tilde{t},x)\tilde{\Phi}=0,$

1369
$$+\frac{1}{(2\pi)^{3/2}} \int_{\mathbb{R}^3} d^3 k \tilde{C}^+(k) \exp(-i < k, x > +i |k| \tilde{t}) =$$
(191)

$$= <\nabla, \tilde{A} > +\frac{1}{c} \frac{\partial \tilde{\varphi}}{\partial \tilde{t}} - \frac{\xi}{2\pi} \int_{\mathbb{R}^3} d^3 y \Theta(c(\overline{t} - \tilde{t}), |x - y|) \psi^+(y) \psi(y)$$

1370 in which we put, by definition, the relativistic generalized function

1371
$$\Theta(c(t-t), |x-y|) := \frac{\delta(|x-y|+c(t-t)) - \delta(|x-y|-c(t-t))}{2|x-y|},$$
(192)

1372 dual to the well known generalized solution [123] [124]

1373
$$\delta(|x-y|^2 - c^2(\overline{t} - \tilde{t})^2) = \frac{\delta(|x-y| + c(\overline{t} - \tilde{t})) + \delta(|x-y| - c(\overline{t} - \tilde{t}))}{2|x-y|}$$

(190)

1374 to the relativistic wave equation.

1375

1376 **Remark 5.** It is here worthy to mention that the above defined operator $\tilde{C}(\bar{t}): \tilde{\Phi} \to \tilde{\Phi}$, 1377 depending parameterically on the bosonic temporal parameter $\bar{t} \in \mathbb{R}$, satisfies the relativistic 1378 wave equation

1379

$$\frac{1}{c^2} \frac{\partial^2 \tilde{C}}{\partial \tilde{t}^2} - \Delta \tilde{C} = 0,$$
(193)

that can be easily checked, if to make use of the rewritten in the Fock space wave equations(167):

1382

$$\frac{1}{c^2}\frac{\partial^2 \tilde{A}}{\partial \tilde{t}^2} - \Delta \tilde{A} = 0, \frac{1}{c^2}\frac{\partial^2 \tilde{\varphi}}{\partial \tilde{t}^2} - \Delta \tilde{\varphi} = 0.$$
(194)

1383 Moreover, as it can be shown by means of direct calculations, the transformed bosonic 1384 Hamiltonian operator (179) on the reduced via the modified Lorenz type constraints (190) Fock 1385 space $\tilde{\Phi}$ persists to be, as before, non-negative definite.

1386

1387 Now we can proceed to deriving the quantum Maxwell equations starting from the 1388 operator equations (194) and the suitably transformed to the Fock space $\tilde{\Phi}$ electromagnetic 1389 fields definitions (158) and (159):

1390
$$(\nabla \times \tilde{B} - \frac{1}{c} \frac{\partial \tilde{E}}{\partial \tilde{t}}) \tilde{\Phi} = \frac{\xi}{2\pi} \nabla \int_{\mathbb{R}^3} y \Theta(c(\bar{t} - \tilde{t}), |x - y|) \psi^+(y) \psi(y) \tilde{\Phi}$$
(195)

1391 and

1396

1392
$$\langle \nabla, \tilde{E} \rangle \tilde{\Phi} = -\frac{\xi}{2\pi} \frac{\partial}{\partial \tilde{t}} \int_{\mathbb{R}^3} d^3 y \Theta(c(\bar{t} - \tilde{t}), |x - y|) \psi^+(y) \psi(y) \tilde{\Phi},$$
 (196)

which are considered in the weak operator sense. Taking now into account the relationships
(182) and (184), one can obtain strong operator relationships for the electrical and magnetic
fields

$$\tilde{E} = -\frac{1}{c} \frac{\partial \tilde{A}}{\partial \tilde{t}} - \nabla \tilde{\varphi} = -\frac{1}{c} \frac{\partial \tilde{A}}{\partial \tilde{t}} - \nabla \tilde{\varphi}, \quad \tilde{B} = \nabla \times \tilde{A}.$$
(197)

1397 with respect to the common reference frame K_r . Similarly one can easily calculate the weak 1398 operator relationship

1399 $(\frac{1}{c}\frac{\partial\tilde{\varphi}}{\partial t} + \langle\nabla,\tilde{A}\rangle)\tilde{\Phi} = 0,$ (198)

1400 which holds for the common temporal parameter $t \in \mathbb{R}$. Now we will calculate the weak 1401 Maxwell type operator relationships (195) and (196) with respect to the common reference 1402 frame K_t :

1403
$$(\nabla \times \tilde{B} - \frac{1}{c} \frac{\partial \tilde{E}}{\partial \tilde{t}})|_{\tilde{t}=\tilde{t}=t} \tilde{\Phi} = \frac{\xi}{2\pi} \nabla \int_{\mathbb{R}^3} d^3 y \Theta(c(\tilde{t}-\tilde{t}),|x-y|) \psi^+(y) \psi(y)|_{\tilde{t}=\tilde{t}=t} \tilde{\Phi} = 0$$
(199)

1404 and

1405
$$<\nabla, \tilde{E} > \tilde{\Phi} = -\frac{\xi}{2\pi} \frac{\partial}{\partial \tilde{t}} \int_{\mathbb{R}^3} d^3 y \Theta(c(\overline{t} - \tilde{t}), |x - y|) \psi^+(y) \psi(y)|_{\tilde{t} = \tilde{t} = t} \tilde{\Phi} = \xi \psi^+ \psi \tilde{\Phi},$$
(200)

1406 where there was used the known [121] [124] generalized function relationship

1407
$$\frac{1}{c}\frac{\partial}{\partial s}\Theta(cs,|z|)|_{s=0} = -2\pi\delta(z)$$
(201)

1408 for all $z \in \mathbb{R}^3$. To calculate further the expression (199), we need to make use of the strong 1409 operator relationships (184) and find that

1410
$$\frac{\partial \tilde{E}}{\partial \tilde{t}}|_{\tilde{t}=\tilde{t}=t} = \frac{\partial \tilde{E}}{\partial t} - \frac{\partial \tilde{E}}{\partial \tilde{t}}|_{\tilde{t}=\tilde{t}=t} = \frac{\partial \tilde{E}}{\partial t} - \frac{i}{\hbar} [\tilde{H}_{f}^{(int)}, \tilde{E}] = \frac{\partial \tilde{E}}{\partial t} + \xi \psi^{+} \alpha \psi.$$
(202)

1411 Thus, from (202) and (199) one can obtain that

1412
$$(\nabla \times \tilde{B} - \frac{1}{c} \frac{\partial E}{\partial t})\tilde{\Phi} = \xi \psi^+ \alpha \psi \tilde{\Phi}$$
(203)

1413 with respect to the common reference frame K_i . The combined together weak operator 1414 relationships (200) and (203)

1415
$$(\nabla \times \tilde{B} - \frac{1}{c} \frac{\partial \tilde{E}}{\partial t}) \tilde{\Phi} = \xi \psi^+ \alpha \psi \tilde{\Phi}, < \nabla, \tilde{E} > \tilde{\Phi} = \xi \psi^+ \psi \tilde{\Phi}$$
(204)

1416 in the reduced by the weak constraint (198) Fock space $\tilde{\Phi}$ jointly with the evident strong 1417 operator relationships

1418
$$\nabla \times \tilde{E} + \frac{1}{c} \frac{\partial \tilde{B}}{\partial \tilde{t}} = 0, \nabla \times \tilde{B} = 0$$
(205)

1419 compile the complete system of quantum Maxwell equations with respect to the common 1420 reference frame K_{r} .

1421 Really, from the Heisenberg evolution equations (181) one easily obtains the strong 1422 operator charge conservative flow relationship

1423
$$\frac{\partial}{\partial t}(\xi\psi^{+}\psi) + \langle \nabla, \xi\psi^{+}\alpha\psi \rangle = 0, \qquad (206)$$

1424 in which the quantity

1425

1427

 $\rho := \xi \psi^+ \psi \tag{207}$

1426 is interpreted as the operator charge density and the quantity

$$J := \xi \psi^+ c \alpha \psi \tag{208}$$

1428 is naturally interpreted as the operator current density in the space R^3 . Whence the weak

1429 operator equations (204) can be rewritten, taking into account the definitions (207) and (208),

1430 in the standard Maxwell equations weak form:

1431
$$(\nabla \times \tilde{B} - \frac{1}{c} \frac{\partial \tilde{E}}{\partial t})\tilde{\Phi} = \frac{J}{c} \tilde{\Phi}, < \nabla, \tilde{E} > \tilde{\Phi} = \rho \tilde{\Phi}$$
(209)

1432 under the Fock space $\tilde{\Phi}$ constraint (198). Moreover, based on the weak operator Maxwell 1433 equations (209) and the Lorenz constraint (198), one can derive easily the following weak 1434 operator linear wave equations

1435 $(\frac{1}{c^2}\frac{\partial^2\tilde{\varphi}}{\partial t^2} - \Delta\tilde{\varphi})\tilde{\Phi} = \rho\tilde{\Phi}, (\frac{1}{c^2}\frac{\partial^2\tilde{A}}{\partial t^2} - \Delta\tilde{A})\tilde{\Phi} = \frac{J}{c}\tilde{\Phi}$ (210)

in respect to the common laboratory reference frame K_r , allowing to calculate the induced by the charged fermionic field causal quantum bosonic potentials $(\tilde{\varphi}_{\xi}, \tilde{A}_{\xi}): \tilde{\Phi} \to \tilde{\Phi}^4$ in the analytical form:

$$\tilde{\varphi}_{\xi} = \frac{1}{4\pi} \int_{\mathbb{R}^3} \frac{\rho(t', y) d^3 y}{|x - y|}, \quad \tilde{A}_{\xi} = \frac{1}{4\pi c} \int_{\mathbb{R}^3} \frac{J(t', y) d^3 y}{|x - y|}, \quad (211)$$

where the "retarded" temporal parameter $t' := t - |x - y|/c \in \mathbb{R}$, making the equations (210) exactly satisfied modulo the solutions to their uniform forms. Moreover, owing to (206), the expressions (211) satisfy exactly the strong operator Lorenz constraint

1443

$$\frac{1}{c}\frac{\partial\tilde{\varphi}_{\xi}}{\partial t} + \langle \nabla, \tilde{A}_{\xi} \rangle = 0$$
(212)

1444 with respect to the laboratory reference frame K_{t} .

1445 From the analysis of the quantum charged particle fermionic field model, interacting 1446 with the self-generated quantum bosonic electromagnetic field, one can infer the following 1447 important consequences: 1448

- the physical effective evolution of the fermionic-bosonic system with respect to the common reference frame K_r is governed by the reduced fermionic Hamiltonian operator (177), acting on the canonically transformed Fock space $\tilde{\Phi}$, reduced by means of the weak Lorenz type operator constraint (198);
- the compatibility of evolutions of the quantum fermionic and bosonic fields with respect to the common temporal reference frame K_i entails the reciprocal influence of the fermionic field on the bosonic one and vice versa, being clearly demonstrated both by the weak field potentials operator equations (210) and the Lorentz type weak constraint (198) imposed on the Fock space $\tilde{\Phi}$;
- 1459

1453

- subject to the basic self-interacting fermionic-bosonic system described by the joint Hamiltonian operator (172) in the transformed Fock space $\tilde{\Phi}$, one can claim that the bosonic electromagnetic impact into the quantum charged particle dynamics is decisive, as owing to it the fermionic system can realize its charge interaction property through the physical vacuum deformation, caused by the related deformation of the weak Lorenz type operator constraint (190), and resulting into the weak operator potential equations (211).
- 1466

1467 The consequences formulated above subject to the quantum fermionic-bosonic self-1468 interacting phenomenon, as it was shown in **Error! Reference source not found.**, appeared to 1469 be very important from classical point of view, especially for physical understanding the inertial 1470 properties of a charged particle under action of the self-generated electromagnetic field.

1471

1472 6. Classical reduction of the quantum charged particle and 1473 electromagnetic field evolutions

1474

1475 Let's consider the vector position operator $\hat{x}: \tilde{\Phi} \to \tilde{\Phi}^3$ and its weak in the reduced Fock 1476 space $\tilde{\Phi}$ evolution with respect to the complete and suitably renormalized charged particle

1477 Hamiltonian operator (177). Taking into account that the Hamiltonian operator $\tilde{H}_{f}^{(int)}: \tilde{\Phi} \to \tilde{\Phi}$ 1478 can be represented as

1479 $\tilde{H}_{f}^{(int)} = \int_{\mathbb{R}^{3}} d^{3}x \psi^{+} < c\alpha, \, \hat{p}_{x} > \psi + \int_{\mathbb{R}^{3}} d^{3}x (\xi \psi^{+} \psi \tilde{\varphi}_{\xi} - \xi \psi^{+} < c\alpha, \, \tilde{A}_{\xi} > \psi), \quad (213)$

1480 within which the operators $(\tilde{\varphi}_{\xi}, \tilde{A}_{\xi}): \tilde{\Phi} \to \tilde{\Phi}^3$ are given by the nonlocal integral expressions 1481 (211) and $\hat{p}_x: \tilde{\Phi} \to \tilde{\Phi}^3$ is the locally defined charged particle ξ momentum operator 1482 $\hat{p}_x:=\frac{\hbar}{i}\nabla_x$, canonically conjugated [71] to the position operator $\hat{x}: \tilde{\Phi} \to \tilde{\Phi}^3$, that is

1483

$$[\hat{p}_{y}, \hat{x}] = \frac{\hbar}{i} \delta(x - y)$$
(214)

1484 for any $x, y \in \mathbb{R}^3$. This also, in particular, means that the position operator $\hat{x}: \tilde{\Phi} \to \tilde{\Phi}^3$ is a 1485 priori given in the diagonal representation: $\hat{x}\tilde{f} := x\tilde{f}$ for any vector $\tilde{f} \in \tilde{\Phi}$.

1486 As a result of a simple calculation one finds the expression

1487
$$d\hat{x} / dt = \psi^+ c \alpha \psi, \qquad (215)$$

1488 which can be used for obtaining the classical charged particle ξ velocity $u(t, x) \in T(\mathbb{R}^3)$ as

1489
$$u(t,x) := (\Omega, d\hat{x} / dt\Omega) = (\Omega, \psi^+ c \alpha \psi \Omega), \qquad (216)$$

1490 where the vector $\Omega \in \tilde{\Phi}$ is the ground state of the Hamiltonian operator (213) acting in the 1491 Lorenz type reduced and suitably renormalized [71] [88] [121] [123] Fock space $\tilde{\Phi}$. 1492 Substituting (215) and (207) into the Hamiltonian expression (213) one obtains the expression

1493
$$\tilde{H}_{f}^{(int)} = \int_{\mathbb{R}^{3}} d^{3}x < d\hat{x} / dt, \, \hat{p}_{x} > + \int_{\mathbb{R}^{3}} d^{3}x (\rho \tilde{\varphi}_{\xi} - \frac{1}{c}J, \, \tilde{A}_{\xi} >), \quad (217)$$

1494 whose classical counterpart looks as

1495
$$\bar{H}_{f}^{(int)} = \int_{\mathbb{R}^{3}} d^{3}x (\rho \tilde{\varphi}_{\xi} - \langle \frac{1}{c}J, \tilde{A}_{\xi} \rangle),$$
(218)

1496 within which there was taken into account the previously assumed quantum massless charged 1497 particle ξ fermionic field. The expression (218) jointly with the renormalized bosonic field 1498 Hamiltonian (162) gives rise to the complete classical Hamiltonian function

1499
$$\overline{H}_{f-b}^{(int)} = \int_{\mathbb{R}^3} d^3 x [\frac{1}{2} (|\tilde{E}|^2 + |\tilde{B}|^2) + \rho \tilde{\varphi}_{\xi} - \langle \frac{1}{c} J, \tilde{A}_{\xi} \rangle],$$
(219)

1500 governing the temporal evolution both of the charged particle ξ and of the electromagnetic 1501 fields with respect to the laboratory reference frame K_i . The obtained above Hamiltonian 1502 function and its corresponding Lagrangian form (166) have been effectively used before in [125] 1503 for describing the classical self-interacting charged particle dynamics and its inertial properties.

Being experienced with the analysis of a self-interacting charged quantum particle 1504 fermionic field with the self-generated quantum bosonic electromagnetic field, we understand 1505 1506 well that the influence of the electromagnetic field on the charged particle should be 1507 considered as crucial, strongly modifying the related fermionic Hamiltonian operator, 1508 describing the charged particle dynamics. As the simultaneously modified bosonic electromagnetic operator depends, owing to the self-interaction, on the charge and current 1509 1510 particle field densities, the joint impact on the charged particle dynamics can be effectively 1511 classically modeled by means of its inertial mass parameter. In the quantum operator case the

1512 physical charged particle mass parameter $m_{ph} \in \mathbb{R}_+$ can be naturally defined by means of the 1513 least quantum renormalized Hamiltonian (172) eigenvalue

- 1514
- 1515

$$m_{ph} := c^{-2} \inf_{\tilde{f} \in \tilde{\Phi}, \|\tilde{f}\|=1} (\tilde{f}, \tilde{H}_{f-b}^{(int)} \tilde{f}), \tilde{H}_{f-b}^{(int)} := \tilde{H}_{f}^{(int)} + \tilde{H}_{b},$$
(220)

in the suitably transformed and reduced by means of the operator Lorenz type constraint (198) Fock space $\tilde{\Phi}$ with respect to the common reference frame K_r . As the quantum spectral problem (220) is very complicated, new tools are needed to be developed for its successful analysis.

1521 7. Classical self-interacting charged particle dynamics and its 1522 inertial properties

1523

1524 Being experienced with the analysis of a self-interacting charged quantum particle 1525 fermionic field with the self-generated quantum bosonic electromagnetic field, we understand well that the influence of the electromagnetic field on the charged particle should be 1526 1527 considered as crucial, strongly modifying the related fermionic Hamiltonian operator, describing the charged particle dynamics. As the simultaneously modified bosonic 1528 electromagnetic operator depends, owing to the self-interaction, on the charge and current 1529 1530 particle field densities, the joint impact on the charged particle dynamics can be effectively 1531 classically modeled by means of its inertial mass parameter. In the quantum operator case the physical charged particle mass parameter $m_{ph} \in \mathbf{R}_+$ can be naturally defined by means of the 1532 least quantum renormalized Hamiltonian (172) eigenvalue (220) in the suitably transformed 1533 and reduced by means of the operator Lorenz type constraint (198) Fock space $ilde{\Phi}$ with 1534 1535 respect to the common reference frame $K_{.}$ As the quantum spectral problem (220) is very 1536 complicated, we will try below to analyze it from the classical point of view.

1537 The guantum operator Hamiltonian approach of Section 4. makes it possible to treat 1538 analytically the charged particle self-interaction mechanism, which can be described by means of the following two steps. The first one consists in producing the charged particle dynamics 1539 governed by the gauge type component of the charged particle Hamiltonian operator (177), 1540 and the second one - consists in modifying this dynamics by means of the self-generated 1541 1542 electromagnetic field, whose influence is governed by the bosonic Hamiltonian (179), perturbed 1543 by the dependence of the electromagnetic field potentials on the related charge and current 1544 densities through the differential relationships (210). This mechanism can be classically realized 1545 analytically by means of the alternative and already before mentioned Lagrangian least action 1546 formalism, following the well known slightly modified [5] Landau-Lifschitz scheme. Namely, the Lagrangian function for the classical charged particle ξ , interacting with the self-generated 1547 1548 electromagnetic field, is easily derived from the corresponding Hamiltonian function (219), 1549 giving rise to the classical Lagrangian expressions (166) in the following slightly extended form:

$$\tilde{\mathcal{L}}_{(f-b)} = \int_{\mathbb{R}^{3}} d^{3}x \left(\left\langle \frac{1}{c}J, \tilde{A} \right\rangle - \rho \tilde{\varphi} \right) + \frac{1}{2} \int_{\mathbb{R}^{3}} d^{3}x \left(\left\langle \nabla \tilde{\varphi} + \frac{1}{c} \frac{\partial \tilde{A}}{\partial t}, \nabla \tilde{\varphi} + \frac{1}{c} \frac{\partial \tilde{A}}{\partial t} \right\rangle -$$
(221)

 $- \langle \nabla \times \tilde{A}, \nabla \times \tilde{A} \rangle - \langle k, dx / dt \rangle,$

where vector $k := k(t, x) \in E^3$ models the related radiation reaction momentum, caused by the 1551 accelerated charged particle ξ with respect to the laboratory reference frame K, as well as 1552 1553 there implied that the classical Lorenz type constraint (198) is satisfied *a priori*. Here we need to 1554 mention that the first part of the Lagrangian (221) is responsible for the internal gauge type 1555 charged particle self-interaction and the second one is responsible for the external charged 1556 particle self-interaction induced by the suitably perturbed electromagnetic field, depending on 1557 the particle charge and current densities. The physical difference between these two phenomena proves to be especially important for calculation of an effective Lagrangian 1558 1559 function for the related dynamical properties of the self-interacting charged particle.

Before proceeding further we need to make an important comment concerning the least action properties of the classical relativistic self-interacting Lagrangian (221). Namely, taking into account a deep quantum vacuum origin [121] of the electromagnetic field and its effective measuring only with respect to the common laboratory reference frame K_i , we can state that the related Maxwell equations should be naturally derived from the following least action principle: the variation $\delta \tilde{S}_{f-b}^{(t)} = 0$, where by definition, the action functional

1566

1550

$$\tilde{S}_{f-b}^{(t)} := \int_{t_1}^{t_2} \tilde{L}_{(f-b)} dt$$
(222)

is calculated with respect to the laboratory reference frame K, on a fixed temporal interval 1567 $[t_1, t_2] \subset \mathbb{R}$. Yet, as it is easy to check, the above action functional (222) fails to derive the 1568 1569 corresponding Lorentz type dynamical equations for the self-interacting charged particle ξ , if to take into account that the related charged particle is considered to be pointwise, located at 1570 point $x(t) \in E^3$ for $t \in R$ and endowed with the current density vector $J = \rho dx(t) / dt \in E^3$ and 1571 the charge density $\rho := \xi \delta(x - x(t)), x \in E^3$. This, evidently, means that the action functional 1572 (222) should be suitably modified with respect to the [1] [51] Feynman proper time reference 1573 1574 frame paradigm, owing to which the action functional for the charged particle dynamics has a 1575 physical sense if and only if it is considered with respect to the proper time reference frame 1576 K_{τ} :

1577
$$\tilde{S}_{f-b}^{(\tau)} := \int_{\tau_1}^{\tau_2} \tilde{L}_{(f-b)} \sqrt{(1+|\dot{x}|^2/c^2)} d\tau$$
(223)

1578 on a fixed temporal interval $[\tau_1, \tau_2] \subset \mathbb{R}$, where we took into account, that 1579 $dt := \sqrt{(1+|\dot{x}|^2/c^2)} d\tau$ and, by definition, the velocity $\dot{x} := dx/d\tau$ with respect to the proper 1580 temporal parameter $\tau \in \mathbb{R}$. Then from the least action condition $\delta \tilde{S}_{f-b}^{(\tau)} = 0$ on the fixed temporal interval $[\tau_1, \tau_2] \subset \mathbb{R}$ one easily obtains the well known classical Lorentz dynamical equation

1583

$$\frac{d}{dt}(mu) = \xi \tilde{E} + \xi u \times \tilde{B},$$
(224)

written with respect to the laboratory reference frame K_i . When deriving (224) there was put, by definition, the inertial mass definition $m := -\tilde{\varphi}/c^2$. The reasonings presented above will be in part employed below when analyzing a suitably reduced Lagrangian function (221).

For the self-interacting charged particle to be physically specified by the mentioned above phenomena in detail, we will consider below a so-called shell model particle, whose charge is uniformly distributed on a sphere of a very small yet fixed radius. Then, following the similar calculations from [5], one can obtain from (221) that

$$\tilde{\mathcal{L}}_{(f-b)} = \frac{1}{2} \int_{\mathbb{R}^3} d^3 x (\tilde{\varphi} < \nabla, \tilde{E} > + \frac{1}{c} < \tilde{A}, \frac{\partial \tilde{E}}{\partial t} > -\frac{1}{c} < \tilde{A}, J + \frac{\partial \tilde{A}}{\partial t} >) -$$

1591
$$-\frac{1}{2c}\frac{d}{dt}\int_{\mathbb{R}^3} d^3x < \tilde{A}, \tilde{E} > +\int_{\mathbb{R}^3} d^3x \left(<\frac{1}{c}J, \tilde{A} > -\rho\tilde{\varphi}\right) -$$

$$-\frac{1}{2}\lim_{r\to\infty}\int_{S_r^2} <\tilde{\varphi}\tilde{E}+\tilde{A}\times\tilde{B}, dS_r^2>-< k, dx/dt>=$$

1592

$$= -\int_{\Omega_{-}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) - d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right) + \int_{\Omega_{+}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi}\right$$

 $-\frac{1}{2c}\frac{d}{dt}\int_{\mathbb{R}^3} d^3x < \tilde{A}, \tilde{E} > - < k, dx / dt > =$

1593

$$= \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi} \right) + \frac{1}{2} \int_{\Omega_{-}(\xi) \cup \Omega_{+}(\xi)} d^{3}x \left(\langle \frac{1}{c}J, \tilde{A} \rangle - \rho \tilde{\varphi} \right) - \frac{1}{2c} \frac{d}{dt} \int_{\mathbb{R}^{3}} d^{3}x \langle \tilde{A}, \tilde{E} \rangle - \langle k, dx / dt \rangle,$$

(225)

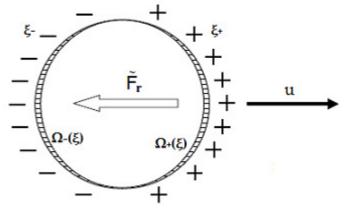


Fig. 1. The courtesy picture from [31]

1594 where we took into account that $\lim_{r\to\infty}\int_{S_r^2} \langle \tilde{\varphi}\tilde{E} + \tilde{A} \times \tilde{B}, dS_r^2 \rangle = 0$, meaning the vanishing of 1595 the radiated by the accelerated charged particle energy, as well as we denoted by $\Omega_-(\xi) :=$ 1596 $\operatorname{supp} \xi_- \subset S^2$ and by $\Omega_+(\xi) := \operatorname{supp} \xi_+ \subset S^2$ the corresponding charge ξ parts supports, located 1597 on the electromagnetic field shadowed rear and on the electromagnetic field exerted on semi-1598 spheres of the charged particle spherical shell $\Omega(\xi) := \Omega_-(\xi) \cup \Omega_+(\xi)$, respectively (see Fig 1.) 1599 to its motion with velocity $u := dx / dt \in E^3$ with respect to the laboratory reference frame K_t .

1600 The expression (225) demonstrates explicitly that during the charged particle motion 1601 the self-generated electromagnetic field interacts effectively only with its frontal part 1602 $\Omega_+(\xi) \subset S^2$ of the particle spherical shell S^2 , as the rear part $\Omega_-(\xi) \subset S^2$ of the particle shell 1603 enters during its motion into the shadowed interior region of the sphere, where the net electric 1604 field $\tilde{E} \in E^3$ is vanishing owing to the charged particle spherical symmetry. To proceed further 1605 we need to calculate the electromagnetic potentials $(\tilde{\varphi}, \tilde{A}): M^4 \to R \times E^3$, using the 1606 determining expressions (211) as $1/c \to 0$:

$$\tilde{\varphi} = \int_{\mathbb{R}^{3}} d^{3}y \frac{\rho(t', y)}{|x - y|} \Big|_{t' = t - |x - y|/c} = \lim_{\varepsilon \downarrow 0} \int_{\mathbb{R}^{3}} d^{3}y \frac{\rho(t - \varepsilon, y)}{|x - y|} + \\ + \lim_{\varepsilon \downarrow 0} \frac{1}{2c^{2}} \int_{\mathbb{R}^{3}} d^{3}y |x - y| \partial^{2}\rho(t - \varepsilon, y) / \partial t^{2} + \\ + \lim_{\varepsilon \downarrow 0} \frac{1}{6c^{3}} \int_{\mathbb{R}^{3}} d^{3}y |x - y|^{2} \partial\rho(t - \varepsilon, y) / \partial t + O(1/c^{4}) =$$

$$= \int_{\Omega_{+}} (\xi) d^{3}y \frac{\rho(t, y)}{|x - y|} + \frac{1}{2c^{2}} \int_{\Omega_{+}} (\xi) d^{3}y |x - y| \partial^{2}\rho(t, y) / \partial t^{2} + \\ + \frac{1}{6c^{3}} \int_{\Omega_{+}} (\xi) d^{3}y |x - y|^{2} \partial\rho(t, y) / \partial t + O(1/c^{4}),$$
(226)

1608

$$\tilde{A} = \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t', y)}{|x - y|} \Big|_{t' = t - |x - y|/c} = \lim_{\varepsilon \downarrow 0} \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3} d^3 y \frac{J(t - \varepsilon, y)}{|x - y|} - \frac{1}{c} \int_{\mathbb{R}^3$$

$$-\lim_{\varepsilon \downarrow 0} \frac{1}{c^2} \int_{\mathbb{R}^3} d^3 y \partial J(t-\varepsilon, y) / \partial t +$$

1609

$$+\lim_{\varepsilon \downarrow 0} \frac{1}{2c^3} \int_{\mathbb{R}^3} d^3 y \, | \, x - y \, | \, \partial^2 J(t - \varepsilon, y) \, / \, \partial t^2 + O(1/c^4) =$$

$$= \frac{1}{c} \int_{\Omega_{+}(\xi)} d^{3}y \frac{J(t, y)}{|x - y|} - \frac{1}{c^{2}} \int_{\Omega_{+}(\xi)} d^{3}y \partial J(t, y) / \partial t +$$

+
$$\frac{1}{2c^3}\int_{\Omega_+}(\xi)d^3y |x-y|\partial^2 J(t,y)/\partial t^2 + O(1/c^4),$$

1610 where the limit $\lim_{\varepsilon \downarrow 0} (...)$ was treated physically, that is taking into account the assumed above 1611 spherical shell model of the charged particle ξ and its corresponding self-interaction during its 1612 motion. Now, as a result of simple enough calculations based on the electromagnetic potentials 1613 (226), the effective expression for the classical Lagrangian (225) can be equivalently rewritten 1614 up to $O(1/c^4)$ accuracy with respect to the laboratory reference frame K_r as

1615
$$\tilde{L}_{(f-b)}^{(t)} = \frac{E_{es}}{2c^2} |u|^2, \qquad (227)$$

1616 if to make use of the following integral expressions:

$$\int_{\Omega_{+}(\xi)\cup\Omega_{-}(\xi)} d^{3}x \int_{\Omega_{+}(\xi)\cup\Omega_{-}(\xi)} d^{3}y \rho(t,y)\rho(t,y) := \xi^{2},$$

$$\frac{1}{2} \int_{\Omega_{+}(\xi)\cup\Omega_{-}(\xi)} d^{3}x \int_{\Omega_{+}(\xi)\cup\Omega_{-}(\xi)} d^{3}y \frac{\rho(t,y)\rho(t,y)}{|x-y|} := \mathbf{E}_{es},$$

$$\int_{\Omega_{+}(\xi)} d^{3}x \rho(t,x) \int_{\Omega_{+}(\xi)} d^{3}y \frac{\rho(t;y)}{|y-x|} = \frac{1}{2} \mathbf{E}_{es},$$

$$\int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \int_{\Omega_{-}(\xi)} d^{3}y \frac{\rho(t;y)}{|y-x|} = \frac{1}{2} \mathbf{E}_{es},$$
(228)

$$\int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \int_{\Omega_{+}(\xi)} d^{3}y \frac{\rho(t;y)}{|x-y|} |\frac{\langle y-x,u\rangle}{|y-x|}|^{2} >:= \frac{E_{es}}{6} |u|^{2}$$

$$\int_{\Omega_{+}(\xi)} d^{3}x \rho(t,x) \int_{\Omega_{+}(\xi)} d^{3}y \frac{\rho(t;y)}{|x-y|} |\frac{\langle y-x,u\rangle}{|y-x|}|^{2} >:= \frac{E_{es}}{6} |u|^{2},$$

1618 obtained owing to the reasonings similar to those in [2] [126]. Now, to derive from the reduced Lagrangian function (227) the corresponding dynamic equation for the charged shell model 1619 particle ξ , we need within the discussed above the Feynman proper time paradigm to 1620 transform this Lagrangian with respect to the charged particle proper time reference frame 1621 1622 K₇ :

$$\tilde{\mathcal{L}}_{(f-b)}^{(t)} \to \tilde{\mathcal{L}}_{(f-b)}^{(\tau)} = \frac{\overline{m}_{es}}{2} |\dot{x}|^2 - \langle k, \dot{x} \rangle,$$
(229)

where we denoted by 1624

1623

1617

$$\overline{m}_{es} := m_{es} \sqrt{1 - |u|^2 / c^2}$$
(230)

1626 the so-called relativistic rest mass of the charged particle with respect to the proper time 1627 reference frame K_{τ} , and by

1628

$$m_{es} := \mathcal{E}_{es} / c^2 \tag{231}$$

the so-called charged particle electromagnetic mass with respect to the laboratory reference 1629 frame K. Based on the Lagrangian function (229) one can construct up to $O(1/c^2)$ the 1630 generalized charged particle inertial momentum 1631 1632

$$\tilde{\pi}_f := m_{ph} u - k \tag{232}$$

1633

1634

as

$$\tilde{\pi}_{f} = \partial \tilde{\mathbf{L}}_{(f-b)}^{(\tau)} / \partial \dot{x} = m_{es} u - k , \qquad (233)$$

satisfying with respect to the proper time reference frame $\,K_{\!\tau}\,$ the evolution equation 1635

1636
$$d\tilde{\pi}_{f} / d\tau = \partial \tilde{L}_{(f-b)}^{(\tau)} / \partial x = 0,$$
(234)

1637 which is equivalent to Lorentz type equation

 $d(m_{as}u) / dt = dk(t) / dt := \tilde{F}_r$

with respect to the laboratory reference frame K_{μ} , where the right hand side of (235) means, 1639

by definition, the corresponding radiation reaction force \tilde{F}_r . Having applied to the Lagrangian 1640

function (229) the standard Legendre transformation, one easily finds the quasi-classical 1641 1642 conserved Hamiltonian function

 $\mathbf{H}_{f-b}^{(t)} := <\tilde{\pi}_{f}, \dot{x} > -\tilde{\mathbf{L}}_{(f-b)}^{(\tau)} = \frac{m_{es} |u|^{2}}{2} (1 + \frac{1}{2} |u|^{2} / c^{2}),$ (236) 1643

satisfying with respect to the laboratory reference frame K_t the condition $dH_{f-b}^{(t)}/dt = 0$ for all 1644 $t \in \mathbb{R}$. Yet, the most interesting and important consequence from (236) and the dynamic 1645 equation (235) consists in coinciding the electromagnetic mass parameter $m_{es} \in \mathbf{R}_+$: 1646

1647
$$m_{nbys} := m_{as}, \tag{237}$$

1648 defined by (231), with the naturally related and physically observed inertial mass $m_{nhvs} \in \mathbb{R}_+$, as it was conceived by H. Lorentz and M. Abraham more than one hundred years ago. 1649

1650 1651

8. The radiation reaction force analysis

1652

To calculate the radiation reaction force (235) one can make use of the classical Lorentz 1653 type force expression (224) and obtain in the case of the charged particle shell model, similarly 1654 to [2] [126], up to $O(1/c^4)$ accuracy, the resulting self-interacting Abraham-Lorentz type force 1655 expression with respect to the laboratory reference frame K. Owing to the zero net force 1656 condition, we have that 1657

$$d\tilde{\pi}_{f} / dt + \tilde{F}_{s} = 0, \qquad (238)$$

where, by definition, $\tilde{\pi}_{t} := m_{ph}u$, the Lorentz force can be rewritten in the following form: 1659

$$\tilde{F}_{s} = -\frac{1}{2c} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \frac{d}{dt} \tilde{A}(t,x) -$$

$$\frac{1}{2c} \int_{\Omega_{+}(\xi) \cup \Omega_{-}(\xi)} d^{3}x \rho(t,x) \frac{d}{dt} \tilde{A}(t,x) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) - \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}x \rho(t,x) \nabla \tilde{\varphi}(t,x) (1 - |u/c|^{2}) + \frac{1}{2} \int_{\Omega_{-}(\xi)} d^{3}$$

1660

$$-\frac{1}{2}\int_{\Omega_{+}}(\xi)\cup\Omega_{-}(\xi)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)\nabla\tilde{\varphi}(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)d^{3}x\rho(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)d^{3}x\rho(t,x)\left(1-|u/c|^{2}\right)d^{3}x\rho(t,x)d^{3}$$

Based on calculations similar to those of [2] [126], from (239) and (226) one can obtain, within 1661 the charged particle shell model, for small |u/c| = 1 and slow enough acceleration that 1662

(239)

(235)

$$\begin{split} \tilde{F}_{s} &= \sum_{n \in \mathbb{Z}_{+}} \frac{(-1)^{n+1}}{2n!c^{n}} (1 - |u/c|^{2}) [\int_{\Omega_{-}(\xi)} \rho(t, x) d^{3}x(\cdot) + \\ &+ \int_{\Omega_{+}(\xi) \cup \Omega_{-}(\xi)} \rho(t, x) d^{3}x(\cdot)] \int_{\Omega_{+}(\xi)} d^{3}y \frac{\partial^{n}}{\partial t^{n}} \rho(t, y) \nabla |x - y|^{n-1} + \\ &+ \sum_{n \in \mathbb{Z}_{+}} \frac{(-1)^{n+1}}{2n!c^{n+2}} [\int_{\Omega_{-}(\xi)} \rho(t, x) d^{3}x(\cdot) + \\ &+ \int_{\Omega_{+}(\xi) \cup \Omega_{-}(\xi)} \rho(t, x) d^{3}x(\cdot)] \int_{\Omega_{+}(\xi)} d^{3}y |x - y|^{n-1} \frac{\partial^{n+1}}{\partial t^{n+1}} J(t, y] = \\ &= \sum_{n \in \mathbb{Z}_{+}} \frac{(-1)^{n+1}}{2n!c^{n+2}} (1 - |u/c|^{2}) [\int_{\Omega_{-}(\xi)} \rho(t, x) d^{3}x(\cdot) + \\ &+ \int_{\Omega_{+}(\xi) \cup \Omega_{-}(\xi)} \rho(t, x) d^{3}x(\cdot)] \int_{\Omega_{+}(\xi)} d^{3}y \frac{\partial^{n-2}}{\partial t^{n+2}} \rho(t, y) \nabla |x - y|^{n+1} + \end{split}$$

1663

$$+ \int_{\Omega_+(\xi) \cup \Omega_-(\xi)} \rho(t, x) d^3 x(\cdot) \left[\int_{\Omega_+(\xi)} d^3 y \frac{\partial^{n-2}}{\partial t^{n+2}} \rho(t, y) \nabla |x-y|^{n+1} + \frac{\partial^{n-$$

$$+\sum_{n\in\mathbb{Z}_{+}}\frac{(-1)^{n+1}}{2n!c^{n+2}}[\int_{\Omega_{-}}(\xi)\rho(t,x)d^{3}x(\cdot)+$$

$$+ \int_{\Omega_{+}(\xi) \cup \Omega_{-}(\xi)} \rho(t, x) d^{3}x(\cdot)] \int_{\Omega_{+}(\xi)} d^{3}y |x - y|^{n-1} \frac{\partial^{n+1}}{\partial t^{n+1}} J(t, y).$$

The relationship above can be rewritten, owing to the charge continuity equation (206)-(208) 1664 1665 and the rotational symmetry property, giving rise to the radiation force differential-integral expression: 1666

$$\tilde{F}_{s} = \frac{d}{dt} \left[\sum_{n \in \mathbb{Z}_{+}} \frac{(-1)^{n+1}}{6n! c^{n+2}} \left[\int_{\Omega_{-}(\xi)} \rho(t, x) d^{3}x(\cdot) + \int_{\Omega_{+}(\xi) \cup \Omega_{-}(\xi)} \rho(t, x) d^{3}x(\cdot) \right] \times \right]$$

1667

$$\times \int_{\Omega_{+}(\xi)} d^{3}y |x-y|^{n-1} \frac{\partial^{n}}{\partial t^{n}} J(t,y) - \sum_{n \in \mathbb{Z}_{+}} \frac{(-1)^{n} |u|^{2}}{6n! c^{n+4}} \left[\int_{\Omega_{-}(\xi)} \rho(t,x) d^{3}x(\cdot) + \right]$$
(241)

$$+ \int_{\Omega_+} (\xi) \cup \Omega_-(\xi)^{\rho(t,x)} d^3 x(\cdot) \int_{\Omega_+} (\xi)^{\sigma(t,x)} d^3 y |x-y|^{n-1} \frac{\partial^n}{\partial t^n} J(t,y)]$$

From the latter, taking into account the integral expressions (228), one finds from (241) 1668 up to the $O(1/c^4)$ accuracy the final radiation reaction force expression 1669

$$\tilde{F}_{s} = -\frac{d}{dt} \left(\frac{\mathbf{E}_{es}}{c^{2}} u \right) + \frac{2\xi^{2}}{3c^{3}} \frac{d^{2}u}{dt^{2}} =$$
(242)

1670

$$= -\frac{d}{dt}(m_{es}u) + \frac{2\xi^2}{3c^3}\frac{d^2u}{dt^2} = -\frac{d}{dt}(m_{es}u - \frac{2\xi^2}{3c^3}\frac{du}{dt})$$

holds. We mention here that following the reasonings from [7] [31] [35] [105] [106], in the expressions above there is taken into account an additional hidden and the velocity $u \in T(\mathbb{R}^3)$ directed electrostatic Coulomb surface self-force, acting only on the *front half part* of the spherical electron shell. As a result, from (238), (239) and the relationship (232) one obtains that the generalized charged particle momentum

1676
$$\tilde{\pi}_{p} := m_{es}u - \frac{2\xi^{2}}{3c^{3}}\frac{du}{dt} = m_{es}u - k,$$
(243)

1677 thereby defining both the radiation reaction momentum $k(t) = \frac{2\xi^2}{3c^3} \frac{du(t)}{dt}$ for all $t \in \mathbb{R}$ and the

- 1678 corresponding radiation reaction force
- 1679 $\tilde{F}_r = \frac{2\xi^2}{3c^3} \frac{d^2u}{dt^2},$ (244)

1680 coincides exactly with the classical Abraham-Lorentz--Dirac expression. From (243) one easily1681 follows that the observable physical charged particle shell model inertial mass

1682 $m_{ph} = m_{es} = E_{es} / c^2$ (245)

is of the electromagnetic origin, coinciding exactly with the result (237) obtained above.Moreover, (243) ensues the final force expression

1685 $\frac{d}{dt}(m_{es}u) = \frac{2\xi^2}{3c^3}\frac{d^2u}{dt^2} + O(1/c^4).$ (246)

The latter means, in particular, that the real physically observed " inertial" mass m_{ph} of 1686 the charged shell model particle ξ is strongly determined by its electromagnetic self-1687 interaction energy E_{es} with respect to the laboratory reference frame K_t . A similar statement 1688 1689 there was recently discussed in [31] [35], based on the vacuum Casimir effect type 1690 considerations. Moreover, the assumed above boundedness of the electrostatic self-energy E. 1691 appears to be completely equivalent both to the presence of the so-called intrinsic Poincaré type " tensions", analyzed in [7] [31] [118], and to the existence of a special compensating 1692 1693 Coulomb " pressure", suggested in [35], guaranteeing the assumed electron stability in the 1694 works of H. Lorentz and M. Abraham.

- 1695 1696
- 8.1. Comments
- 1697

1698 The electromagnetic mass origin problem was reanalyzed in details within the Feynman 1699 proper time paradigm and related vacuum field theory approach by means of the fundamental 1700 least action principle and the Lagrangian and Hamiltonian formalisms. The resulting electron 1701 inertia appeared to coincide in part, in the quasi-relativistic limit, with the momentum

expression obtained more than one hundred years ago by M. Abraham and H. Lorentz [53] 1702 1703 [54] [55] [64], yet it proved to contain an additional hidden impact owing to the imposed 1704 electron stability constraint, which was taken into account in the original action functional as 1705 some preliminarily undetermined constant component. As it was demonstrated in [31] [35], 1706 this stability constraint can be successfully realized within the charged shell model of electron 1707 at rest, if to take into account the existing ambient electromagnetic " dark" energy fluctuations, 1708 whose inward directed spatial pressure on the electron shell is compensated by the related 1709 outward directed electrostatic Coulomb spatial pressure as the electron shell radius satisfies 1710 some limiting compatibility condition. The latter also allows to compensate simultaneously the 1711 corresponding electromagnetic energy fluctuations deficit inside the electron shell, thereby 1712 forbidding the external energy to flow into the electron. In contrary to the lack of energy flow inside the electron shell, during the electron movement the corresponding internal momentum 1713 1714 flow is not vanishing owing to the nonvanishing hidden electron momentum flow caused by the 1715 surface pressure flow and compensated by the suitably generated surface electric current flow. 1716 As it was shown, this backward directed hidden momentum flow makes it possible to justify the 1717 corresponding self-interaction electron mass expression and to state, within the electron shell 1718 model, the fully electromagnetic electron mass origin, as it has been conceived by H. Lorentz 1719 and M. Abraham and strongly supported by R. Feynman in his Lectures [1]. This consequence is 1720 also independently supported by means of the least action approach, based on the Feynman 1721 proper time paradigm and the suitably calculated regularized retarded electric potential impact into the charged particle Lagrangian function. 1722

The charged particle radiation problem, revisited in this Section, allowed to conceive the explanation of the charged particle mass as that of a compact and stable object which should be exerted by a vacuum field self-interaction energy. The latter can be satisfied if to impose on the intrinsic charged particle structure [30] some nontrivial geometrical constraints. Moreover, as follows from the physically observed particle mass expressions (245), the electrostatic potential energy being of the self-interaction origin, contributes into the inertial mass as its main relativistic mass component.

1730 There exist different relativistic generalizations of the force expression (246), which 1731 suffer the common physical inconsistency related to the no radiation effect of a charged 1732 particle in uniform motion.

Another deeply related problem to the radiation reaction force analyzed above is the search for an explanation to the Wheeler and Feynman reaction radiation mechanism, called the absorption radiation theory, strongly based on the Mach type interaction of a charged particle with the ambient vacuum electromagnetic medium. Concerning this problem, one can also observe some of its relationships with the one devised here within the vacuum field theory approach, but this question needs a more detailed and extended analysis.

1739

9. Supplement: the "minimum" interaction principle and itsgeometric backgrounds

1742

1743 In this Section we will sketch analytical backgrounds of the "minimum" interaction 1744 principle widely used in modern theoretical and mathematical physics. For description of a moving point charged particle under external electromagnetic field, we will make use of the geometric approach [64]. Namely, let a trivial fiber bundle structure $\pi : M \to R^3, M = R^3 \times G$, with the abelian structure group $G := R \setminus \{0\}$, equivariantly act on the canonically symplectic coadjoint space $T^*(M)$. The latter possesses the canonical symplectic structure

1749
$$\omega^{(2)}(p, z; x, g) := d(pr_*)^* \alpha^{(1)}(x, g) = \langle dp, \wedge dx \rangle + + \langle dz, \wedge g^{-1} dg \rangle_G + \langle z dg^{-1}, \wedge dg \rangle_G$$
(247)

for all $(p, z; x, g) \in T^*(M)$, where $\alpha^{(1)}(x, g) := \langle p, dx \rangle + \langle z, g^{-1}dg \rangle_G \in T^*(M)$ is the corresponding Liouville form on $T^*(M)$ and $\langle \cdot, \cdot \rangle$ is the usual scalar product in E^3 . On the fibered space M one can define a connection Γ by means of an one-form $A: M \to T^*(M) \times G$, determined as

1754
$$A(x,g) := g^{-1} < \xi A(x), dx > g + g^{-1} dg$$
(248)

1755 with $\xi \in G^*, (x,g) \in \mathbb{R}^3 \times G$. The corresponding curvature 2-form $\Sigma^{(2)} \in \Lambda^2(\mathbb{R}^3) \otimes \mathbb{G}$ is

1756
$$\Sigma^{(2)}(x) := dA(x,g) + A(x,g) \wedge A(x,g) = \xi \sum_{i,j=1}^{3} F_{ij}(x) dx^{i} \wedge dx^{j},$$
(249)

1757 where

1758

$$F_{ij}(x) \coloneqq \frac{\partial A_j}{\partial x^i} - \frac{\partial A_i}{\partial x^j}$$
(250)

for $i, j = \overline{1,3}$ is the spatial electromagnetic tensor with respect to the reference frame K_i. For 1759 an element $\xi \in G^*$ to be compatibly fixed, we need to construct the related momentum 1760 mapping $l:T^*(M) \to G^*$ with respect to the canonical symplectic structure (247) on $T^*(M)$, 1761 and put, by definition, $l(x, p) := \xi \in G^*$ to be constant, $P_{\xi} := l^{-1}(\xi) \subset T^*(M)$ and 1762 $G_{\xi} = \{g \in G : Ad_{G}^{*}\xi\}$ to be the corresponding isotropy group of the element $\xi \in G^{*}$. Next we 1763 can apply the standard [47] [64] [96] invariant Marsden-Weinstein-Meyer reduction scheme to 1764 the orbit factor space $\tilde{P}_{\xi} := P_{\xi} / G_{\xi}$ subject to the corresponding group G action. Then, as a 1765 1766 result of the Marsden-Weinstein-Meyer reduction, one finds that G_{ξ} ; G, the factor-space $ilde{P}_{\mathcal{E}}$; $T^*(\mathbb{R}^3)$ becomes Poisson space with the suitably reduced symplectic structure 1767 $\bar{\omega}_{\xi}^{(2)} \in T^*(\tilde{P}_{\xi})$. The corresponding Poisson brackets on the reduced manifold \tilde{P}_{ξ} equal to 1768

1769
$$\{x^{i}, x^{j}\}_{\xi} = 0, \{p_{j}, x^{i}\}_{\xi} = \delta^{i}_{j},$$
$$\{p_{i}, p_{j}\}_{\xi} = \xi F_{ij}(x)$$
(251)

1770 for $i, j = \overline{1,3}$, being considered with respect to the laboratory reference frame K_i . Based on 1771 (251) is worth to observe that a new so called "*shifted*" momentum variable 1772 $\tilde{\pi} := p + \xi A(x)$ (252)

1773 on \tilde{P}_{ξ} gives rise to the symplectomorphic transformation $\overline{\omega}_{\xi}^{(2)} \rightarrow \tilde{\omega}_{\xi}^{(2)} := \langle d\tilde{\pi}, \wedge dx \rangle \in$ 1774 $\Lambda^2(T^*(\mathbb{R}^3))$. The latter gives rise to the following important in theoretical physics "minimal 1775 interaction" canonical Poisson brackets:

$$\{x^{i}, x^{j}\}_{\bar{\omega}_{\xi}^{(2)}} = 0, \{\tilde{\pi}_{j}, x^{i}\}_{\bar{\omega}_{\xi}^{(2)}} = \delta^{i}_{j}, \{\tilde{\pi}_{i}, \tilde{\pi}_{j}\}_{\bar{\omega}_{\xi}^{(2)}} = 0$$
(253)

1777 for $i, j = \overline{1,3}$, represented already with respect to some new reference frame K_{t} , 1778 characterized by the phase space coordinates $(x, \tilde{\pi}) \in \tilde{P}_{\xi}$ and a new evolution parameter 1779 $t' \in \mathbb{R}$, as the spatial Maxwell field compatibility equations

1780

$$\partial F_{ij} / \partial x_k + \partial F_{jk} / \partial x_i + \partial F_{ki} / \partial x_j = 0$$
(254)

are identically satisfied on \mathbb{R}^3 for all $i, j, k = \overline{1, 3}$, owing to the electromagnetic curvature tensor (250) definition.

- 1783 1784
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